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The undersigned, appointed by the dean of the Graduate School, have examined the dissertation entitled:

## Radioanalytical Multi-elemental Analysis: New Methodology and Archaeometric Applications

Presented by Magen E. Coleman, a candidate for the degree of Doctor of Philosophy, and hereby certify that, in their opinion, it is worthy of acceptance.

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#### Abstract

For my parents who always supported me no matter what I wanted to do;


## For my professors

 who believed that I could accomplish great things;For my friends, fellow graduate students, who helped me survive with my sanity (relatively) intact;

## And for Beau,

who was always there to help, support, listen, and just love me.

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#### Abstract

Several projects are covered in this dissertation: the application of instrumental neutron activation analysis (INAA) and rigorous statistical analyses to the sourcing of Egyptian limestone and Kenyan obsidian, and the development of a method to determine titanium, barium, and arsenic concentrations in obsidian using epithermal neutron activation.

INAA, when coupled with rigorous statistical methods, including principal component analysis and clustering techniques, can provide the precision and confidence needed to accurately determine the source of material. However, this technique has not been fully explored for the provenancing of Egyptian limestone sculpture. A combination of the elemental concentration data and rigorous statistical methods is used to study the compositional differences between known ancient quarries. Preliminary results show that INAA has the potential to be effective for limestone provenance studies.

Obsidian is an important component of East African artifacts. Although compositional studies of obsidian from East Africa conducted in the 1980s showed great potential, a comprehensive database has not been developed. Here, samples from Kenya were examined via X-ray fluorescence (XRF) and INAA. The results indicate that there is a clear correlation between geographic proximity and chemical composition. Using more rigorous statistical methods, including principal component analysis, subgroups can be classified and better correlation to geographical groups is observed.

Titanium and barium are often used for characterizing obsidian, especially in areas in Africa. A method has been developed to analyze for these elements using epithermal neutron activation analysis, which takes advantage of larger natural abundances of parent nuclides and a higher probability for epithermal neutron reactions. It produces results with improved accuracy and precision for titanium and barium analyses in obsidian. It is also now possible to analyze for arsenic, which has not previously been reported in obsidian studies.

Archaeology provides a link to the history of man and, through the application of rigorous analytical techniques, insight into the procurement, distribution, and exchange of resources can be achieved. The use of chemical and radiochemical signatures to distinguish artifacts in an accurate and precise manner brings the histories locked in these pieces closer to becoming clear.


## Chapter 1:

## Multi-Elemental Analysis and Archaeometry

### 1.1 Introduction

Elemental studies of artifacts have been an important part of chemistry and archaeology since the 1800s. ${ }^{1-3}$ Many different analytical techniques have been used over the years, from simple gravimetric techniques to determine the concentration of a single element to more complex spectroscopic methods used for multi-elemental analysis. ${ }^{1-2}$ The application of these techniques to artifacts also served various purposes. In many cases, it began with an archaeological question about the source, the age, the manufacturing technology, or the authenticity. In a few cases, though, artifacts were used as a convenient sample set for new or improved techniques. Whatever the original motive, this type of work has demonstrated the ability of analytical chemistry to aid in answering archaeological questions. ${ }^{3}$

Archaeological artifacts provide a unique challenge for the analytical chemist. ${ }^{3}$ These are precious samples that cannot be duplicated. Often, archaeologists and museums are hesitant to give up samples for destructive analysis, so non-destructive methods or methods that are minimally destructive have to be employed. The archaeologist also wants to get as much information as possible out of each sample to aid
in the studies of ancient peoples. In addition, there is the issue of weathering and other chemical changes that may occur over time that may interfere with the measurement or the final analysis of the results. Artifacts usually have complex matrices that can present a myriad of challenges on their own. These are all things that the analyst must keep in mind when choosing an analytical technique for elemental analysis of archaeological samples.

Different definitions of the idea of destructive analysis have been used throughout the years as archaeologists and chemists work to find a good balance between getting enough information out of an artifact with the smallest amount of destruction possible. ${ }^{3}$ Some believe that the artifact must remain completely unchanged in order for the technique to be considered nondestructive, while others believe that a small amount removed from the sample that does not detract from an understanding of its purpose or form is perfectly acceptable. For some techniques, especially extremely sensitive ones such as inductively-coupled plasma mass spectrometry (ICP-MS) and neutron activation analysis (NAA), only milligram amounts of samples are needed for the analysis. In these cases, a small hole can be drilled into the artifact or a small chip broken off, so that the damage is almost imperceptible. There are also techniques, such as x-ray fluorescence and electron microprobe analysis, which can provide elemental analysis without any damage to the artifact at all. These techniques tend to be less sensitive than NAA and ICP-MS, but can still provide valuable information about the sample. ${ }^{3-4}$

One major area of research in archaeology that has benefited from multielemental analysis is the sourcing of raw materials. ${ }^{3,5}$ The most common materials that have been applied to these studies are ceramics, lithics, and building materials. In order
to accurately determine the source for a material, the chemical composition needs to act as a fingerprint for the source. Neutron activation analysis (NAA) and x-ray fluorescence (XRF) are the most common techniques applied to sourcing studies because they both offer multi-elemental analysis with good sensitivity for many of the elements of interest. The trace elements in the composition of the samples are usually used to discriminate among sources. These elements are mostly the rare earth elements, though other elements can be used. Both XRF and NAA offer sensitivity down to the part-per-million level for many of the rare earth elements.

This chapter will describe both the techniques of XRF and NAA and how they have been applied to the analysis of obsidian and limestone in this dissertation. The advantages and disadvantages of each will be presented. Also, the choice of NAA for the examination of the sources will be explained in more detail.

### 1.2 X-ray fluorescence

X-ray fluorescence (XRF) analysis is a nondestructive technique that uses secondary x-rays to characterize the chemical composition of a sample. In XRF analysis, a primary beam of x-rays strikes the sample, interacting with the atoms and ejecting inner-shell electrons. This puts the atom in an excited state, and electrons from the outer shells are quickly transferred to fill the vacancies, giving off an x-ray in the process. These secondary x-rays are characteristic of the element and their energy is equal to the difference in binding energy between the shell initially occupied by the electron and the shell it occupies after filling the vacancy. ${ }^{6}$ A diagram of this process is shown in Figure 1.1.

Figure 1.1: Illustration of $X$-ray Fluorescence


XRF analysis can be used for both qualitative and quantitative analysis. The energy (and wavelength) of the secondary x-rays are characteristic for a particular element and the number of x-rays of a certain energy is proportional to the amount of that element in a sample. ${ }^{7}$

XRF analysis offers several advantages for archaeological applications. First and foremost, it is completely nondestructive. With the trend of XRF instruments toward the completely portable, the size of an artifact is becoming less of an issue, since it does not need to fit in a small sample chamber. Portable instruments also allow the analysis to be performed on site instead of requiring the sample to be brought to a laboratory. For labbased instruments, as long as the sample fits inside the sample chamber, the analysis is nondestructive. In addition, very little sample preparation is needed. The sample needs to have a clean surface and, in the case of bulk analysis, the spot for analysis should be representative of the whole object. XRF also offers simultaneous multi-elemental analysis that is very rapid, usually on the order of several minutes per sample. Along with the improvements in technology allowing for portable instruments, with similar sensitivities as the larger lab-based instruments for many elements, the instruments are also relatively inexpensive and accessible.

XRF analysis also has several disadvantages. Because of the limited penetrating power of both the primary x-ray beam and the secondary x-rays emitted from the sample, the analysis is often limited to the first few millimeters of the sample, and the matrix plays a very important role in the analysis. XRF analysis is also not applicable to all elements. Most instruments only work well for elements with atomic numbers greater than 13 , though better sensitivity for low-Z elements can be achieved by analyzing under
vacuum. XRF analysis tends not to be as sensitive as semi-destructive analytical techniques such as ICP-MS and NAA, and the limits of detection tend to be higher.

### 1.2.1 XRF instruments

XRF instruments consist of four basic components: an excitation source, a means of separating and isolating characteristic x-ray lines, a detection system, and a data collection and processing system. The excitation source is usually a sealed x-ray tube with a rhodium, tungsten, or silver anode, but radioisotopes such as ${ }^{241} \mathrm{Am},{ }^{109} \mathrm{Cd}$, and ${ }^{155} \mathrm{Eu}$ can be used as $\gamma$-sources, especially for the excitation of high-Z elements. The detection system can vary depending on the type of XRF analysis being done. Most commonly, either solid-state scintillation detectors or semiconductor detectors are used. ${ }^{7,8}$

XRF analysis can be divided into two general categories, based on the means of separating the characteristic x-rays. Wavelength-dispersive XRF (WDXRF) was the first type of XRF to be regularly used in the lab. Here, the signal from the sample is separated by Bragg diffraction from a single crystal. ${ }^{6}$ Then each wavelength is detected by two detectors in sequence: a proportional counter for low energy x-rays and a scintillation detector for high energy x-rays. Multielemental analysis can either be done sequentially, using a single appropriate crystal set to the appropriate angle, or simultaneously. If the analysis is to be done simultaneously, several spectrometers are set out around the sample, each with a crystal tuned to the appropriate angle for the analyte. ${ }^{6}$ The major advantage of WDXRF is the resolution that can be achieved; modern instruments can have resolutions of about 5 eV .

Most XRF instruments used in laboratories today are energy-dispersive XRF (EDXRF) instruments. EDXRF differs from WDXRF in that the detector, usually a semiconductor or solid-state scintillation detector, operates as both a means to separate out the characteristic x-rays and a detector for the intensity of the peaks. EDXRF instruments are relatively simple compared to WDXRF instruments, and though they do not have as good a resolution (usually about $100-300 \mathrm{eV}$ ), they have the ability to quickly and simultaneously analyze many elements accurately. ${ }^{6,8}$ Because of the simplicity of the instrument and advances in technology, portable EDXRF units are becoming more widely available.

### 1.2.2 Basic principles of XRF analysis

When using XRF analysis to determine the chemical composition of a sample, several factors must be considered. First is the interaction of x-rays with the sample matrix. As an x-ray beam passes through matter, the intensity of that beam is attenuated as it interacts with the sample. Two basic interaction processes tend to dominate: either the x-ray is scattered or the x-ray transfers its energy to an electron and ejects it from the atom in the sample (photoelectric effect). ${ }^{6}$ Though important information can be achieved through x-ray scattering (i.e., x-ray diffraction), this discussion will focus on the photoelectric effect. When an inner-shell electron is ejected, the atom is left in an excited state. This atom can then de-excite by filling the vacancy with an electron from an outer shell. During this process, either an x-ray is emitted that has an energy equal to the difference between the two energy levels or an Auger electron is emitted. ${ }^{6,8}$

In XRF analysis, the secondary x-rays are detected for qualitative and quantitative analysis of the sample. Each element has a series of characteristic x-ray lines whose energy is dependent on the shell from which the electron was ejected. The relationship between the wavelength of the x-ray $(\lambda)$ and the atomic number $(Z)$ of the excited element can be described by Equation 1.1, which was established by Moseley. ${ }^{7}$

$$
\frac{1}{\lambda}=K\left[Z-\sigma^{2}\right]
$$

## Equation 1.1

where K is a constant for each spectral series and $\sigma$ is a shielding constant with a value usually just less than 1.

The first, and usually most intense, set of lines is the K-lines, which are emitted when an electron from a higher energy shell fills a vacancy in the innermost shell (the Kshell) (Figure 1.2). ${ }^{6}$ The K-lines offer the highest-energy x-rays emitted from an element and are most often used in quantitative analysis. Next come the L-lines (due to a vacancy in the L-shell). These lines are lower in energy than the K-lines and, while there may be a few K-lines for a particular element, there can be several L-lines, often forming multiplets. The L-lines are usually only used for analysis when the K-lines of a particular element are too high in energy to be excited by the x-ray source and cannot be detected. The last major set of x-ray lines are the M-lines (due to a vacancy in the M-shell). These are the lowest in energy and are often overlapping with K-lines and L-lines from other elements, making quantification difficult.

Figure 1.2 Electronic transitions in the atom for characteristic $x$-ray lines.


The intensity of the x-ray fluorescence of the sample is dependant on two major factors: the excitation yield and the fluorescence yield. ${ }^{8}$ The excitation yield describes the probability of photoelectric absorption of the incoming radiation by the sample. This is dependent on the energy of the incoming radiation (hv), the binding energy of the electron (related to the Coulomb interaction between the electron and the nucleus), and atomic number. This relationship can be easily seen in Equation 1.2: ${ }^{8}$
$\tau_{p h}=a \frac{Z^{5}}{(h v)^{3}}$
Equation 1.2
where Z is the atomic number and $a$ is a proportionality constant. $\tau_{\mathrm{ph}}$ is the cross-section, or probability of photon absorption for an individual atom. ${ }^{8}$ When working with larger samples, the probability can also be expressed in terms of a mass (photoelectric) absorption coefficient, $\mu_{\mathrm{ph}}$, which describes the probability of photon absorption per unit mass of the absorber (Equation 1.3): ${ }^{8}$
$\mu_{p h}=a \frac{Z^{5}}{(h v)^{3}} \frac{N_{A}}{A}$
Equation 1.3
where $\mathrm{N}_{\mathrm{A}}$ is Avogadro's number and A is the atomic mass of the sample. Based on these equations, the probability of photoelectric absorption increases with atomic number, and therefore binding energy, assuming everything else remains constant. However, because the excitation yield is also dependent on the mass number, A , which also increases as Z increases, the effect is somewhat modified. ${ }^{8}$ The dependence of excitation yield on the energy of the incoming radiation (hv) also plays a strong role.

The excitation yield is also strongly dependent on the energy of the incoming radiation. However, this relationship for a particular atomic number, $Z$, is not a continuous function, which might be thought considering the above equations. Instead, there are discontinuities when the energy of the incoming radiation is exactly the binding energy of the electron. For example, when $h \nu$ is just below the binding energy of an electron in the K-shell, there is a sharp decrease in the excitation yield for K electrons. Resonance behavior is also seen in the relationship between $\mathrm{h} v$ and the excitation yield. When the energy of the incoming radiation is just above the binding energy of the electron, there is a sharp increase in the excitation yield. The discontinuities are due simply to the fact that a vacancy can only be created when the energy of the photon (hv) is greater than or equal to the binding energy of the electron. This, as mentioned above, greatly affects the excitation yield for elements of high Z . Not only is the atomic mass, A, greater, reducing the excitation yield slightly, but higher energies of incoming radiation are needed in order to excite the sample. This means that for high Z elements, the probability of photoelectric absorption ( $\mu_{\mathrm{ph}}$ ) is several orders of magnitudes lower than for low Z elements. This most strongly affects the K-lines for high Z elements, though
the L-lines and M-lines follow the same trend. This also means that for high Z elements, the excitation yield is often higher for L-shell electrons than for K-shell electrons. ${ }^{8}$

The fluorescence yield is the ratio of x-ray photons emitted from the sample to the number of vacancies caused by the excitation. This ratio describes the relative efficiency of the two opposing de-excitation processes that can occur after excitation: emission of an x-ray photon or emission of an Auger electron. The fluorescence yield ( $\omega$ ) is defined for each emission series as the number of x-ray photons ( n ) of that series emitted divided by the total number of vacancies $(\mathrm{N})$, shown in Equation 1.4 for the K series. ${ }^{6}$

$$
\omega_{K}=\frac{\Sigma(n)_{K}}{N_{K}}=\frac{n\left(K \alpha_{1}\right)+n\left(K \alpha_{2}\right)+n\left(K \beta_{1}\right)+n\left(K \beta_{2}\right)+\cdots}{N_{K}}
$$

Equation 1.4

The fluorescence yield is always less than one. As the atomic number increases, the probability of the emission of an Auger electron decreases, making the fluorescence yield increase. In addition, the fluorescence yield for K series x-rays is always larger than that for the $L$ series. ${ }^{8}$

X-ray lines chosen for quantitative analysis are based on optimizing and balancing both the excitation yield and the fluorescence yield. The excitation yield is dependant on the energy of the incoming radiation, as well as the composition of the sample, so the energy of the incoming radiation can be chosen to optimize the analysis for a particular set of elements or can be high enough to induce excitation in the majority of the elements of interest. The fluorescence yield, however, is not affected by any of the experimental parameters. Thus, based on the combination of the excitation yield and the fluorescence yield, the K series x-ray lines are often the best for quantification for lower Z elements, while the L series x -ray lines are more sensitive for the higher Z elements. ${ }^{8}$

Because x-rays can be easily scattered and absorbed in a sample, the sample matrix and size play an important role in XRF analysis. Firstly, the incident x-ray beam does not penetrate very far into samples, usually on the order of a few millimeters. ${ }^{7}$ For bulk analyses, it is crucial that the spot on the sample being analyzed is characteristic of the entire piece. Differences in composition at different depths, as well as different spots, may exist. Secondly, once the secondary x-rays are emitted, they then can be absorbed by other atoms in the sample and not make it out of the sample at all. ${ }^{7}$ If the secondary x rays have enough energy, they can excite another atom in the sample, which will then emit its own x-ray. Lastly, the matrix can scatter the x-rays so that they are not detected by the instrument. ${ }^{7-8}$ These challenges are usually taken into account by using matrixmatched standards for calibration of the instrument and by analyzing the sample at various locations to ensure homogeneity. When possible, thin samples can be used to ensure a reduction in the self-absorption of secondary x-rays.

Figure 1.3 is an example of an x-ray spectrum from a sample of obsidian analyzed by EDXRF. As stated before, the energies of the peaks in the spectra can be used to identify the elements in the sample. The intensities of the peaks are used in quantitative analysis.

Figure 1.3: XRF spectrum of obsidian


For accurate quantitative analysis, matrix-matched standards are analyzed under the same conditions to be used for the unknown sample and calibration curves are created for the elements of interest. Most XRF instrument software comes with a quantification program to create calibration curves and allow for quick, quantitative, multi-elemental analysis of the unknowns. Sample preparation is very simple; usually a clean, flat surface is the only requirement.

### 1.2.3 XRF analysis at MURR

In the Archaeometry group at MURR, there are two XRF instruments that are used regularly for analysis: an Elva-X benchtop XRF spectrometer and a Bruker handheld XRF spectrometer. Both of these instruments have energy-dispersive detection
systems. The Elva-X uses a tungsten anode for the x-ray source, while the Bruker uses a rhodium anode.

The concentrations of ten elements ( $\mathrm{K}, \mathrm{Mn}, \mathrm{Fe}, \mathrm{Zn}, \mathrm{Ga}, \mathrm{Rb}, \mathrm{Sr}, \mathrm{Y}, \mathrm{Zr}$, and Nb ) are typically determined, though other elements are occasionally added depending on the sample. In order to relate the intensity of the peak seen in the spectrum for each element, a calibration is created using well-characterized obsidian source standards. Since both the standards and the unknowns are run under the same conditions and have very similar matrices, factors that affect the intensity of each peak, such as instrumentation parameters and matrix effects, can be assumed to be the same for both. A linear regression model is then used to create the calibration curve to which unknown samples are compared, directly correlating intensity to concentration. Samples are typically analyzed for three to five minutes on the detector, set to aim for a dead time on the detector of about $10-15 \%$.

Because of the quick, nondestructive capabilities of XRF, it is a common choice of archaeologists and museums for the analysis of artifacts. Though fewer elements are determined at higher levels than with other techniques, the results are usually adequate for most purposes. In sourcing studies for lithics (which usually have the relatively homogenous composition needed for bulk analysis by XRF), usually a few elements are identified as important for analysis and source determination.

### 1.3 Neutron activation analysis

Neutron activation analysis (NAA) has become a prominent technique in archaeometric studies requiring precise, multi-elemental analysis. With the long list of elements that can be quantified and sensitivity (limited artifact destruction needed for
analysis), NAA has many qualities desired in a technique being applied to archaeological studies. This section will describe the theory of this technique and the current procedures in place at MURR.

### 1.3.1 A brief overview of NAA

Neutron activation analysis (NAA) is an analytical method that uses nuclear reactions to form radioactive products or excited states of the target nucleus, followed by the use of gamma spectroscopy for identification and quantification. ${ }^{4,9-10}$ NAA is the most common form of activation analysis and uses neutrons to induce radioactivity. Then, gamma rays emitted from the sample as the radioactive products decay are detected (Figure 1.4).

Figure 1.4 Neutron Activation Analysis


NAA has many advantages that make it a very useful technique for the multielemental analysis of archaeological artifacts. First are the very high sensitivities for the majority of the elements in the periodic table. These sensitivities can be on the same order as those for techniques such as ICP-MS. ${ }^{9}$ In addition, NAA provides simultaneous
multi-elemental analysis, so most of the elements of interest can be determined at one time. Many of the experimental parameters can be adjusted to provide a very accurate and precise measurement. Unlike methods such as ICP-MS, NAA requires very little sample manipulation, meaning that the method is unsusceptible to contamination. The analysis is also independent of matrix, so complex matrices can be analyzed without the extensive wet chemistry often required to remove matrix interferents. The matrix can contribute significantly to the background in the gamma ray spectrum if it contains significant quantities of elements that easily activate and have half-lives similar to or longer than the analytes of interest. If the elements of interest produce short-lived product isotopes, then the analysis can be very rapid, often completed in minutes. Also, if a large number of samples are being analyzed, there can be a low unit cost. ${ }^{9}$

Some of the disadvantages of NAA include the lack of information on chemical form and difficulty measuring some elements, especially the low-Z elements. In addition, NAA is only semi-nondestructive. While small samples can be placed whole in the vials and are not completely destroyed in the process, the analysis leaves the samples radioactive. Depending on the composition of the sample, it can remain radioactive for many years and may not be returned to the archaeologist or museum for that time. Another drawback is the limited access to neutron facilities, though many research facilities have programs to provide neutron activation analysis to archaeologists. ${ }^{11-14}$

### 1.3.2 Types of NAA

There are several types of NAA that can be performed. The most common that is applied to archaeometric studies is instrumental neutron activation analysis (INAA).

Here, the sample is simply irradiated and the delayed gamma rays counted, without any additional sample processing. Other types of NAA include radiochemical NAA (RNAA), where the analyte of interest is separated from the matrix after irradiation and before the sample is counted, and prompt gamma NAA (PGNAA), where the prompt gamma rays emitted from the sample during irradiation are detected. ${ }^{9}$ For the studies described later in this dissertation, INAA was used for the analyses.

NAA can also be divided by the energy of the neutrons being used for the irradiation. Thermal NAA (TNAA) is the most common and offers the best sensitivity for most elements in the period table. ${ }^{9}$ In TNAA, reactions involving thermal neutrons with a mean energy of about 0.025 eV are used for the analysis. High fluxes of thermal neutrons are typically generated from a fission reactor and moderation of the fission spectrum neutrons. Most of the reactions that occur with thermal neutrons are called ( $\mathrm{n}, \gamma$ ) reactions, where a neutron is absorbed by the target nucleus, followed by the emission of a prompt gamma ray. Because slower neutrons are more likely to interact with the nucleus of the atom, $(\mathrm{n}, \gamma)$ reactions tend to have very high probabilities (cross sections).

Epithermal NAA uses reactions involving the epithermal (also called epicadmium) neutrons that have energies between 0.1 eV and $1 \mathrm{keV} .{ }^{9}$ For ENAA, a cadmium or boron shield around the samples is used to reduce the flux of thermal neutrons from the reactor interacting with the samples, but allows the higher-energy neutrons to pass through. ENAA is mostly used to increase the sensitivity for elements that have larger cross sections for reactions with epithermal neutrons compared to thermal neutrons and to reduce the background due to elements that are high in abundance and have large thermal cross sections.

The last type is fast NAA (FNAA), where fast neutrons of energies greater than 0.5 MeV interact with the sample. ${ }^{9}$ These fast neutrons have the ability to cause different reactions other than $(\mathrm{n}, \gamma)$. An $(\mathrm{n}, \mathrm{p})$ reaction involves a neutron being absorbed, followed by the emission of a proton; $(n, \alpha)$ emits an alpha particle after the absorption of the neutron; and ( $\mathrm{n}, 2 \mathrm{n}$ ) emits two neutrons as a result. These reactions are less probable than thermal and epithermal neutrons due to the reduced wavelength of the neutron, and the barriers (Coulomb and angular momentum) associated with particle emission from the compound nucleus. FNAA can be performed using a cadmium shield around the samples, just as in ENAA, or by using a reactor neutron source that provides high energy neutrons. One common form is $14-\mathrm{MeV}$ INAA, where a small accelerator known as a neutron generator is used to produce $14-\mathrm{MeV}$ neutrons from the reaction of a deuterium beam with a tritium target. FNAA works best for lighter elements and can provide better sensitivity than typical TNAA.

### 1.3.3 Basic principles of NAA

As stated before, NAA offers the ability to fine-tune the analysis for the desired analyte(s) while offering the ability to determine the concentrations of many elements simultaneously. This is because the activity produced during the irradiation of the sample is dependant on variables such as the number of target nuclei $(\mathrm{n})$, the cross-section ( $\sigma$, measured in barns or $10^{-24} \mathrm{~cm}^{2}$ ) and the decay constant $(\lambda)$ of the target, the neutron flux $\left(\varphi\right.$, neutrons $\left./ \mathrm{cm}^{2} / \mathrm{s}\right)$, and the irradiation time $(\mathrm{t})$, according to the following equation: $:^{9-10}$ $A=n \phi \sigma\left(1-e^{-\lambda t}\right)$

Equation 1.5

Isotopes of the elements of interest are usually chosen to maximize the amount of activity that can be detected. Thus, isotopes that have larger abundances (so a larger n) and higher cross sections, with appropriate decay constants, are usually chosen. The irradiation time can be adjusted as well to maximize the sensitivity for the isotope.

The neutron flux in a light-water reactor, such as MURR, contains mostly thermal neutrons. Epithermal and fast neutron fluxes are usually about 2-3\% and 7-10\% of the thermal neutron flux, respectively. ${ }^{5}$ Thus, most of the neutrons interacting with the sample are thermal neutrons. Depending on the type of reaction that is of interest (using thermal, epithermal, or fast neutrons), the flux can also be adjusted with the use of cadmium or boron shields to reduce the thermal flux. ${ }^{9-10}$ This can be useful in that some isotopes have larger cross sections for epithermal neutrons than thermal neutrons. Fast neutron reactions tend to have smaller cross sections than either thermal or epithermal neutron reactions. For example, fast neutron reactions tend to have cross sections on the order of millibarns, while thermal and epithermal neutron reactions are usually on the order of barns. However, fast neutron reactions can still provide better sensitivity for lighter elements, especially when the thermal neutron flux has been reduced using cadmium or boron shields. ${ }^{10}$

The choice of irradiation time is based on several factors. If the isotopes of interest are short-lived with larger cross-sections, then short irradiation times are usually chosen so as to reduce the amount of longer-lived isotopes produced which might interfere with the measurement of the short-lived species. ${ }^{5,10}$ However, if the isotopes have smaller cross sections, then longer irradiation times are chosen to maximize the
amount of the radioactive species that is produced. Irradiation times can be anywhere from seconds to days.

Other factors that can be adjusted to increase the sensitivity for the analytes are the decay and count times. Once the radioactive species is produced, then it will immediately begin to spontaneously and randomly decay. The overall rate of decay is described by the following equation:

$$
A_{t}=A_{0} e^{-\lambda t}
$$

Equation 1.6
where $\lambda$ is the decay constant, which can be calculated from the known half-life of the isotope:

$$
t_{1 / 2}=\frac{\ln (2)}{\lambda}
$$

Equation 1.7

This decay can be used to the advantage of the analyst. If the isotope is short-lived, then the sample can either be counted immediately, or is allowed to decay a short amount of time, usually to make the sample less hazardous to handle and to reduce the amount of dead time in the detector. However, if the isotope of interest is long-lived, then the short-lived isotopes can be allowed to decay away before counting, reducing the overall activity of the sample, the background in the gamma-ray spectrum, and the amount of potential interferents in the sample. ${ }^{5}$ The count time is also chosen accordingly. For short-lived isotopes that are decaying rapidly, shorter count times are used. However, for isotopes that are decaying more slowly, longer count times are needed to achieve good statistics for the counts.

### 1.3.4 Gamma spectroscopy and quantitative analysis

Delayed gamma rays emitted from the activated sample are detected for identification and quantification of the components. Gamma rays are high-energy photons that are emitted from the nucleus. Because they have no mass or charge, they tend to interact with matter in an all-or-nothing way. Denser materials, like lead, are needed to increase the probability of interaction. ${ }^{9}$ Common detectors for gamma rays are scintillation detectors such as sodium iodide ( NaI ) and semiconductor detectors such as high-purity germanium (HPGe). ${ }^{4}$ Because so many elements are being analyzed simultaneously in NAA, the HPGe detectors are most often used because of their high resolution (typically around $1-2 \mathrm{keV}$ ). ${ }^{4}$ An example gamma ray spectrum achieved from a sample of obsidian is shown in Figure 1.5 below.

Figure 1.5 Gamma spectrum acquired from NIST SRM 278 (obsidian)


Because the gamma rays are characteristic of the isotopes that emitted them, they can be used to qualitatively determine the composition of the sample. ${ }^{9-10}$ Typically, a peak search is performed on the spectrum and then the energies of those peaks are compared to a table of gamma ray energies for the identification of the isotopes. The gamma rays are chosen for the analysis based on their abundance (branching ratio, BR ) and energy. ${ }^{10}$ Because of the background due to the Compton scattering of all the gamma rays emitted from the sample, gamma rays with low energy are sometimes difficult to see above background and yield a less sensitive signal, even though they typically have a higher efficiency than higher energy gamma rays. The number of counts in the photopeak $(\mathrm{R})$ is dependent on the activity produced during the irradiation (A, calculated using Equation 1.1), the efficiency of the detector ( $\varepsilon$ ), the abundance of the gamma ray (BR), and the count time $(\mathrm{t})$, according to Equation 1.8. . $^{-10}$

$$
R=A_{0} \varepsilon(B R) \int_{t 1}^{t 2} e^{-\lambda t}
$$

Equation 1.8

For quantitative analysis using the gamma spectrum, the total area under the peak is calculated, and then the background is subtracted to find the total counts due to the sample ( $\mathrm{R}_{\text {tot }}$ ). The intensity of the observed signal is directly proportional to the quantity of that isotope in the sample, which can then be related back to elemental concentration. However, since quantities such as the neutron flux are difficult to measure with good accuracy and to account for uncertainties in the cross section for the isotope, the comparator method is typically used for quantification rather than Equation 1.4. In the comparator method, a standard of a similar matrix and known concentrations of the
analytes is irradiated under the same conditions as the unknown sample. Then, the mass of the analyte in the unknown can be determined using the following equation: ${ }^{4,9}$

$$
\frac{R_{s t d}}{R_{\text {sam }}}=\frac{W_{s t d}\left(e^{-\lambda t}\right)_{s t d}}{W_{s a m}\left(e^{-\lambda t}\right)_{s a m}}
$$

## Equation 1.9

where R is the count rate for either the standard (std) or the sample (sam), W is the mass of the element, and $t$ is the decay time. If the decay time and the count time are identical for the standard and the sample, then a slightly different form of the equation can be used: ${ }^{5}$

$$
\begin{equation*}
c_{\text {sam }}=c_{\text {std }} \frac{m_{\text {std }} R_{\text {sam }}}{m_{\text {sam }} R_{\text {std }}} \tag{Equation 1.10}
\end{equation*}
$$

where c is the concentration of the element, m is the mass of the sample, and R is the count rate.

As stated before, NAA can offer very high sensitivities for many elements in the periodic table. Detection limits for NAA can be in the nanogram range for single elements in simple matrices. Detection limits calculated as three times the standard deviation of the background, which is the probability that a signal can be observed above background with $95 \%$ confidence. ${ }^{5}$ Several standards are usually analyzed along with the samples: one is used as a standard for the quantification of the element concentrations and the others are used as quality controls.

### 1.3.5 Potential interferences

As in any analytical method, the problem of interferences must be considered. For NAA, there are several different types of interferences. Primary interferences are when the desired product radionuclide can be produced in more than one way., ${ }^{4,9}$ For
example, scandium-46 can be produced via an ( $\mathrm{n}, \gamma$ ) reaction from $\mathrm{Sc}-45$ or from an (n,p) reaction from Ti-46 with fast neutrons. Secondary interferences are when the desired product radionuclide is produced by the decay of another product nuclide. ${ }^{4,9}$ So, for example, $\mathrm{Sc}-47$ can be produced from an ( $\mathrm{n}, \mathrm{p}$ ) reaction from Ti-47 and fast neutrons, but it is also produced from the beta decay of $\mathrm{Ca}-47$, a product of an $(\mathrm{n} . \gamma)$ reaction from Ca46. These two types of interferences are mostly avoided by choosing products that can only be produced via the desired reaction from the target nuclide. A third type of interference can occur in the gamma spectroscopy. ${ }^{9}$ Two radionuclides may emit gamma rays of the same energy or of similar energies that cannot be resolved. This can be resolved by using other gamma rays that are emitted from that same radionuclide for quantification. For example, mercury can be measured via NAA using the Hg -202 ( $\mathrm{n}, \gamma$ ) $\mathrm{Hg}-203$ reaction $(\sigma($ thermal $)=4.89$ barns $) . \mathrm{Hg}-203$ then decays with a half-life of 46.6 days and emits a single gamma ray of $279.2 \mathrm{keV}(81 \%)$. At the same time, selenium is often measured via NAA using the $\mathrm{Se}-74(\mathrm{n}, \gamma) \mathrm{Se}-75$ reaction $(\sigma($ thermal $)=51.8 \mathrm{~b}, \mathrm{t}$ 1/2 $(S e-75)=119.78$ days). Se-75 emits a gamma ray of 279.54 keV that has a $24.99 \%$ abundance, which cannot be resolved from the $\mathrm{Hg}-203$ peak. For the measurement of selenium, the other gamma rays emitted, particularly one at 264.66 keV ( $58.90 \%$ ), need to be used for the analysis. For the measurement of mercury in the presence of selenium, another reaction altogether needs to be considered, even though the sensitivity will be lowered.

### 1.3.6 NAA at MURR

The Archaeometry Group at MURR handles hundreds of samples on a regular basis that have been submitted for elemental analysis via NAA. This section will describe the standard procedures used for NAA for these samples from preparation to analysis, focusing on the procedures for obsidian and limestone. ${ }^{5}$ The concentrations of 28 elements are typically determined during the common NAA procedure.

Sample preparation is relatively simple. For obsidian, clean samples that are free of cortex (the weathering that can appear on the outside of the obsidian) are broken into pieces that can fit inside the irradiation vials. For limestone, the outside of the sample piece is cleaned by removing the outer layer using a dremmel tool, then the sample is crushed into a powder, using an agate mortar and pestle, and the powder is then weighed into the irradiation vials. Sample weights for analysis usually range between $50-100 \mathrm{mg}$ for short irradiations and 100-200 mg for long irradiations.

The analysis of samples is divided up into two parts: a procedure for the analysis of the short-lived isotopes that uses the pneumatic tube system in place at the reactor and a procedure for the analysis of longer-lived isotopes which involves placing the samples inside an aluminum can in the flux trap of the reactor. For the short irradiation, samples are weighed into polyvials that are then placed inside a rabbit. The sample is irradiated in the reactor for five seconds, allowed to decay for 25 minutes, and then counted on an HPGe detector for 12 minutes. The elements determined from the short irradiations are aluminum $\left({ }^{28} \mathrm{Al}, \mathrm{t}_{1 / 2}=2.24 \mathrm{~min}\right)$, barium $\left({ }^{139} \mathrm{Ba}, \mathrm{t}_{1 / 2}=83.06 \mathrm{~min}\right)$, dysprosium $\left({ }^{165} \mathrm{Dy}, \mathrm{t}_{1 / 2}=\right.$ $2.33 \mathrm{~h})$, potassium $\left({ }^{42} \mathrm{~K}, \mathrm{t}_{1 / 2}=12.36 \mathrm{~h}\right)$, manganese $\left({ }^{56} \mathrm{Mn}, \mathrm{t}_{1 / 2}=2.58 \mathrm{~h}\right)$, sodium $\left({ }^{24} \mathrm{Na}, \mathrm{t}_{1 / 2}\right.$ $=14.96 \mathrm{~h})$, and vanadium $\left({ }^{52} \mathrm{~V}, \mathrm{t}_{1 / 2}=3.74 \mathrm{~min}\right)$.

For the long irradiation, samples are weighed out into high purity quartz vials. Samples are placed in the reactor for a 48-hour irradiation. After this, the samples are counted twice: once for 30 minutes after a 7-day decay, and again for 168 minutes after a 28-day decay. The first count determines elements that have a mid-range half-life [including lanthanum $\left({ }^{140} \mathrm{La}, \mathrm{t}_{1 / 2}=40.27 \mathrm{~h}\right)$, lutetium $\left({ }^{177} \mathrm{Lu}, \mathrm{t}_{1 / 2}=6.73 \mathrm{~d}\right)$, neodymium $\left({ }^{147} \mathrm{Nd}, \mathrm{t}_{1 / 2}=10.98 \mathrm{~d}\right)$, samarium $\left({ }^{153} \mathrm{Sm}, \mathrm{t}_{1 / 2}=46.28 \mathrm{~h}\right)$, uranium $\left({ }^{239} \mathrm{~Np}, \mathrm{t}_{1 / 2}=2.36 \mathrm{~d}\right)$ and ytterbium $\left.\left({ }^{175} \mathrm{Yb}, \mathrm{t}_{1 / 2}=4.185 \mathrm{~d}\right)\right]$ and the second count determines the elements that produce long-lived isotopes [cerium $\left({ }^{141} \mathrm{Ce}, \mathrm{t}_{1 / 2}=32.5 \mathrm{~d}\right)$, cobalt $\left({ }^{60} \mathrm{Co}, \mathrm{t}_{1 / 2}=5.27 \mathrm{y}\right)$, cesium $\left({ }^{134} \mathrm{Cs}, \mathrm{t}_{1 / 2}=2.06 \mathrm{y}\right)$, europium $\left({ }^{152} \mathrm{Eu}, \mathrm{t}_{1 / 2}=13.54 \mathrm{y}\right)$, iron $\left({ }^{59} \mathrm{Fe}, \mathrm{t}_{1 / 2}=44.5 \mathrm{~d}\right)$, hafnium $\left({ }^{181} \mathrm{Hf}, \mathrm{t}_{1 / 2}=42.39 \mathrm{~d}\right)$, rubidium $\left({ }^{86} \mathrm{Rb}, \mathrm{t}_{1 / 2}=18.63 \mathrm{~d}\right)$, antimony $\left({ }^{124} \mathrm{Sb}, \mathrm{t}_{1 / 2}=60.2\right.$ d), scandium $\left({ }^{46} \mathrm{Sc}, \mathrm{t}_{1 / 2}=83.79 \mathrm{~d}\right)$, strontium $\left({ }^{85} \mathrm{Sr}, \mathrm{t}_{1 / 2}=64.84 \mathrm{~d}\right)$, tantalum $\left({ }^{182} \mathrm{Ta}, \mathrm{t}_{1 / 2}=\right.$ $114.43 \mathrm{~d})$, terbium $\left({ }^{160} \mathrm{~Tb}, \mathrm{t}_{1 / 2}=72.3 \mathrm{~d}\right)$, thorium $\left({ }^{233} \mathrm{~Pa}, \mathrm{t}_{1 / 2}=27.0 \mathrm{~d}\right)$, zinc $\left({ }^{65} \mathrm{Zn}, \mathrm{t}_{1 / 2}=\right.$ $244.26 \mathrm{~d})$ and zirconium $\left.\left({ }^{95} \mathrm{Zr}, \mathrm{t}_{1 / 2}=64.02 \mathrm{~d}\right)\right]$.

A peak search and quantification is performed using commercial software (Canberra), which detects the presence of a peak and then determines the area of the peak by fitting a Gaussian curve to it. In situations where several peaks seem to overlap, the program deconvolutes the multiplet and fits Gaussian curves to each of the contributing peaks. This program also calculates the peak areas and their uncertainties, the background that has been subtracted, and constants using the data from the standards for the calculation of concentrations of the analytes in the unknowns.

Several standards are analyzed along with each batch of samples. For obsidian, three standards are run with each set of samples, including one fly ash standard (SRM 1633a), one obsidian standard reference material (SRM 278), and one obsidian source
standard (JJO, from the Alca source in Peru), which is used for quality control. For limestone, the same standards are used, except instead of the JJO obsidian standard, the Ohio Red clay standard is used. The precision obtained using this method for obsidian and limestone is usually about $5 \%$ RSD or less for most elements. The limits of detection for the elements usually determined in obsidian are listed in Table 1.1.

NAA offers may advantages for the multi-elemental analysis of archaeological samples. The small sample size needed, the high sensitivity for many of the elements of interest, the small chance for contamination due to the limited sample preparation, and the ability to analyze many elements simultaneously offers a large amount of accurate and precise data that can be used for the characterization of the samples. In sourcing studies, data for a large number of elements allows for more discrimination among sources. This can lead to a stronger ability to differentiate between source groups, which can allow archaeologists to better answer questions about the movement of raw materials during prehistory.

Table 1.1 95\% confidence level detection limits for elements analyzed in obsidian

| ELEMENT | TYPICAL DETECTION <br> LIMIT |
| :---: | :---: |
| Na | 100 ppm |
| Al | 500 ppm |
| Cl | 25 ppm |
| K | 700 ppm |
| Sc | 0.001 ppm |
| Mn | 1 ppm |
| Fe | 50 ppm |
| Co | 0.02 ppm |
| Zn | 1 ppm |
| Rb | 1 ppm |
| Sr | 10 ppm |
| Zr | 15 ppm |
| Sb | 0.02 ppm |
| Cs | 0.02 ppm |
| Ba | 10 ppm |
| La | 0.5 ppm |
| Ce | 0.1 ppm |
| Na | 1 ppm |
| Sm | 0.02 ppm |
| Eu | 0.005 ppm |
| Tb | 0.05 ppm |
| Dy | 0.5 ppm |
| Yb | 0.1 ppm |
| Lu | 0.1 ppm |
| Hf | 0.01 ppm |
| Ta | 0.03 ppm |
| Th | 0.05 ppm |
| U | 0.5 ppm |
|  |  |

### 1.4 Comparison of XRF and NAA

Though XRF analysis and NAA can offer similar information about the elemental composition of artifacts, these techniques can also serve different purposes. For sourcing studies, it is important to have the source thoroughly characterized and to have as much information as possible about the source. This is where the advantages of NAA become useful. NAA can offer results for more elements with better accuracy and precision than XRF analysis. The results from NAA can then be examined to find which elements are important for distinguishing between two sources (most often by using multivariate analysis, discussed in the next chapter). In some cases, the differences in composition are so slight that more precise data are crucial. This especially applies when new sources or new regions are being characterized. In addition, the sources are usually geological samples, which are not as precious as archaeological artifacts, so destructive analysis is easier to perform.

XRF has is better suited for the analysis of artifacts as it requires less preparation and alteration of the samples. Once the composition of a source has been characterized by NAA, then the source samples can be analyzed by XRF to provide a comparative database for artifacts. The elements of interest for each source become the focus for the analysis. XRF can also be useful for the initial screening of artifacts. If an artifact has a similar chemical composition to two sources that are very similar, then those samples can undergo NAA at a later time. Since portable XRF units have become so widely available, the instrument can then be taken to the sample, rather than the sample needing to be transported to the lab.

In the research described in later chapters, NAA was used to characterize the sources of Egyptian limestone and Kenyan obsidian. Since only geological samples were considered in these studies, the semi-destructive nature of NAA was not a problem. In addition, it has been shown that limestone can be very difficult to source unless very precise data is used (discussed in detail in Chapter 3). Obsidian from many parts of the world has been found to work almost equally well with both XRF and NAA, but NAA was used here since the new sources investigated in this work are not as well characterized as obsidian sources from North, Central and South America, in order to fully examine the differences between the sources.

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# Statistical Methods of Data Analysis and Their Application to 

## Provenance Studies

### 2.1 Introduction

Data analysis is a critical part of the analysis process. When working with analytical techniques that produce large amounts of data, statistical techniques can help determine and clearly quantify relationships between different independent and dependent variables including subtle underlying relationships between several variables. In this chapter, several statistical methods will be described, as well as examples given of their use. The methods used include cluster analysis and principal component analysis. These methods have been applied to the various projects described throughout this dissertation.

### 2.2 Statistics in archaeometry

The use of statistical techniques applied to archaeological and chemical analysis is so diverse that it would prove difficult to discuss every application individually here. Thus, this section will focus on the applications relevant to later chapters of this work, especially those methods with application to sourcing archaeological materials.

Since their infancy, provenancing studies have relied on rigorous statistical methods to help distinguish source groups and characterize samples. ${ }^{1}$ The goal of these studies is to characterize samples in order to tell one group from another and how these groups relate to each other. This is often achieved through the use of statistical methods that look for the variance within a data set, as well as the similarities and differences between samples in that data set. ${ }^{1-2}$ Multivariate techniques, such as principal component analysis (PCA) and clustering analysis have proved especially useful. A wide variety of materials have been studied in this way including metals, ceramics, obsidian, limestone, chert, and other lithic materials. ${ }^{3}$ These materials range from being relatively easily sourced, such as obsidian ${ }^{2}$, to those more difficult due to the potential for chemical changes between the source and the artifact, such as ceramics. ${ }^{3}$

In order to be sourced, a material must follow the provenance postulate, which states that the variance within a source must be significantly less than the variance between different sources. ${ }^{1}$ If the variance within the source is too large or the variance between two sources is too small, then the exact source of a sample cannot be accurately or precisely determined with a good degree of confidence. Materials such as obsidian and ceramics have been found to follow the provenance postulate in many parts of the world. ${ }^{1-2}$

Generally, sourcing studies focus on the variation of the elemental composition of the material being studied, though other traits, such as color in the case of obsidian, have also been examined. In order to use the elemental composition, very accurate and precise multi-elemental techniques must be applied to the sample, such as x-ray fluorescence and neutron activation analysis (discussed more thoroughly in Chapter 1). Routine multi-
elemental analysis of samples can produce large amounts of data. For example, a standard NAA protocol can determine the concentrations for over 30 elements. Multiply that by the hundreds of samples being analyzed, and thousands of individual data points are created. The large amount of data can prove difficult to examine thoroughly, especially when some variables are correlated. In these cases, multivariate techniques, such as PCA have proved invaluable.

As stated before, obsidian has been proven to follow the provenance postulate and is relatively easily sourced. This is due to the fact that the artifact is chemically unchanged and so its composition should match that of the source very closely. In addition, simple scatter plots of the concentration of one element versus another can be used to distinguish source groups. ${ }^{2}$ Then distance calculations, such as the Euclidean distance calculations described later, are used to group samples with a $90 \%$ confidence. ${ }^{2}$ Here, the usefulness of multivariate techniques comes in pinpointing which elements will be able to best characterize the source groups and in further examining possible subgroups.

Limestone is quite different from obsidian in that it has been shown that simple scatter plots cannot distinguish source groups on their own. ${ }^{4}$ The nature of the matrix, a porous calcium carbonate whose general composition can change over time due to weathering, presents more difficulty in looking at the variance within each source and between different sources. In this case, clustering techniques, both hierarchical and nonhierarchical, and PCA allow the entire data set to be examined as each variable in relation to the others is considered. With a larger amount of information being utilized, the differences between sources can be more thoroughly characterized.

### 2.3 Standardization of data

When applying statistical techniques to large data sets where the data points may have a large variance in their magnitudes, the question of standardization or transformation becomes important. ${ }^{5}$ There are several possibilities for standardization and transformation and each has its advantages and disadvantages. Some of the most common are ${ }^{5}$ :

1) centering, $\left(y_{i j}=x_{i j}-\bar{x}_{j}\right)$, where $x_{i j}$ is any data point of sample $i$ and variable $j$ and $\bar{x}_{j}$ is the mean of variable $j$.
2) standardizing, $\left(y_{i j}=\left(x_{i j}-\bar{x}_{j}\right) / \mathrm{s}_{\mathrm{j}}\right)$, where $s_{j}$ is the standard deviation for variable j .
3) taking the $\log _{10},\left(y_{i j}=\log x_{i j}\right)$
4) ratio-transforming the data, $\left(y_{i j}=x_{i j} / x_{i p}\right)$

These can also be combined; for example, standardizing logged data.
Most statistical techniques applied to archaeometric data are looking for the variance within the data. Thus the results will be dominated by variables that have a large variance. Variables measured in different units or that have widely different magnitudes have a potential to skew the results. In these cases, standardization or transformation using $\log _{10}$ is usual. Centered data still presents the same problem where variables with large variances will dominate. Standardization will scale the data so that each variable has an equal variance and allows each to play an equal role in the analysis. Taking the $\log _{10}$ of the data will make the data be of a similar magnitude and comparable variance. The data is then said to be lognormal. This has been the typical choice for archaeometric data, though variables that have a large variance may still dominate. ${ }^{5-6}$

Standardization has often been discounted on the basis that is arbitrarily reweights the variables. This is not necessarily true, as standardization actually converts the data to units of standard deviation, and it has been shown that using standardized data still gets comparable results to that of log transformed data. ${ }^{5-6}$ For most of the work discussed in later chapters, standardized data was used for statistical analysis. This choice was made to reduce the effect of variance due to the wide range of concentration values, from tens of ppm to thousands of ppm .

### 2.4 Multivariate statistics

For the following discussions of multivariate statistics, a small data set containing five variables measured for 10 samples will be used as an illustration. The data set was standardized by dividing the difference between each datum and the mean of that variable by the standard deviation for that variable. The standardized data appears in the table below.

Table 2.1: Standardized Sample Data Set

| Sample | Al | $\mathbf{K}$ | $\mathbf{M n}$ | $\mathbf{N a}$ | $\mathbf{F e}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| A | 0.60145 | 0.72825 | -0.80663 | -0.15716 | -0.77335 |
| B | -0.09115 | -0.49447 | -0.86184 | -0.53460 | -0.89564 |
| C | 0.72148 | -1.20404 | -0.91608 | -0.34513 | -0.94371 |
| D | -1.12247 | -1.01227 | 1.47751 | 1.04605 | 1.42670 |
| E | 1.62583 | 2.32687 | 1.49388 | -2.06958 | 1.26339 |
| F | 0.39939 | 0.36260 | -0.90497 | -0.20627 | -0.92529 |
| G | 0.32126 | -0.22046 | 0.00096 | 0.03172 | 0.00073 |
| H | -1.70333 | -0.20859 | 0.68397 | 1.29877 | 0.77857 |
| I | -0.99226 | -0.33157 | 0.64183 | 1.18377 | 0.93233 |
| J | 0.23981 | 0.05367 | -0.80865 | -0.24758 | -0.86371 |

### 2.4.1 Variance, covariance, and correlation

Multivariate statistical methods are designed to group and classify objects as well as model relationships between the different variables in the data set. In doing so, common trends between variables become clear. Within the large data sets generated by multi-elemental analysis on archaeological artifacts are several variables that are positively correlated with each other. This is particularly prominent in geological samples, such as obsidian, for the rare earth elements (REEs). ${ }^{2}$ Thus, it is quite useful to calculate both the variance and the covariance in the data set.

Variance is defined as the spread of the samples around the mean for that variable and is calculated as the square of the standard deviation according to the following equation:

$$
\sigma_{i j}^{2}=\frac{\sum_{i=1}^{n}\left(x_{i j}-\bar{x}_{j}\right)^{2}}{n-1}
$$

where $n$ is the number of samples and $j=1 \ldots p$ for a matrix of $p$ variables. ${ }^{7}$ The measured sample variance $\left(\sigma_{\mathrm{m}}{ }^{2}\right)$ is the sum of the natural variance in composition $\left(\sigma_{\mathrm{n}}{ }^{2}\right)$ and the variance due to sampling and analytical errors $\left(\sigma_{a}{ }^{2}\right) .{ }^{2}$
$\sigma_{m}^{2}=\sigma_{n}^{2}+\sigma_{a}^{2}$
Equation 2.2

Covariance, like variance, describes the spread of samples around a mean, except here it is the shared variability between two variables around a common mean. This can be calculated for two variables, j and k , according to the following equation ${ }^{7}$ :
$\operatorname{cov}(j, k)=\frac{\sum_{i=1}^{n}\left(x_{i j}-\bar{x}_{j}\right)\left(x_{i k}-\bar{x}_{k}\right)}{n-1}$
Equation 2.3

When working with a large matrix of data that has $p$ variables, a variance-covariance matrix (C) can be calculated for all pairs of variables. ${ }^{7}$
$C=\left[\begin{array}{cccc}\sigma_{11}^{2} & \operatorname{cov}(1,2) & \ldots & \operatorname{cov}(1, p) \\ \operatorname{cov}(2,1) & \sigma_{22}^{2} & \ldots & \operatorname{cov}(2, p) \\ \vdots & \vdots & & \vdots \\ \operatorname{cov}(p, 1) & \operatorname{cov}(p, 2) & \ldots & \sigma_{p p}^{2}\end{array}\right]$
Equation 2.4

The correlation between two variables can also be calculated. Correlation coefficients are calculated for each pair of variables according to the following equation:

$$
r_{j k}=\frac{\operatorname{cov}(j, k)}{\sigma_{j} \sigma_{k}}
$$

## Equation 2.5

where $\sigma_{j}$ and $\sigma_{k}$ are the standard deviations of variables $j$ and $k$. Values for the correlation coefficient range from -1 to +1 , where -1 indicates an inverse relationship, 0 means no relationship, and +1 is perfect correlation. ${ }^{2,7}$ In geological samples, coefficients for the REEs can be as high as $0.9 .{ }^{2}$ When the data has been standardized, the covariance matrix and the correlation matrix are equivalent. ${ }^{7}$ The covariance and correlation matrices also form the basis for many multivariate statistical techniques such as principal component analysis (PCA), discussed in depth later.

### 2.4.2 Distance measurements and clustering methods

While covariance and correlation matrices provide information on the relationship between different variables in the dataset, other methods need to be used for initial studies of the relationship between samples. Measurement of the distance between samples in multidimensional space is often a first step in determining their similarity. ${ }^{2,6}$ There are
several distance measures commonly used: city-block (or Manhattan), Euclidean, and Mahalanobis. ${ }^{7}$ The city-block distance is calculated by

$$
d_{i j}=\sum_{k=1}^{n}\left|x_{i k}-x_{j k}\right|
$$

where n is the number of variables and $i$ and $j$ are the indices for samples $i$ and $j$. Graphically, this distance can be represented in two dimensions as the following:

Figure 2.1: City-block distance


One of the most useful distance measurements is the Euclidean distance. This is the straight-line distance between the two objects (the shortest distance between the two points). The Euclidean distance ${ }^{7}$ for a pair of objects in multidimensional space is:

$$
d_{i j}=\left[\sum_{k=1}^{n}\left(x_{i k}-x_{j k}\right)^{2}\right]^{1 / 2}
$$

Equation 2.7
and can be graphically represented in two dimensions by:

## Figure 2.2: Euclidean Distance



Many clustering methods (discussed in detail later) use either the Euclidean distance or the squared-mean Euclidean distance as a way of hierarchically grouping samples. The squared-mean Euclidean distance (SMED) ${ }^{2,5}$ is calculated by the following equation:
$d_{i j}^{2}=\frac{\sum_{k=1}^{n}\left(x_{i k}-x_{j k}\right)^{2}}{n}$

## Equation 2.8

The Mahalanobis distance is defined as the squared Euclidean distance between a sample and a group centroid, divided by the group standard deviation in that direction. ${ }^{2,6,7}$ Here, instead of looking at the relationship between two individual samples, the relationship between a cluster of samples and an individual sample is elucidated. Mahalanobis distance between the centroid of group X and sample $k$ is calculated as shown below.

$$
D_{k X}^{2}=\sum_{i=1}^{n} \sum_{j=1}^{n}\left[x_{i k}-\bar{x}_{i}\right] \cdot \mathrm{C}_{\mathrm{ij}} \cdot\left[x_{j k}-\bar{x}_{j}\right]
$$

Equation 2.9
where $\bar{x}_{i}$ and $\bar{x}_{j}$ are the means of variables $i$ and $j$, and $\mathrm{C}_{i j}$ is the $i j$ th element in the covariance matrix. The Mahalanobis distance is calculated using the covariance matrix, so any covariance between variables is taken into account. This distance measurement is especially useful when assigning samples to already existing groups (i.e., assigning artifacts to source groups). ${ }^{2}$ Other common uses of the Mahalanobis distance are to test the distance between two neighboring groups and to create confidence ellipses for clusters, typically at $90 \%$ or $95 \%$ confidence. ${ }^{2}$

As mentioned before, distance measurements are often used in cluster analysis. Cluster analysis endeavors to place samples in groups based on their similarity and distance measures are used as a way to calculate that similarity. ${ }^{7}$ The two major types of cluster analysis are hierarchical and nonhierarchical. ${ }^{7}$ Hierarchical cluster analysis is usually agglomerative and starts with individual samples, grouping them one at a time according to their distances. This can be graphically represented in dendrograms. Nonhierarchical cluster analysis partitions the samples independently of each other and does so without ordering them hierarchically. Here, the number of groups is determined initially and each sample is mathematically tested to prove its membership in the group.

Hierarchical cluster analysis is the more common in archaeological applications. ${ }^{2,}$ ${ }^{6}$ Samples are grouped together based on similarity, then those groups are linked to other groups, and the iterations continue until all the samples are related to each other. This is often depicted as a dendrogram, with the samples on the x -axis and the distance between those samples plotted on the y-axis. ${ }^{7}$ In most common clustering routines, the Euclidean or squared-mean Euclidean distance is used to determine similarity between two samples. ${ }^{2,7}$

To mathematically determine if the samples are similar enough to be linked, one of several different linkage methods can be used. These linkage methods, or clustering algorithms, each have their own advantages and disadvantages and can often produce different results for the clustering of the same data set. Several examples of the different linkage methods are given in Table C for linking groups A and B into a new cluster $k$ by comparing them to object $i .^{7}$

Table 2.2: Linkage methods ${ }^{7}$

| Method | Equation | Description |
| :---: | :---: | :--- |
| Single <br> Linkage | $d_{k i}=\frac{d_{A i}+d_{B i}}{2}-\frac{\left\|d_{A i}-d_{B i}\right\|}{2}=\min \left(d_{A i}, d_{B i}\right)$ | The shortest distance <br> between opposite clusters is <br> calculated. |
| Complete <br> Linkage | $d_{k i}=\frac{d_{A i}+d_{B i}}{2}+\frac{\left\|d_{A i}-d_{B i}\right\|}{2}=\max \left(d_{A i}, d_{B i}\right)$ | The largest distance <br> between opposite clusters is <br> calculated. |
| Unweighted <br> Average <br> Linkage | $d_{k i}=\frac{d_{A i}+d_{B i}}{2}$ | Average distance from <br> objects A and B is <br> calculated. |
| Weighted <br> Average <br> Linkage | $d_{k i}=\frac{n_{A}}{n} d_{A i}+\frac{n_{B}}{n} d_{B i}$ with $n=n_{A}+n_{B}$ | The number of objects (n) in <br> the clusters is used for <br> weighting the cluster <br> distances. |
| Centroid <br> Linkage | $d_{k i}=\frac{n_{A}}{n} d_{A i}+\frac{n_{B}}{n} d_{B i}-\frac{n_{A} n_{B}}{n^{2}} d_{A B}$ | The distance between the <br> centroids of the clusters is <br> calculated. |
| $d_{k i}=\frac{d_{A i}}{2}+\frac{d_{B i}}{2}-\frac{d_{A B}}{4}$ <br> Linkage | The median instead of the <br> centroid of each cluster is <br> used for calculation. |  |
| Ward's <br> Method | $d_{k i}=\frac{n_{A}+n_{i}}{n+n_{i}} d_{A i}+\frac{n_{B}+n_{i}}{n+n_{i}} d_{B i}-\frac{n_{i}}{n+n_{i}} d_{A B}$ | Clusters are combined based <br> on a minimum increase in <br> the within-group variance <br> from the centroid. |

Each of these linkage methods can produce different results because they stress different characteristics in the strategy for grouping samples. For example, single linkage can produce loosely-bound clusters, while complete linkage tends to form small, wellseparated clusters. ${ }^{7}$

The sample data set described above is used as an example of hierarchical clustering using the clustering routines in Matlab©. Distances between samples were determined using the Euclidean distance measures and then two different linkage methods were applied, the average linkage and Ward's linkage. The resulting dendrograms for both are shown below.

Figure 2.3: Dendrogram of standardized sample data set (Table 2.1) using average linkage


Figure 2.4: Dendrogram of standardized sample data (Table 2.1) using Ward's linkage


In both of these figures, similar groups are seen in the higher clustering levels, while differences between the two methods are more visible in the initial clusters formed. However, both methods suggest the presence of at least two clusters, with one or two outliers. The first cluster is made up of samples A, B, C, F, and J (and possibly G), while the second includes samples $\mathrm{D}, \mathrm{H}$, and I, with sample E as an outlier in that group.

Nonhierarchical clustering is another way to view potential groups within the data. One method for nonhierarchical clustering is K-means clustering. ${ }^{6}$ This method works as a partitioning method, where samples are placed within clusters that are mutually exclusive based on their similarities, rather than on proximities. Rather than creating structure in the data to represent the levels of similarity (as in a dendrogram), a single level of clusters is formed. The method defines each cluster by the samples belonging to
it and the centroid of the cluster. Samples are then added to the clusters so that the sum of the distance measures to the centroid in that cluster is minimized. Usually this is done iteratively, moving samples from cluster to cluster until all the distances from sample to centroid cannot be minimized any further. While this can be useful for clustering large amounts of data (since a large, complicated dendrogram is not produced), the number of clusters, K, must be determined beforehand. This means that using a different number of clusters to begin with can change the results of the clustering, sometimes dramatically. ${ }^{6}$

Again, the sample data set is used as an illustration. Here, K-means clustering was performed using the kmeans function in Matlab©. This function uses squared Euclidean distances to determine the distance from each point to the centroid of the cluster. The output is a vector that contains the cluster indices for each sample. Below are three tables showing the results of the K-means clustering when K (the number of clusters) is three, four, and five, respectively.

Table 2.3: Results of K-means clustering for standardized sample data set (Table 2.1)

| A) $\mathrm{K}=3$ |  | B) $\mathrm{K}=4$ |  | C) $\mathrm{K}=5$ |  |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Sample | Cluster | Sample | Cluster | Sample | Cluster |
| D | 1 | D | 1 | D | 1 |
| H | 1 | H | 1 | H | 1 |
| I | 1 | I | 1 | I | 1 |
| A | 2 | A | 2 | A | 2 |
| B | 2 | B | 2 | F | 2 |
| C | 2 | C | 2 | J | 2 |
| F | 2 | F | 2 | B | 3 |
| G | 2 | J | 2 | C | 3 |
| J | 2 | G | 3 | G | 4 |
| E | 3 | E | 4 | E | 5 |

It is apparent that all three analyses get similar results for some samples. Samples D, H, and I form one cluster that remains the same in all three approaches, while sample E is placed in a group on its own. K-means clustering with $\mathrm{K}=4$ seems to get results that most closely match those from the hierarchical clustering methods. When $\mathrm{K}=5$, further division is seen between samples $\mathrm{A}, \mathrm{F}$, and J , and samples B and C .

Though it is useful as a starting point in distinguishing groups, cluster analysis cannot be used alone as a quantitative measure for the differentiation of groups. ${ }^{2}$ Cluster analysis assumes that the variables in the data are uncorrelated, which is known not to be the case with archaeological data sets. ${ }^{2}$ If the data were truly uncorrelated, then the dendrograms might accurately represent the similarities between samples. However, when sourcing geological materials, it has been shown that many elements, such as the REEs, are correlated. Thus, cluster analysis can only be used as an initial step and should not be taken as faithfully representing differences between groups of samples. For truly accurate distinction between groups, other statistical techniques, such as principal component analysis, should be used.

### 2.4.3 Principal component analysis

Principal component analysis (PCA) is a statistical method which reduces the amount of variables needed to describe a data set when there is correlation present. ${ }^{7-10}$ It allows data from a high dimensional space to be projected into fewer dimensions (typically either two or three dimensional space) while maintaining the distances between each point. In addition, while the number of dimensions needed to graphically represent the data are reduced, the data points still include all of the original data. PCA does this by
finding the principal components (PCs), linear combinations of the original variables, which account for the variation in the data set. ${ }^{7}$ This technique is especially useful when looking at archaeometric data. Since there are so many variables, it is quite useful to be able to reduce the number of variables to a few that can describe the majority of the variance in the data set. It is also then important to relate these new variables to the original ones to understand how they affect the variance between groups within the data.

PCA takes a large data set with $p$ variables that may be correlated and linearly transforms it into $p$ uncorrelated variables so that the best clustering of groups within the data is achieved. ${ }^{5-7,10}$ The standardized data set is then projected onto the new set of axes created, where the first axis (principal component), PC1, accounts for most of the variation, the second (PC2) accounts for the next greatest amount of variation and is orthogonal to PC1, and so forth. The same number of PCs is produced as there were original variables. The linear transformation is done using the following equation for each principal component:
$P 1=a_{11} X_{1}+a_{12} X_{2}+a_{13} X_{3}+\ldots . a_{1 n} X_{n}$ for PC1,
Equation 2.10
$\mathrm{P} 2=\mathrm{a}_{21} \mathrm{X}_{1}+\mathrm{a}_{22} \mathrm{X}_{2}+\mathrm{a}_{23} \mathrm{X}_{3}+\ldots \mathrm{a}_{2 \mathrm{n}} \mathrm{X}_{\mathrm{n}}$ for PC 2,
etc, where a is the score for each variable and X is the standardized data. ${ }^{2,6,9}$

When PCA is performed on a particular data set, three matrices are typically returned. ${ }^{7-8}$ The first is the set of eigenvalues, which describe the amount of variance described by each principal component. The second is the loadings matrix which describes the influence of each variable on each principal component. The last is the scores matrix, which is the projection of the data set into principal component space.

PCA was performed on the sample data set above (Table 2.1) and will be used to illustrate the results of a PC analysis.

The eigenvalues, or the variance described by each PC, provide useful information about how the principal components describe the variance in the sample and they allow an informed number of principal components to be chosen for examination. ${ }^{8-9}$ The eigenvalues are the sum of the variances for each principal component. When the data is standardized, the total variance in the data set equals the sum of the variances of each PC. ${ }^{8}$ So, for the data set containing 5 standardized variables, $\sigma_{\mathrm{i}}^{2}=1$ for $i=1$ to 5 for each variable. When PCA was performed on the sample data set described above (Table 2.1), the following eigenvalues were obtained (Table 2.4).

Table 2.4: Eigenvalues of standardized sample data set (Table 2.1)

| Principal <br> Component | Eigenvalue | \% Variance <br> explained by each PC | \% Cumulative <br> variance |
| :---: | :---: | :---: | :---: |
| PC1 | 2.6555 | 53.11 | 53.11 |
| PC2 | 1.9976 | 39.95 | 93.06 |
| PC3 | 0.2824 | 5.648 | 98.71 |
| PC 4 | 0.06191 | 1.238 | 99.95 |
| PC5 | 0.00252 | 0.05048 | 100 |

Here, the sum of the eigenvalues also adds up to 5 .
The eigenvalues are particularly useful for determining the number of principal components that are appropriate for consideration for the analysis of the data set. Here, several factors are considered. First, the amount of variance explained by each PC is calculated by dividing each eigenvalue by the total variance (here, 5 ) and multiplying by $100 \%$ (Table 2.4). The cumulative variance is also calculated. Usually a fixed
percentage of variance is specified, for example $90 \%$, and the number of PCs needed to reach that cumulative variance are chosen. ${ }^{7}$ In this data set, the first two PCs explain about $93 \%$ of the variance, so only the first two PCs would be considered for further analysis. In addition, a Scree plot can also be used to determine the number of PCs. ${ }^{7-8}$ The Scree plot can either be created by plotting the eigenvalues for each PC (Figure 2.5a) or by plotting the percent variance explained for each PC (Figure 2.5 b ).

Figure 2.5: Scree plots from PCA of standardized sample data set (Table 2.1)
A) Eigenvalues vs. PCs

B) $\%$ Variance vs. PCs


In each of these plots, the amount of variance levels off when the appropriate number of PCs has been reached. These justify the choice of the first two principal components to describe the majority of the variance in the data set.

The loadings matrix describes how each of the variables contributes to the variance in the data set and thus to each PC. ${ }^{5,7-8}$ It also shows the correlations between each variable within the data set. The loadings are the coefficients $\left(\mathrm{a}_{\mathrm{ij}}\right)$ shown in equation 2.10 above. These coefficients weight the original variables to project them into principal component space. Shown below is a plot of the loadings values for each variable in the first two PCs.

Figure 2.6: Loadings plot from PCA of standardized sample data set (Table 2.1)


In this plot, the size of the value reflects the contribution to each PC and the direction (positive or negative) can give information on the relationship between variables. For example, aluminum and sodium both strongly contribute to the variance explained by PC1 but they are inversely correlated to each other. Thus it seems possible that sodium and aluminum can be used to differentiate between groups within this data set. Meanwhile, the second principle component is driven by variance from potassium, manganese, and iron, all of which are strongly intercorrelated with each other in this dimension. Another way of plotting the loadings is shown in Figure 2.7. ${ }^{7}$ Here, the strong correlation between manganese and iron is clearly visible as is the negative correlation between aluminum and sodium.

Figure 2.7: Loadings vector plot from PCA of standardized sample data set (Table 2.1)


The last matrix that results from PCA is the scores matrix. This matrix is the projection of the data onto the new axes defined by the principal components. ${ }^{7,9-10}$ PCA maintains the distance between each data point in multidimensional space, so the scores represent the same data as before, just with a redefined origin and set of axes. The scores are the new coordinates of the data. Figure 2.8 is a plot showing the scores values for PC2 plotted versus those from PC1 for the sample data set. These principal components were proven above to describe over $90 \%$ of the variance in the data set.

Figure 2.8: Scores plot from PCA of standardized sample data set (Table 2.1)


As can be seen in the plot, there are several distinct groups formed by the samples. For samples to have similar scores values, they must have similar values in the original data set. In this data set, samples A, B, C, F, and J form one group while D, H, and I for another. Samples E and G seem to be outliers in the data set. If the loadings vectors and the scores are plotted on the same graph, then the relationship between the samples and the principal components becomes even clearer (Figure 2.9). In this plot, the loadings vectors have been scaled up for clarity.

Figure 2.9: Plot of scores and loadings vectors from PCA of standardized sample data set (Table 2.1)


In this plot, it is clearly seen how aluminum and sodium drive the separation of the (A,B,C,F,J) group and the (D,H,I) group. The latter also has relatively high levels of iron and magnesium. High levels of potassium distinguish sample E from the rest of the samples, while moderate levels of all the measured elements separates sample G. Thus, not only are the differences between groups and similarities within groups more evident using the results from PCA but also these results can be related back to the original variables. ${ }^{7-9}$

PCA is applied to sourcing studies because the focus of those studies is to distinguish groups of similar elemental composition and to determine what elements
actually drive the variance between source groups. By applying PCA, large data sets can be reduced to much fewer variables, through analysis of the eigenvalues, which can then be used to study the variance within the data set. Plots of the loadings and the scores then describe how the original variables (the elemental concentrations) drive the variance and can be used on their own to distinguish between compositional groups. Finally, the scores plot itself shows how the samples in the data set are related to each other and the most probable compositional groups are formed.

### 2.5 Conclusion

The goal of studies involving the sourcing of geological materials is to accurately describe the chemical composition of the source so that the source of samples of unknown origin might be determined with confidence. In order to accomplish this, multivariate statistical techniques can be used on the large data sets that result from multi-elemental analysis to decipher the variations within source groups and between different source groups. ${ }^{2}$ Initial analysis using covariance matrices and cluster analysis can be used to determine the relationship between the measured variables and the individual samples, respectively. More rigorous statistical methods, such as principal component analysis, can then be used to accurately model the groups within the data. This allows for a complete characterization of the data set, using all the information available from the multi-elemental analysis.

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## Chapter 3:

## Sourcing Egyptian Limestone

### 3.1 Introduction

Neutron activation analysis (NAA) has performed an important role in sourcing ancient artifacts including ceramics and a wide variety of lithic materials, including flint, basalt and obsidian. When coupled with rigorous statistical methods, including principal component analysis and various clustering techniques, it can provide the precision needed to determine the source of the material with high confidence. However, this technique has not been fully exploited for the provenancing of Egyptian limestone sculpture. Limestone has played an important role in the building and sculpting history of many cultures all over the world, but sourcing studies have been mostly limited to France and other regions in Europe. ${ }^{1-3}$ A plethora of ancient Egyptian monuments and sculptures has been made from limestone and the Egyptian people have continued to use limestone as a building material since antiquity. Most Egyptian provenance studies have focused on the use of petrology rather than the chemical composition. ${ }^{4-5}$ This study examines the use of elemental analysis by INAA as a complementary method for sourcing Egyptian limestone. Here, a combination of the elemental concentration data and statistical methods is used to study the compositional differences between known ancient quarries.

### 3.1.1 The geology of limestone

Limestone is a sedimentary rock that is mostly calcium carbonate in the form of calcite. ${ }^{6}$ Most limestone is formed from the remains of microscopic marine species which have been compressed into the solid rock. The calcite needed to form limestone can also be deposited by evaporation of a solution rich in calcium carbonate. Limestone is a relatively porous and soft rock, making it easily carved. It is also widely available all over the world. These last two qualities are what made limestone a popular choice as a building material and for sculptures.

Because limestone is mostly made from the remains of small marine species, especially those with calcite shells, the mineral and fossil composition of limestone can vary from location to location. Different species in limestone can be identified through optical and scanning electron microscopy. These differences are often used to source the limestone, though analysis of the trace elemental composition has also been proven useful in sourcing studies involving limestone., 6-7

The chemical composition of the limestone is dependent on its formation and weathering history, just like the petrology of the sample. Over time, limestone that contains easily dissolved minerals will become more porous and wear away, while limestone that is more "weather-resistant" will remain behind. The more resistant limestone was frequently chosen by ancient people as being good building material. ${ }^{6}$ This deliberate choice limits the number of limestone formations that were actively used as quarries. However, each of these still has a unique formation and weathering history,
which supports the hypothesis that the chemical composition of these sources can act like a fingerprint.

### 3.1.2 Limestone sourcing studies

Sourcing studies for limestone began as a way to assign a provenance to limestone sculptures that have been moved or "lost their histories". ${ }^{1}$ Initial studies used a combination of petrological and compositional techniques to characterize and assign the unknown samples. Extensive research has been done in sourcing limestone from France and other areas of Europe using NAA. These studies, mostly performed at Brookhaven National Laboratory (BNL) by Holmes and colleagues, ${ }^{1-2,7-10}$ have shown that using the elemental composition and rigorous statistical techniques, including cluster analysis and multivariate analysis, the limestone sources can be distinguished. This has been especially useful in determining not only the limestone quarries used but also in tracing sculptures back to their original locations. ${ }^{2,11}$

In contrast to Europe, very little has been done on sourcing Egyptian limestone. Only three major papers have been published on the subject in the last 20 years. In addition, two out of the three papers focus mostly on the petrological analysis rather than elemental composition. The first major study done was by Meyers and van Zelst ${ }^{3}$ in 1977. Their work describes the analysis of limestone artifacts from Egypt and Spain. The Egyptian artifacts were from a variety of locations and it was assumed that each was made from local limestone deposits. However, it was found that the Egyptian limestone seemed to be relatively inhomogeneous in composition, since the variation between samples that came from the same location was about the same as the variation between
samples from different locations. Cluster analysis was applied to the samples and they found that only one small group of artifacts from Thebes seemed to be distinguishable from the rest of the samples. The other limestone artifacts were indistinguishable from each other compositionally and there was no consistent relationship between composition and geographic origin. However, one of the drawbacks of this study was that quarry samples were not analyzed along with the artifacts. Therefore it is possible that these artifacts were made from limestone from the same quarry or quarries close together that are similar in composition. A compositional study of the quarries is needed to confirm that there is no distinction between geographical sources and chemical composition. In addition, more rigorous statistical methods need to be applied and reported in order for statistical confirmation of these results.

The second major study of Egyptian limestone artifacts was published by Middleton and Bradley in 1989.5 In this study, a combination of petrographic examination using both optical and scanning electron microscopes (SEM) and mineralogical analysis using x-ray diffraction was used to characterize three major groups of Egyptian limestone. Again, mostly artifacts were studied and grouped according to their similar composition. However, this paper showed promising results: the sources were distinguishable from each other by the petrological analysis. Finer distinctions could also be made using the elemental composition determined using the energy-dispersive x-ray spectrometer attached to the SEM instrument. The three regions that could be distinguished were centered around Cairo, El Bersha, and the Thebes/Abydos region. However, very few geological source samples were analyzed and the few that had been sampled were painted reliefs. Quarry samples from geological
sources need to be analyzed in order for these results to be confirmed and to begin applying this method to unknown artifacts.

More recent work done on sourcing Egyptian limestone is that by Dr. Harrell of the University of Toledo. In his 1992 paper, ${ }^{4}$ Harrell describes a combination of thinsection petrography and x-ray fluorescence spectrometry for the analysis of quarry samples. These methods were applied to samples of limestone from 23 known quarries in Egypt, with a total of 28 samples analyzed. These known quarries are located in six main limestone formations in Egypt: Tarawan, Serai, Drunka, Minia, Salamut, and Mokattam formations (listed in order from north to south), all along the Nile. The main results of this study centered around distinguishing between these six major limestone formations. In a few instances, Harrell reports being able to distinguish specific quarries, but for the most part, only the general region can be determined for the limited number of limestone samples. The geochemical results were used mostly to confirm the petrological results, since there was not much variation seen between quarries for the elements analyzed.

At about the same time, a book by Rosemarie Klemm and Dietrich D. Klemm also considers the idea of sourcing limestone from Egypt. ${ }^{6}$ In the chapter on limestone, most of the emphasis is placed on using the petrology of the sample to determine the origin. However, a smaller section described the use of elemental analysis, performed by x-ray fluorescence, atomic absorption, and inductively-coupled plasma spectrometry, to also aid in distinguishing sources. The elements chosen to distinguish between sources were magnesium and strontium, which provided satisfactory results. These elements are rather mobile in the environment, but because of insufficient limits of detection for the
less mobile trace elements, they provided the best possible results. The authors also caution that in order to achieve the best possible separation between source groups, a large number of samples from each quarry location need to be analyzed. Analyzing many samples from each location provides information not just on the variation between source groups, but also characterizes the intra-source variation.

### 3.1.3 The current project

As a combination of neutron activation analysis and rigorous statistical methods have been proven useful in determining the source of limestone from other areas of the world, a similar technique is applied to quarry samples of Egyptian limestone. Since xray fluorescence was shown to be a useful complementary technique to the petrological analysis, NAA should also provide complementary information about the samples. In addition, NAA has the ability to precisely measure more elements, especially the rareearth trace metals that have been useful in sourcing studies of other lithic materials.

### 3.2 Methodology

Two sets of samples were used in this study. The first is a set of 75 samples that were inherited by the Archaeometry Group at MURR from BNL. These samples had little identification other than the list of the regions from which they came and they were already in powder form. Some limited data had been collected on the elemental composition using INAA at BNL. To make up for the missing elements, these samples were reanalyzed using standard INAA procedures (as described in Chapter 1) at MURR.

The second group of samples is a set of 65 samples acquired from Dr. Harrell at the University of Toledo. These samples were taken from many known ancient quarries along the Nile and had more detailed geographical information, including geospatial coordinates for the specific locations. About half of these samples came in powder form, while the other half were rock fragments. The fragments where cleaned by burring off the surface and rinsing in DI water. After drying in an oven overnight, they were then crushed into a powder using an agate mortar and pestle. The powders were placed overnight in an oven at $110{ }^{\circ} \mathrm{C}$ to dry the samples.

Samples of about 150 milligrams (for the short irradiation procedure) and 200 milligrams (for the long irradiation procedure) were weighed out into high-density polyethylene vials (shorts) and quartz tubes (longs). The samples were then analyzed via INAA according to standard procedures for limestone (as described in Chapter 1). Standards used in the analysis included SRM1633a (fly ash), SRM 278 (obsidian), and Ohio Red clay, which acted as a QC sample for the analyses. Once the data was collected, several multivariate statistical techniques were used to characterize the samples in the data set. Each set of samples was treated separately for the statistical analysis.

### 3.3 Statistical results and interpretation

### 3.3.1 The BNL samples

The Egyptian limestone samples from BNL came to the Archaeometry Group at MURR as part of a large collection of limestone samples, mostly from France, which had been previously analyzed by Lore Holmes and colleagues. However, among the plethora of information that came with the samples, there was very little about these particular
samples. There are 75 samples in total from five different quarry locations in Egypt:
Beni Hasaan (BH), Deir Abu Hennis (DAH), Maasara (M), Quseir el Amarna(QA), and Tura (T). Very little else is known about these samples. A map of the locations of these five quarries is shown in Figure 3.1.

Figure 3.1: Map of sampled quarry locations in Egypt (BNL samples)


As stated before, there were many missing values for certain elements, such as manganese, so these samples were reanalyzed at MURR. Then data (given in Appendix 2) were subjected to several multivariate statistical techniques. First, cluster analysis was performed on the data set, using Euclidean distances and single linkages. The resulting dendrogram is shown in Figure 3.2. From this, it can be seen that the samples form two main groups, along with a few outliers. In addition, samples from the same source tend to cluster together more than those from different sources.

Figure 3.2: Dendrogram resulting from hierarchical cluster analysis of Egyptian limestone samples (BNL samples)


K-means clustering was also performed on the data set, with $\mathrm{k}=2,3$, and 5 . The results from this analysis for $\mathrm{k}=3$ are shown in Table 3.1. Here the samples are still grouping in similar ways as before. When $\mathrm{k}=3$, the results almost perfectly match those achieved with hierarchical clustering, if groups 1 and 3 are combined. The separation between groups 1 and 3 is most likely due to a large intra-source variation.

Table 3.1: Results from k-means clustering of Egyptian limestone samples (BNL samples)

| anid | Group | IDX3 |
| :--- | :--- | ---: |
| EB0102 | M | 1 |
| EB0103 | M | 1 |
| EB0104 | M | 1 |
| EB0107 | M | 1 |
| EB0108 | M | 1 |
| EB0112 | M | 1 |
| EB0078 | T | 1 |
| EB0079 | T | 1 |
| EB0080 | T | 1 |
| EB0082 | T | 1 |
| EB0083 | T | 1 |
| EB0084 | T | 1 |
| EB0085 | T | 1 |
| EB0100 | M | 3 |
| EB0101 | M | 3 |
| EB0105 | M | 3 |
| EB0106 | M | 3 |
| EB0109 | M | 3 |
| EB0110 | M | 3 |
| EB0111 | M | 3 |
| EB0081 | T | 3 |
| EB0038 | BH | 2 |
| EB0051 | BH | 2 |
| EB0052 | BH | 2 |
| EB0053 | BH | 2 |
| EB0054 | BH | 2 |
| EB0055 | BH | 2 |
| EB0056 | BH | 2 |
| EB0057 | BH | 2 |
| EB0058 | BH | 2 |
| EB0059 | BH | 2 |
| EB0060 | BH | 2 |
| EB0061 | BH | 2 |
| EB0062 | BH | 2 |
| EB0063 | BH | 2 |
| EB0037 | T | 2 |
| EB0086 | T | 2 |
|  |  |  |


| anid | Group | IDX3 |
| :--- | :--- | ---: |
| EB0039 | DAH | 2 |
| EB0040 | DAH | 2 |
| EB0041 | DAH | 2 |
| EB0042 | DAH | 2 |
| EB0043 | DAH | 2 |
| EB0044 | DAH | 2 |
| EB0045 | DAH | 2 |
| EB0046 | DAH | 2 |
| EB0047 | DAH | 2 |
| EB0048 | DAH | 2 |
| EB0049 | DAH | 2 |
| EB0050 | DAH | 2 |
| EB0064 | QA | 2 |
| EB0065 | QA | 2 |
| EB0066 | QA | 2 |
| EB0067 | QA | 2 |
| EB0068 | QA | 2 |
| EB0069 | QA | 2 |
| EB0070 | QA | 2 |
| EB0071 | QA | 2 |
| EB0072 | QA | 2 |
| EB0073 | QA | 2 |
| EB0074 | QA | 2 |
| EB0075 | QA | 2 |
| EB0076 | QA | 2 |
| EB0087 | QA | 2 |
| EB0088 | QA | 2 |
| EB0089 | QA | 2 |
| EB0090 | QA | 2 |
| EB0091 | QA | 2 |
| EB0092 | QA | 2 |
| EB0093 | QA | 2 |
| EB0094 | QA | 2 |
| EB0095 | QA | 2 |
| EB0096 | QA | 2 |
| EB0097 | QA | 2 |
| EB0098 | QA | 2 |
| EB0099 | QA | 2 |
|  |  |  |

From the clustering results, it appears as though the samples from these five locations are actually from two different limestone formations. This is supported by the geographical locations of the quarries. $\mathrm{BH}, \mathrm{DAH}$, and QA are quarries that are very close in proximity and tend to group together in the data set. The same goes for quarries M and T . These quarries are located farther north on the Nile than the other three.

Principal component analysis was also performed on the data set. The concentrations of the elements in each sample were standardized by dividing by the standard deviation. Then Matlab ${ }^{\circledR}$ was used to calculate the loadings, scores, and eigenvalues. A plot of the loadings for the first three principal components (Figure 3.3) shows that almost every element plays a role in the first PC, indicating that there is a lot of variance among the samples. PC2 provides more useful information. PC2 is more strongly driven by fewer elements which could prove useful in distinguishing between sources.

A plot of the scores for PC1 and PC2 (Figure 3.4) shows similar results to the cluster analysis. The data seems to be forming two groups, though those groups are very spread out. The more tightly-bound group is made up of the samples from BH, DAH, and QA, while the samples from M and T are much more spread out. In comparing these two groups, it seems as though there is much less intra-source variation in the samples from $\mathrm{BH}, \mathrm{DAH}$, and QA than there is in the samples from M and T .

Figure 3.3a: Loadings bar plot for Egyptian limestone samples (BNL samples)


Figure 3.3b: Loadings bar plot for Egyptian limestone samples (BNL samples)


Figure 3.4: Plot of PC2 vs. PC1 for the BNL Egyptian limestone samples


Since the groups in the scores plot were so spread out, bivariate plots were tried in order to refine the grouping of the samples. In examining the loadings, several elements were chosen for this analysis, including manganese, iron, and strontium. Figure 3.5 is a plot of the standardized concentrations of iron versus those of strontium. Figure 3.6 is a similar plot using the standardized concentrations of manganese versus strontium. In both plots, the two groups of samples are easily seen, with one group being more spread out in each plot than the other. In Figure 3.5, quarries BH, DAH, and QA form a very tightly-packed group. In Figure 3.6, quarries M and T are the more tightly-packed group, but still show significant variance within the group.

Figure 3.5: Plot of strontium versus iron using the standardized concentrations (BNL samples)


Figure 3.6: Plot of strontium versus manganese using the standardized concentrations (BNL samples)


Standardized [Mn]

The elements chosen here are more mobile in the environment, so the variation seen between the two groups is subject to the extent of weathering; however, substitution of rare earth elements which have similar loadings values for the elements used in these plots shows very similar results.

The results from this set of limestone samples are promising for using elemental composition as a complementary method to sourcing Egyptian limestone. Though the intra-source variance for some sources may be large, like that of the samples from quarries M and T , it still is possible to distinguish between source groups. The large variation contained in source groups for limestone is often seen in other limestone data sets from Europe. ${ }^{3,11}$ The large variation within a source is often attributed to the weathering of the limestone over time, which can leach more mobile elements from the rock. However, a combination of cluster analysis, PCA, and simple bivariate plots allows the extraction of the most information possible from the samples and aids in accurately interpreting the results.

### 3.3.2 Samples from Dr. Harrell

As the samples from BNL's collection seemed to be representative of only two sources of limestone, additional samples from other locations were acquired from Dr. Harrell of the University of Toledo. Dr. Harrell had previously reported the usefulness of petrological analysis for sourcing Egyptian limestone, but also examined the use of elemental analysis as an aid in characterizing the source more thoroughly. It was hoped that using the elemental composition might enable distinction between limestone samples with similar petrology. To this end, 62 samples from 42 locations were acquired from Dr.

Harrell and analyzed by INAA at MURR. The source locations are shown in Figure 3.8 below (listed in Appendix 3).

As for the BNL samples described above, the results from the chemical analysis of these samples, listed in Appendix 4, were subjected to several multivariate techniques, including clustering techniques and principal component analysis. Hierarchical cluster analysis was first performed several times, using both Euclidean and standardized Euclidean distances and using both complete and Ward's linkage methods. Figure 3.8 is a dendrogram resulting from cluster analysis using the standardized Euclidean distance and Ward's linkage method.

Figure 3.7: Map showing the sampled locations in Egypt for the limestone from Dr. Harrell


Figure 3.8 Results from cluster analysis using Euclidean distances and Ward's linkages for the Harrell limestone samples


While there are some small groups that are seen in the dendrogram that do accurately represent samples from about the same location, such as the group of samples 1 and 2, 12 and $13,22-25$, and 61 and 62 , few of the rest of the clusters correlate with geographic location. Since cluster analysis does not always accurately predict groups of similar samples, principal component analysis was performed on the data set. ( $\mathrm{k}_{\text {means }}$ clustering was also performed, but the results also did not show much promise in distinguishing groups of samples.)

Before PCA was performed, the data was standardized as described in Chapter 2 and then PCA was performed using Matlab ${ }^{\circledR}$ as previously described. In examining the eigenvalues resulting from the analysis, the first three principal components described about $85 \%$ of the variance in the data set. The loadings for the first three principal components were then plotted (Figure 3.9).

Figure 3.9: Loadings bar plot for Egyptian limestone samples (Harrell)


Element

In the first PC, the majority of the elements is correlated and appears to have the same contribution to the variance and only calcium is inversely correlated. As the concentration of minor and trace elements increases, this reflects an increase of impurities in the limestone, the amount of pure calcium carbonate decreases. More useful information can be extracted from the second and third principal components.

Here, elements such as strontium (which has been previously shown useful in the elemental analysis of limestone), ${ }^{6}$ hafnium, iron, and manganese play important roles. A biplot of the loadings projected on the scores was created to see how the variation caused by these elements affected the differentiation between source groups. Figure 3.10 depicts PC2 plotted against PC1, while Figure 3.11 shows PC3 versus PC2.

Figure 3.10 Biplot of PC2 vs. PC1 for the Harrell Egyptian limestone samples


Figure 3.11 Biplot of PC3 vs. PC2 for the Harrell Egyptian limestone samples


Figures 3.9 and 3.10 show that only a few samples from some of the sampled locations seem to separate out, while the rest form an oblong cluster in the center. The samples that do separate seem to reflect the geographical locations. For example, Figure 3.12 shows a plot of just the scores for the data, PC3 versus PC2. Figure 3.13 shows the groups of samples highlighted in Figure 3.12 plotted on a map.

Figure 3.12 PC2 versus PC3 for the Egyptian limestone samples (Harrell)


Figure 3.13 Map of three chemical groups of Egyptian limestone samples (Harrell)


In addition to examining the scores plots, several bivariate plots were created using the elements of interest from the loadings matrix, such as strontium, iron, manganese, hafnium, and cobalt. However, the bivariate plots only echoed the results of the scores plots above. The same samples could be separated out to form groups, while the rest of the data set formed one large group with a few outliers. From these results, it seems like the intra-source variation is large enough in many cases to overshadow the inter-source variation.

Though these results are not as promising as those achieved from the BNL data set, there is one major difference between the two that could be a part of the problem. In performing studies of this sort, it is assumed that the composition of a source is relatively homogenous and that the within-source variation is small. However, with the small samples from Dr. Harrell that were used in this analysis, it is possible that the withinsource variation was not characterized well enough to prove this assumption true. With only one or two samples from most of the sources, it is difficult to distinguish the withinsource variation from the variation between sources. In the BNL data set, there were at least ten samples from each quarry site, which allowed the intra-source variation to be characterized more completely. The more samples from each site that can be characterized, the more accurate statistical analysis will be. ${ }^{6,9-11}$ Since limestone is a material that can undergo many changes in composition both during its formation and the natural weathering process, the stone itself can have quite high variations in composition, especially for the more mobile elements like magnesium, sodium, potassium, and iron. ${ }^{6}$ More samples from each source would need to be analyzed in order to better examine the potential of INAA for the characterization of Egyptian limestone sources. However,
based on the results using the BNL data set, it is possible that this could be a powerful technique, especially when used in conjunction with petrological studies.

### 3.4 Conclusions and future work

Using elemental analysis alone to source Egyptian limestone is challenging. The most important of these is the potential for high intra-source variation, which requires the analysis of many samples from each source location. Limestone has been shown to follow the provenance postulate in other parts of the world, but it has not been proven in Egypt, mostly due to the lack of source analyses. With more samples from each known source being analyzed, the intra-source variation can be examined and compared to the inter-source variation. With this variation taken into account, more useful information can be extracted from the samples. As seen with the BNL samples, where the 75 samples were representative of only two source groups, it is possible to distinguish between source groups if the source is well characterized. This information can then be used alongside traditional petrology techniques to provide accurate sourcing of limestone artifacts.

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## CHAPTER 4:

## Sourcing Obsidian from Kenya using Neutron Activation AnAlysis

### 4.1 Introduction

Obsidian, a volcanic stone that is easily chemically characterized and the source of which can be determined with great accuracy, is relatively common in East African artifact assemblages. Although compositional studies of obsidian from East Africa conducted during the 1980s showed great potential, a comprehensive database has not been developed. Instead, many smaller studies have been performed focusing on specific sites and surrounding sources. Similar databases have proven critical to the interpretation of obsidian compositional data in other areas around the world. African archaeology is vital to an understanding of the origins of modern humans, and such a database would be useful to study the movements and interactions of the peoples who lived in this region by studying the distance from recovery sites to the sources to understand the past procurement and exchange systems.

### 4.1.1 Obsidian: a brief overview

Obsidian is a volcanic glass formed when viscous molten lava is cooled too quickly to crystallize. It has many characteristics that make it easy to source. ${ }^{1}$ First, the
sources are limited. Second, the chemical composition of these sources is relatively homogenous within a source, but quite variable between sources. Third, the composition of an artifact made from obsidian is unchanged from that of the source, since the artifact is made simply by flaking the obsidian to create the sharp cutting edges desired. Obsidian does absorb water slowly over time, so hydration layers can form on the outside of the piece, but this crust on the outside is usually avoided for analysis.

### 4.1.1.1 Formation of obsidian

Not every volcano can produce obsidian-forming lava and not every eruption occurs under the right conditions to form obsidian. The lava must have certain characteristics in order to form obsidian. ${ }^{2}$ Specifically, it must be viscous enough to slow the mobility of ions within the lava. Since obsidian forms under unique conditions, the presence of obsidian is limited to those volcanoes, which produce lava that has the opportunity to be quickly cooled, often near a source of water such as a river or lake.

### 4.1.1.2 Composition of obsidian

The chemical composition of obsidian is relatively homogenous within a source, but quite variable between sources. Because of this, obsidian has been found to follow the provenance postulate, which states that in order to be able to determine the source of a sample, the intra-source variations must be statistically less than the inter-source variations. ${ }^{3}$ In some cases, the composition of a particular magma can even be variable over time, making it possible, not only to distinguish between obsidian from different volcanoes, but also to distinguish between obsidian produced from different eruptions of
the same volcano. ${ }^{4}$ Typically, obsidian is primarily silicon dioxide $\left(\mathrm{SiO}_{2} ; 70-75 \%\right)$, but also includes significant amounts of aluminum oxide $\left(\mathrm{Al}_{2} \mathrm{O}_{3} ; 10-15 \%\right)$, sodium oxide $\left(\mathrm{Na}_{2} \mathrm{O} ; 2-5 \%\right)$, potassium oxide $\left(\mathrm{K}_{2} \mathrm{O} ; 1-5 \%\right)$ and iron oxide $\left(\mathrm{FeO}\right.$ and $\left.\mathrm{Fe}_{2} \mathrm{O}_{3} ; 1-5 \%\right) .{ }^{1}$ These concentrations are representative of the typical rhyolitic obsidian. Peralkaline obsidian, which has higher iron concentrations, is also found. Other elements are usually below $1 \% \mathrm{w} / \mathrm{w}$ in concentration. Obsidian also contains a small amount of water, usually $0.1-0.5 \%$. Over time, the concentration of water increases to approximately $3.5 \%$, as the obsidian slowly becomes perlite. ${ }^{4}$

The concentrations of trace elements are used to separate source samples into groups. The best elements are those incompatible with the solid phase due to large ion size or high ionic charges. ${ }^{4}$ The large-ion lithophile elements (LILE) include potassium, rubidium, cesium, strontium, barium, the light rare earth elements (LREEs, La to Sm), thorium, and uranium. The high field-strength elements (HFSE) include yttrium, zirconium, hafnium, niobium, tantalum, and the heavy rare earth elements (HREEs, Eu to $\mathrm{Lu}) .{ }^{1}$ Because lava must cool quickly in order to form obsidian, the incompatible elements are unable to fractionate between solid, crystallizing phases and the rest of the melt. ${ }^{2}$ This gives the obsidian formed a relatively homogenous composition relative to the trace elements. In this way, the incompatible elements tend to have higher and more variable concentrations in obsidian than in other volcanic rocks, such as granite. Many of these elements can be measured with excellent precision and accuracy using either XRF or NAA.

### 4.1.2 African obsidian and anthropological implications

Obsidian use has a long history in East Africa, which is generally considered the region in which modern human behavior originated. The transition to modern behavior can be observed in the patterns of raw material procurement seen in the artifacts left behind. Typically, procurement of raw materials follows a distance-decay pattern, i.e. as the distance between a site and the source increases, the relative abundance of that source decreases. ${ }^{5}$ However, as the interaction between socio-political groups increases, the slope of the typical distance-decay curve becomes shallower, as more artifacts from nonlocal materials are found at greater distances. In some cases, the distance-decay curve may even be the opposite of what is expected, with some of the more-distant sources having greater abundances, though this is more often seen in later sites. ${ }^{5-6}$

Very few studies have been done on obsidian sources in this region of the world, and most have stemmed from the work of Merrick and Brown in 1984. ${ }^{7-8}$ Their work using electron microprobe in addition to x-ray fluorescence spectrometry provides a good start toward the development of a database for Kenya and Tanzania. As a result, they were able to identify at least 35 chemically distinct sources from 54 geographically distinct source localities, based on the more well-known locations of obsidian-bearing localities. In addition, the analyses of artifacts indicated the presence of possibly 22 more unknown sources. Merrick and Brown noticed that they were able to distinguish most groups using major and minor elements (iron, titanium, and calcium), unlike obsidian from other parts of the world which rely on trace elements to characterize sources. They do state that trace element analysis would still be needed to distinguish between chemically similar sources or outcrops of the same source, depending on the precision of
source assignment needed. Their work proved that provenance studies in Kenya and Tanzania had a great potential for studying raw material procurement patterns and interaction patterns of ancient peoples living in this region.

Negash, Shackley, and coworkers have also studied obsidian in Africa, focusing on Ethiopian sites and sources. ${ }^{9-11}$ Their work stems from the preliminary studies done by Muir and Hivernel ${ }^{12}$ in 1976 which showed the potential for sourcing obsidian from the Melka-Konture site in Ethiopia, a site occupied by ancient peoples from the Early Stone Age (ESA) through the Late Stone Age (LSA). Negash et al. have taken this work and expanded it by more extensively sampling known sources of obsidian in Ethiopia and by examining more archaeological sites, including Porc Epic ${ }^{10-11}$, Melka Konture ${ }^{10-11}$, and Beseka ${ }^{9}$, which range from the ESA through to the Holocene ( 1.6 million years ago to 3500 years ago). Similar to the Merrick and Brown studies, they found that the obsidians from sources around Beseka were well-characterized using major and minor elements (iron, titanium, aluminum, and manganese). ${ }^{9}$ Obsidian sources near the other sites, though still using iron as one of the best elements for characterization, used trace elements such as rubidium, zirconium, and zinc to distinguish sources. However, the presence of artifacts from these sites with no known source gives compelling evidence that further study into sourcing obsidian from this region is needed.

Currently, more and more anthropologists are gaining interest in building a database of obsidian sources for this part of the world. More recent work is being done by Ambrose ${ }^{13-14}$, who is expanding on the work of Merrick and Brown in Kenya. His work is currently focused on the Central Rift Valley area of Kenya, which is well-known for its obsidian sources. The Rift Valley sources were commonly used sources because
of their high quality and abundance. ${ }^{14}$ Other researchers are beginning to expand into other parts of Ethiopia and Tanzania and are interested in examining long-distance procurement patterns, which would require consulting a comprehensive database to source artifacts coming from sources quite distant from the site.

Though initial steps have been taken toward studying the geochemistry of the obsidian from this region, the work has mostly been focused on specific archaeological sites or has only brushed the surface of studying all the possible obsidian sources in East Africa. Though many of the known sources are located within the Rift Valley, even these large sources show the possibility of being further divided into smaller subsources, all distinguishable from the chemical data. In addition, one of the major difficulties anthropologists have faced is the large number of artifacts coming from as yet unidentified sources. ${ }^{13,15}$ The presence of unknown sources limits the ability to study procurement patterns and thus limits evaluation of how ancient peoples interacted with their environment and with each other. This can be quite a difficult problem to solve because it is possible that the sources no longer exist or no longer exist in a similar state as when the ancient people were using the source. Because of this, more work needs to be done in this area, first in expanding our geochemical knowledge of currently known sources and subsources, then by extensive fieldwork in collaboration with anthropologists to attempt to discover previously unknown sources, and, lastly, in compiling the work done in previous studies to make it accessible to archaeologists and anthropologists working in this region today.

### 4.1.3 Current project

In this study, geological samples from Kenya were collected and examined via neutron activation analysis to test our ability to distinguish between individual sources and to begin creating a database for future use. The preliminary results indicate that there is a clear correlation between geographic proximity and chemical composition. Most of the source groups can be distinguished by using simple bivariate plots, but a few groups, including possible subgroups, require three-dimensional plots using three elements to show clear separation. Thus far, the results have shown the potential for characterizing Kenyan obsidian based on the major and minor elements rather than trace elements, which has been demonstrated in previous studies.

### 4.2 Experimental methods

Samples of obsidian from the Central Rift Valley in Kenya were acquired over a period of several field seasons. Initially, one hundred samples from sources mostly centering around Lake Naivasha were obtained from Dr. Stanley Ambrose at the University of Illinois, which he had collected over several previous field seasons. An additional 405 samples were collected during eight-week field seasons in Kenya by Dr. Ambrose and two of his students in summer 2008 and 2009. I participated in collecting field samples in the 2008 field session. A total of 505 samples from various parts of the Central Rift Valley have been analyzed. A map showing the locations of the sources of these samples is shown in Figure 4.1.

Figure 4.1: Map of sampled locations of obsidian in the Central Rift Valley of Kenya


Neutron activation analysis was performed at the University of Missouri Research Reactor under the guidance of Dr. Michael Glascock. For the analysis of the source samples, previously developed procedures were used for sample preparation and analysis for both short and long irradiations. The relative concentrations of twenty-eight elements (Al, Ba, Ce, Co, Cs, Cl, Dy, Eu, Fe, Hf, K, La, Lu, Mn, Nd, Na, Rb, Sb, Sc, Sr, Sm, Ta, $\mathrm{Tb}, \mathrm{Th}, \mathrm{U}, \mathrm{Yb}, \mathrm{Zn}$, and Zr ) were determined by INAA at MURR. Concentrations of each element are determined by the comparator method. The standards used included NIST standards SRM-278 (obsidian) and SRM-1633a (fly ash). An additional standard, JJO, an obsidian source sample from the Alca source in Peru, was used as a quality control sample.

For short irradiation samples, samples of between 50 and 100 milligrams were weighed out into polyethylene vials. The vials were then subjected to a five-second irradiation, followed by a 25 -minute decay and a 12 -minute count. The elements determined from the short irradiations are aluminum, barium, calcium, dysprosium, potassium, manganese, sodium, titanium and vanadium.

Samples of between 100 and 200 milligrams were weighed out into quartz vials and sealed for the long irradiation and were placed into the reactor for a 70-hour irradiation. These samples are counted twice, once after a seven-day decay followed by a 33 -minute count and again after a 28 -day decay and a 168-minute count. The first count determines elements that have a mid-range half-life (including lanthanum, lutetium, neodymium, samarium, uranium and ytterbium) and the second count determines the elements that produce long-lived isotopes (cerium, cobalt, chromium, cesium, europium, iron, hafnium, nickel, rubidium, antimony, scandium, strontium, tantalum, terbium,
thorium, zinc and zirconium). All concentration data for the samples can be found in the Appendix 6.

Following data acquisition, the concentration values were standardized as described in Chapter 2. Then principal component analysis was applied to the standardized data set, using the Matlab© program. Results of the chemical analysis were then compared with the geographical locations for the characterization of geographical source groups.

### 4.3 PCA results

As described in earlier chapters, sourcing studies rely on being able to characterize the variance between source groups. In this study, data for over 30 elements was acquired for 505 samples and PCA was applied to this data set in order to extract the most information possible without having to go through each element and each sample individually. As discussed in Chapter 2, PCA can be used to determine how the variance is driven by each variable and how this can be used to distinguish between groups in the data set.

The eigenvalues vector was first examined to see which principal components described the most variance in the data set, which would focus the rest of the interpretation. As seen in Figure 4.2 below, the first three principal components describe almost $90 \%$ of the variance in the data. The Scree plot also starts to level off after the second PC. Hence, we concluded that the first three PCs can be used for examining the variance in the data set.

Figure 4.2 Scree plot from PCA results of the standardized Kenyan obsidian data


Next, a bar plot of the loadings was created to determine the influence of each element on the first three principal components (Figure 4.3). From this plot it can be seen that there is significant covariance among the rare earth elements, which drives the variance explained by PC 1 , while PC 2 is more strongly influenced by many of the minor elements, such as iron, sodium, and manganese, as well as a few trace elements. As discussed earlier, Merrick and Brown had found with their analyses that the minor elements could be used to distinguish source groups and these results are consistent with their findings. Bivariate plots of elements that are not correlated but have a strong influence on either of the first two PCs might be enough to distinguish between source groups, since these elements drive the variance in the data set. For example, plotting iron
versus cesium has been found useful in separating most of the major groups in the data set (discussed in detail later). PC3 is influenced mostly by two of the major elements, aluminum and potassium, and to a smaller extent scandium.

Figure 4.3 Bar plot of loadings from PCA results of the standardized Kenyan obsidian data


Since PC1 and PC2 seem to describe the majority of the variance in the data set, the scores for PC1 and PC2 were plotted to see if groups were discernible (Figure 4.4). Several large groups were quite noticeable in this evaluation. Other smaller groups were then distinguished on closer inspection of the plot. These groups represent samples that are similar to each other in multidimensional space and so have similar compositional patterns, which ideally are linked to geographical location. In addition, a few outliers were identified.

Figure 4.4 Scores plot from PCA results of the standardized Kenyan obsidian data


In order to transition to examining simple bivariate plots, which are most often used with artifacts to simplify the analysis, the loadings for PC1 and PC2 were projected onto the scores plot to determine which pairs of elements might be most useful in distinguishing source groups (Figure 4.5). This plot reiterates how many of the rare earth elements covary with each other and have a strong influence on PC1. It also shows more clearly the relationship between pairs of elements. Pairs of elements that have the potential to aid in distinguishing source groups are usually those that have little correlation between them. So, a plot of iron versus cesium (shown in the following section) would do well in distinguishing groups while a plot of iron versus manganese would not. Other pairs of elements that were considered were dysprosium versus iron,
thorium versus iron, cesium versus scandium, sodium versus scandium, and europium versus thorium.

Figure 4.5 Biplot of scores and loading vectors from PCA results of the standardized Kenyan obsidian data


### 4.4 Characterizing geological source groups using elemental bivariate plots

From the results of the principal component analysis, an initial set of chemical source groups were created based on the samples that were similar to each other in multidimensional space. These initial groups, shown in the scores plot, Figure 4.6, were used as a starting point to characterize the geological sources of obsidian. The locations of these groups were then plotted on a map to examine the results when compared with the different geographical locations of the sources (Figure 4.7). As can be seen in Figure
4.7, several of the geological sources seem to have very similar compositions, even though they are in very different locations, while other sources that exist in close proximity are quite different in composition.

Figure 4.6: Initial groups, using scores plot from PCA results of the standardized Kenyan obsidian data


Figure 4.7 Map showing locations in Kenya of the initial chemical groups


For finer separations of the geological source groups, several bivariate plots were used to distinguish between sources. In some cases, the difference in the composition of two source groups, even if they are quite distant from each other, can be only one or two elements. Starting with the groups found in the initial scores plot, the source groups were refined with each bivariate plot and the results were compared to the geographical locations of the sources. As soon as a group could be reliably distinguished from the rest of the samples, it was removed from the subsequent plots for clarity. In this way, some groups can be characterized using only one plot, while other groups take several bivariate plots to accurately characterize.

Elements for each bivariate plot were chosen based on the results of the PCA performed on the entire data set. The two-dimensional biplot shown above (Figure 4.5) was used to choose the best elements for the separation of each source group. Elements such as iron, cesium, thorium, and many of the rare earth elements have the potential to distinguish between groups, since they contribute significantly to the variance in the data set. The use of iron to distinguish between source groups was also supported by the literature, as many researchers working in this region found iron and other minor elements useful for characterization. ${ }^{7-11,15}$

The following bivariate plots demonstrate the characterization of the groups using this method of analysis. The first of these is a plot of the standardized values for cesium versus those for iron (Figure 4.8). As seen in the plot, many of the original groups found in the plot of PC2 versus PC1 (shown in the same colors as above) are still visible and distinguishable. In fact, this plot can distinguish between many of the geographical source locations and finer separations can be made between source groups that previously
seemed very similar. For example, group 6 appeared in the scores plot to be one larger group. In Figure 4.8, however, three small groups are distinguishable. These finer separations were plotted on the map and confirmed the separation between source groups. However, there are still groups that are geographically distinct and cannot be distinguished using this elemental bivariate plot.

Figure 4.8: Plot of cesium versus iron using the standardized concentrations for the entire Kenyan obsidian data set


In order to better separate the geographical source groups contained in groups 11 and 14, additional bivariate plots are needed.

Groups 11 and 14, shown in Figure 4.8, overlap quite a bit, even though they are completely separated in the scores plot. Another bivariate was made using the
standardized concentrations of dysprosium versus cesium (Figure 4.9). In this plot, groups 11 and 14 are well separated and subgroups within each group are more easily visible. In group 11 , there appear to be three subgroups $(11,18,19)$. When plotted on a map, these three chemical subgroups appear as distinct geological sources (Figure 4.10). Subgroup 19 does seem to characterize two different geological sources, but it is possible that these two sources were formed during the same geological event. The samples would then have very similar compositions, as seen here, even though the sources are several kilometers apart.

Figure 4.9: Plot of dysprosium versus cesium, using standardized concentrations, showing the separation of group 11 into subgroups


Figure 4.10: Map of the locations of chemical group 11 subgroups, identified on Figure 4.9.


The last group to be characterized is group 14, seen in Figure 4.9 above. This group is spread out in many of the bivariate plots without forming distinct subgroups. Geographically, this group covers a large area and overlaps many of the other sources described above. To further examine this "source group", PCA was performed on this subset of samples which contained 110 samples. While one must exercise caution in employing PCA on small sample sets, this analysis provided information about which elements contribute most to the variance in this small subset. From the results of the loadings plot (Figure 4.11), it seems that antimony and scandium play more of a role in the variance described by the first two principal components for this set of samples than in the entire data set. Other elements that had played a larger role before, such as cesium and dysprosium, play a much smaller role here, which is why good separation was not achieved with the previous bivariate plots. A plot of antimony versus iron shows good separation between subgroups within group 14 (Figure 4.12) and these groups match up relatively well geographically (Figure 4.13).

Figure 4.11: Loadings plot for group 14 of the Kenyan obsidian data set


Figure 4.12: Plot of antimony versus iron, using standardized concentrations, for group 14 of the Kenyan obsidian data set


Figure 4.13: Map of subgroups of group 14 identified in Figure 4.12


Though Figure 4.12 shows good separation between most of the subgroups, there is still one subgroup, 28(14), that seems to represent three geographically distinct source groups. This subgroup can be further divided into three groups using a plot of uranium versus thorium, two additional elements that play a large role in the first two principal components (Figure 4.14). These smaller groups match up well with the geographical locations (Figure 4.15).

Figure 4.14: Plot of uranium versus thorium, using standardized concentrations for group 14, subgroup 28


Figure 4.15: Map of subgroup 28 for group 14, identified in Figure 4.14


### 4.5 Conclusion

Through this study, it has been demonstrated that obsidian from the Central Rift Valley in Kenya follows the provenance postulate. Over 25 chemically-distinct sources were characterized from over 40 geographically-distinct sources. Many of the major source groups can be distinguished using one elemental bivariate plot, but a few sources, including several subgroups within those sources, require a series of bivariate plots to characterize them completely. As work on obsidian from this region continues, more samples will allow for additional refinement of known source groups as well as provide information on "new" sources. As this data set is continually added to and refined through collaboration with researchers working in this region, the results can be applied to archaeological and anthropological research to answer questions about the origin of modern humans and their interactions with each other and the environment.

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## Chapter 5:

A new Epithermal Neutron Activation Analysis method for TITANIUM AND BARIUM ANALYSIS IN OBSIDIAN

### 5.1 Introduction

Though NAA and XRF have been the most common analytical techniques used in sourcing studies for obsidian, other techniques, such as electron microprobe analysis, have been used. Each of these techniques varies with the elements that can be determined and the sensitivity for those elements. Even though standard thermal NAA is sensitive for many of the elements used for obsidian analyses, there are some elements that require a different analysis procedure in order to achieve accurate and precise results. By examining alternate neutron-induced reactions for the analysis, not only can better results be achieved for currently measured elements, but the list of measurable elements can be expanded to include those that are usually more difficult to quantify.

### 5.1.1 The challenge

Previous work has described the use of such elements as titanium and barium to distinguish between chemical source groups of obsidian in several areas of the world, including Kenya. The two most common techniques used for these studies have been
electron microprobe analysis (EMPA) and x-ray fluorescence (XRF). In many cases, a combination of the two has been used, such as in Merrick and Brown's work. ${ }^{1-2}$

However, a concern arises about the use of these two techniques for the analysis of titanium and barium. Both of these methods are based on the emission of secondary x rays from the sample to characterize the chemical composition. The typical x-ray line used for titanium analyses is the K-line at 4.51 keV . This falls very close to the x-ray line normally used for barium analysis, the most intense L-line at 4.47 keV . A resolution of better than 40 eV is needed in order to clearly separate these peaks. Many XRF and EMPA instruments use energy-dispersive detection systems which have resolutions of several hundred eV, which means that the peaks would be overlapping, as seen in Figure 5.1.

Figure 5.1: Titanium (K) and barium (L) lines in a typical EDXRF spectrum.


However, there are also many instruments for both techniques that use wavelengthdispersive detection systems, where the resolution can be on the order of $5-10 \mathrm{eV}$. This would easily separate the two peaks and greatly simplify the quantification of barium and titanium. Because concentrations of barium in obsidian can range up to several hundred parts per million, the overlap of the peaks becomes an important consideration. If the peaks cannot be resolved, then there would be a contribution of barium to the titanium peak, causing an inflated value for titanium.

In many of the previous studies, however, the type of instrument is not mentioned. ${ }^{1-2}$ In addition, any calculations used to take this interference into account are not described and a list of which elements were determined by which technique is not given. It cannot be assumed that wavelength-dispersive instruments were used, so a way of independently verifying the previous results is needed to ensure the accuracy of the values included in a large database.

Neutron activation analysis (NAA) is a good candidate as a verification method. Since it is based on nuclear reactions (as described in Chapter 1), there is no interference experienced between the titanium and the barium in the analysis. It is matrixindependent and has many tunable parameters that can be utilized to get the best results possible. A method using NAA to determine the titanium and barium concentrations in obsidian would work well to verify the results previously attained, especially in Kenyan obsidian. These results could then also be used in conjunction with the results achieved from the standard NAA procedure to allow us to make direct comparison to previous studies.

Currently there is no method in place at MURR for the analysis of titanium in obsidian and the current method for barium suffers from poor precision over much of the concentration range observed. New methods need to be developed to determine titanium and barium with good accuracy and precision. In this chapter, the new methodology that has been developed will be described. For each element, the limits of detection have been calculated and the precision is evaluated using the relative standard deviation. This new methodology offers better precision and low detection limits for the analysis of barium and titanium and also provides a way to expand the current list of elements that can be determined in obsidian.

### 5.2 Development of the new methodology

To obtain better results for barium and titanium in obsidian than is achievable with the current method, alternate neutron-induced reactions need to be considered, in particular those using epithermal or fast neutrons. These reactions are typically not used because they are usually masked by the activity generated from the thermal neutron reactions. In most cases, the thermal neutron cross sections are larger than the fast and resonance reactions and thermal neutrons dominate the spectrum; they make up about $90 \%$ of the neutron flux in the graphite moderated irradiation positions. The number of these reactions can cause a large background in the gamma spectrum that overwhelms the peaks from isotopes formed via epithermal and fast neutron activation.

However, using a boron or cadmium shield around the samples can reduce the thermal neutron flux reaching the samples and thus reduce the number of thermal neutron reactions taking place. At MURR, a boron shield is often used for short irradiations,
while a cadmium can is used for longer irradiations. Boron and cadmium both have large cross sections for thermal neutrons and much smaller cross sections for the higher-energy neutrons. Cadmium, in particular, is very efficient at absorbing neutrons with energies lower than $0.5 \mathrm{eV} .{ }^{3-4}$ At greater energies, the ability for cadmium to absorb the neutrons drops off precipitously. This makes cadmium a good choice for specifically filtering thermal neutrons while allowing the epithermal and fast neutron fluxes to remain relatively high. For cadmium-covered irradiations, a standard aluminum canister (3.35 in x 10 in ) has been lined with 40 mils ( 0.040 in or 1.016 mm ) of cadmium. ${ }^{5}$

Epithermal and fast neutron activation analysis has been shown to greatly increase the sensitivity for the analysis of many elements in geological materials. ${ }^{3,6-9} \mathrm{By}$ reducing the thermal flux, and thus reducing the Compton due to the many abundant isotopes that react via thermal $(\mathrm{n}, \gamma)$ reactions, the sensitivity for many of the trace elements can be improved usually by factors of 2-5. ${ }^{7}$ Here, methods for the analysis of titanium and barium using epithermal neutron activation analysis (ENAA) will be considered as well as the standard thermal NAA procedures for the best sensitivity.

### 5.2.1 Titanium

For titanium, several reactions can be considered. First, the only (n, $\gamma$ ) reaction to result in a radioactive isotope of titanium was examined. Here, $\mathrm{Ti}-50$, which has a natural abundance of $5.18 \%$, reacts with thermal neutrons to produce $\mathrm{Ti}-51\left(\mathrm{t}_{1 / 2}=5.76\right.$ $\min , \mathrm{E} \gamma=320.1 \mathrm{keV}) .{ }^{10}$ If this reaction was used for the measurement of titanium, it could be incorporated into the current short-irradiation procedure, since the half-life of Ti-51 is relatively short. However, the cross section for Ti-50 for thermal neutrons is
0.1795 barns. ${ }^{10}$ This is relatively small, meaning that not much activity is produced during the irradiation. In fact, in examining the spectrum for an obsidian standard, NIST SRM-278, which has a titanium concentration of 1470 ppm , the peak for Ti-51 is barely visible above background. If the concentration of titanium was smaller than this, as for another standard, JR-1 ([Ti] = 660 ppm$)$, then a peak would not be distinguishable from background. Thus, other reactions and methodologies must be considered.

Titanium can also undergo several different reactions when the isotopes react with fast neutrons. One potential reaction is the ( $\mathrm{n}, \mathrm{p}$ ) reaction on $\mathrm{Ti}-46$ to produce $\mathrm{Sc}-46$. Ti46 has a natural abundance of $8.25 \%$ and a fast reaction cross section of 12.5 millibarns. ${ }^{10}$ Sc-46 has a half-life of 83.79 days, which would make it a good candidate for a longirradiation procedure that is followed by a longer count time. However, $\mathrm{Sc}-46$ is also produced via an $(\mathrm{n}, \gamma)$ reaction from $\mathrm{Sc}-45\left(100 \%\right.$ abundant, $\sigma_{\text {thermal }}=17.4 \mathrm{~b}, \sigma_{\text {epithermal }}=$ $7.0 \mathrm{~b}) .{ }^{10}$ Since there is usually some scandium present in obsidian and other geological samples, there is no way to distinguish the activity from scandium from the activity due to the reaction with titanium.

Another fast neutron reaction is the ( $\mathrm{n}, \mathrm{p}$ ) reaction on $\mathrm{Ti}-47$ to form $\mathrm{Sc}-47$. Ti-47 has a natural abundance of $7.44 \%$ and a fast neutron cross section of $20.0 \mathrm{mb} .{ }^{10} \mathrm{Sc}-47$ has a half-life of 3.35 days, which would work well for a long irradiation followed by a longer count time. The only concern with this reaction is that $\mathrm{Sc}-47$ is also produced via the beta decay of $\mathrm{Ca}-47$, which decays at a similar rate as $\mathrm{Sc}-47\left(\mathrm{t}_{1 / 2}=4.54\right.$ days for $\mathrm{Ca}-$ 47). ${ }^{10} \mathrm{Ca}-47$ is produced via an $(\mathrm{n}, \gamma)$ reaction on $\mathrm{Ca}-46$, which has a $0.004 \%$ abundance and a thermal cross section of 0.74 b . Ca-47 also emits several gamma rays, the most abundant of which is at $1297.09 \mathrm{keV}(71.00 \%)$. However, since calcium concentrations
in obsidian are usually less than $0.5 \%$ and the abundance of $\mathrm{Ca}-46$ is so low, this interference should be minimal. However, peaks in the gamma spectrum due to the decay of Ca-47 can be monitored for their presence or absence.

The last option for a fast-neutron reaction is an ( $\mathrm{n}, \mathrm{p}$ ) reaction on $\mathrm{Ti}-48$ to make Sc-48. Titanium-48 is the most abundant isotope of titanium at $73.72 \%$. The product, Sc-48, has a half-life of 43.67 hours, and has two gamma rays that are $100 \%$ abundant at 983.52 keV and 1312.10 keV . However, the cross section for the $(\mathrm{n}, \mathrm{p})$ reaction with fast neutrons on Ti-48 is only 0.315 millibarns. ${ }^{10}$

Production rates for all three ( $n, p$ ) reactions were calculated to quantitatively compare them. Table 5.1 shows a summary of these calculations. As can be seen in the table, the reaction Ti-47 (n,p) Sc-47 produces the greatest amount of activity, mostly due to the larger cross section. This reaction was chosen as the best to use for the quantification of titanium using this method, but the Ti-48 (n,p) Sc-48 was also monitored in the spectrum.

Table 5.1: Production rates for fast neutron reactions of titanium using 200 mg of SRM 278 ([Ti] = 1470 ppm ), a fast neutron flux of $1.50 \times 10^{12} \mathrm{n} \mathrm{cm}^{-2} \mathrm{~s}^{-1}$, and an irradiation time of 48 hours ${ }^{10}$

|  | Ti-46 (n,p) Sc-46 | Ti-47 (n,p) Sc-47 | Ti-48 (n,p) Sc-48 |
| :--- | :---: | :---: | :---: |
| Target Abundance | $8.25 \%$ | $7.44 \%$ | $73.72 \%$ |
| Target cross section | 12.5 mb | 20 mb | 0.315 mb |
| Half-life | 83.79 days | 3.35 days | 43.67 hours |
| Saturation Activity | $5.74 \times 10^{3} \mathrm{~Bq}$ | $8.28 \times 10^{3} \mathrm{~Bq}$ | $1.29 \times 10^{3} \mathrm{~Bq}$ |
| Activity at end of <br> irradiation | 94.1 Bq | $2.8 \times 10^{3} \mathrm{~Bq}$ | 689 Bq |

### 5.2.2 Barium

Currently, the standard procedure to measure barium in the Archaeometry Group at MURR uses the thermal neutron $\mathrm{Ba}-138(\mathrm{n}, \gamma) \mathrm{Ba}-139$ reaction. $\mathrm{Ba}-138$ is the most abundant isotope of barium (71.69\%) and has a thermal neutron cross section of 0.36 barns. ${ }^{10}$ Ba-139 has a half-life of 83.06 minutes and emits a gamma ray at 165.85 keV (23.70\%). The current procedure determines the concentration of barium as part of the short-irradiation procedure, where the samples are irradiated for 5 seconds, allowed to decay for 25 minutes, and then counted for 12 minutes. However, because Ba-138 has such a low cross section, not much $\mathrm{Ba}-139$ is produced during the short irradiation. In addition, the half-life of $\mathrm{Ba}-139$ is so long compared to the counting time and the gamma ray emitted has a low abundance that very little is detectable above background. This method suffers from poor precision, especially when the concentration of barium in the sample is low. This method could be improved by irradiating the sample for a longer period of time as well as counting it for longer as well. Alternate methods using a Cd covered irradiation were explored.

One reaction that was of great interest was $\mathrm{Ba}-130(\mathrm{n}, \gamma) \mathrm{Ba}-131 . \mathrm{Ba}-130$ is a less abundant isotope of barium ( $0.106 \%$ ) but has a thermal cross section of 8.8 barns and a resonance cross section of 184 . barns. ${ }^{10} \mathrm{Ba}-131$ has a half-life of 11.5 days, making it suitable for a procedure using a longer counting time, and the most abundant gamma ray has an energy of $496.33 \mathrm{keV}(47.00 \%)$. With a cadmium can reducing the thermal flux reaching the samples, the large resonance cross section can be utilized to produce a strong signal for $\mathrm{Ba}-131$. So, even though the abundance of the target isotope is low, this
reaction can produce results at least ten times better than the reaction using Ba-138 as the target.

### 5.3 Initial experiments

The initial experiments to test this new methodology used a cadmium can to reduce the thermal neutron flux and take advantage of the larger cross sections for the epithermal and fast neutron reactions. Several standards and samples were analyzed to explore a wide range of concentrations for titanium and barium. A procedure resembling the standard long-irradiation procedure was used to provide the longer irradiation times and longer counting times necessary. From the results, the accuracy and precision were evaluated for each method and the limits of detection were calculated.

### 5.3.1 Standards and samples

In order to demonstrate the new methods for a wide range of concentrations of barium and titanium, several different standards and samples were prepared. The standards included NIST SRM 278 (obsidian), USGS AGV-1 (andesite), JGS JA-1 (andesite), JGS JB-2 (basalt), JGS JR-1 (rhyolite), and NIST SRM1633a (fly ash). The concentrations of titanium and barium in these standards are found in Table 5.2.

Table 5.2: Concentrations of titanium and barium in NIST, USGS, and JGS standards ${ }^{10}$

| Standard | Titanium (ppm) | Barium (ppm) |
| :---: | :---: | :---: |
| NIST SRM 278 (obsidian) | 1470 | 881. |
| USGS AGV-1 (andesite) | 6340 | 1221 |
| JGS JA-1 (andesite) | 5100 | 311 |
| JGS JB-2 (basalt) | 7100 | 222 |
| JGS JR-1 (rhyolite) | 660 | 50.3 |
| NIST SRM 1633a (fly ash) | 8230 | 1320 |

As well as the standards, obsidian samples from well-characterized sources in Central and South American were also analyzed to examine the ranges of barium and titanium concentrations found in obsidian. These included samples from Guadalupe Victoria, Otumba, Paredon, El Paraiso, and Pico de Orizaba. In addition, several flux wires were placed inside the cadmium can with the samples to measure the thermal and fast neutron fluxes. For the analyses, two hundred milligrams of each sample was weighed out and sealed into a high-purity quartz vial.

### 5.3.2 Irradiation, decay, and counting procedures

In order to achieve the best possible results for these analyses, a modified longirradiation procedure was used. Samples were irradiated in a cadmium can in the reactor for 48 hours. Then, because the final total activity of the samples after irradiation was unknown, the samples were allowed to decay for 7 days. Because the samples were not as radioactive from this cadmium-covered irradiation as they would be after a typical long irradiation, subsequent analyses allowed the samples to decay for only 5 days. Then the samples were counted on an HPGe detector to acquire the gamma spectra. The samples were counted for 30 minutes each, 5 inches away from the detector. Dead times for the detectors were kept below $15 \%$, usually ranging between 5 and $10 \%$.

### 5.3.3 Gamma spectra and determination of precision and limits of detection

The results from the first experiment showed marked improvement over traditional NAA methodologies. The cadmium can reduced the thermal flux by about a factor of 10 , which allowed the peaks emitted from isotopes resulting from epithermal
and fast neutron reactions to be observed. Limits of detection were lowered and precision was greatly improved. Figure 5.2 shows the comparison between the gamma spectrum acquired during a standard NAA procedure and one acquired after irradiation in the cadmium can for NIST SRM 278 (obsidian). As seen in the spectrum, the activity due to thermal neutron reactions was significantly reduced and allowed gamma peaks resulting from epithermal and fast neutron reactions to be detectable.

Figure 5.2: Comparison of the gamma ray spectrum from the standard NAA procedure and the epithermal NAA procedure


### 5.3.3.1 Titanium results

After examining the results from this initial analysis, the Ti-47 (n,p) Sc-47 reaction was chosen for the analysis of titanium in the sample. As expected, the lower cross section of Ti-48 caused less $\mathrm{Sc}-48$ to be produced, so the peak was not visible above the background. A clear peak was, however, seen for Sc -47 in all the samples analyzed, even those with lower titanium concentrations.

One concern, as stated above, is that $\mathrm{Sc}-47$ is also produced from the beta decay of $\mathrm{Ca}-47$. To examine this potential interference, the gamma peak for $\mathrm{Ca}-47$ at 1297.09 keV was followed in the spectrum. However, because of the small abundance and low cross section of the target isotope, $\mathrm{Ca}-46$, very little $\mathrm{Ca}-47$ activity is produced. The gamma-ray peak due to $\mathrm{Ca}-47$ was not observed in any of the standards or samples. However, this energy was monitored for all samples. If a peak had appeared at 1297 keV , the contribution of the $\mathrm{Ca}-47$ decay to the $\mathrm{Sc}-47$ peak would be taken into account in the calculations.

Figure 5.3 shows two gamma spectra of NIST SRM 278 (obsidian), one resulting from a typical short irradiation analysis and the other from the cadmium-covered, long irradiation. These spectra highlight the contrast between the results for titanium for each irradiation. In the typical short irradiation analysis, the peak for $\mathrm{Ti}-51$ is barely visible above background. Since this standard has one of the higher concentrations of titanium (1470 ppm), it is unlikely that a peak would be detectable above background for samples with lower concentrations. On the other hand, the spectrum resulting from the cadmium-
covered long irradiation shows a distinct peak from $\mathrm{Sc}-47$, which can be used to calculate the concentration of titanium in the sample.

Figure 5.3: Gamma spectra of SRM 278 (obsidian) depicting peaks for titanium analysis


Since each standard was a different matrix, the estimated limit of detection for titanium, percent relative standard deviation, and the signal to noise ratio were calculated for each, for both the typical short irradiation and the cadmium-covered long irradiation analysis. The limit of detection was calculated by estimating the background under the peak and then calculating three times the square root of the background counts. The results of these calculations are shown in Table 5.3. For all matrices, the cadmiumcovered irradiation procedure showed a marked improvement in the limit of detection of
titanium and the precision of the measurements. The accuracy of the results also improved significantly.

Table 5.3: Titanium analysis results: comparison of thermal NAA and fast NAA (Concentrations were calculated using NIST SRM 278 as a primary standard)

|  | Literature Concentrations ${ }^{10}$ | $\begin{gathered} \mathrm{Ti}-51 \\ \text { (thermal NAA) } \end{gathered}$ | $\begin{gathered} \text { Sc-47 (Ti-47) } \\ \text { (fast NAA) } \\ \hline \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| SRM 278 <br> (primary standard) Signal to Noise Ratio LOD | $1470 \mathrm{ppm} \pm 40$ | $\begin{gathered} \text { N/A } \\ 0.062 \\ 436 \mathrm{ppm} \end{gathered}$ | $\begin{gathered} \mathrm{N} / \mathrm{A} \\ 0.47 \\ 58.1 \mathrm{ppm} \end{gathered}$ |
| SRM 1633a (n=3) <br> $[\mathrm{Ti}] \pm$ Std. Dev <br> \% RSD <br> Signal to Noise Ratio <br> LOD | $8230 \mathrm{ppm} \pm 390$ | $\begin{gathered} 11530 \pm 1430 \\ 12.44 \% \\ 0.97 \\ 151 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 7973 \pm 23 \\ 0.28 \% \\ 0.78 \\ 124 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{r} \hline \text { AGV-1 }(\mathrm{n}=3) \\ {[\mathrm{Ti}] \pm \text { Std. Dev }} \\ \% \text { RSD } \end{array}$ <br> Signal to Noise Ratio LOD | $6340 \mathrm{ppm} \pm 300$ | $\begin{gathered} 5751 \mathrm{ppm} \pm 257 \\ 4.47 \% \\ 0.16 \\ 314 \mathrm{ppm} \end{gathered}$ | $\begin{gathered} 6461 \mathrm{ppm} \pm 94 \\ 1.45 \% \\ 2.01 \\ 53.5 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{r} \hline \text { JA-1 }(\mathrm{n}=3) \\ {[\mathrm{Ti}] \pm \text { Std. Dev }} \\ \text { \% RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \\ \hline \end{array}$ | 5100 ppm | $\begin{gathered} \text { 4440. } \mathrm{ppm} \pm 197 \\ 4.44 \% \\ 0.11 \\ 365 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 5239 \mathrm{ppm} \pm 71 \\ 1.36 \% \\ 2.59 \\ 34.4 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{\|r\|} \hline \text { JB-2 }(\mathrm{n}=3) \\ {[\mathrm{Ti}] \pm \text { Std. Dev }} \\ \text { \% RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \\ \hline \end{array}$ | 7100 ppm | $\begin{gathered} 4807 \mathrm{ppm} \pm 604 \\ 12.55 \% \\ 0.12 \\ 505 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 7482 \mathrm{ppm} \pm 118 \\ 1.58 \% \\ 2.14 \\ 40.9 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{aligned} & \text { JR-1 }(\mathrm{n}=3) \\ & {[\mathrm{Ti}] \pm \text { Std. Dev }} \\ & \% \text { RSD } \end{aligned}$ <br> Signal to Noise Ratio <br> LOD | 660 ppm | $\begin{gathered} 486.2 \mathrm{ppm} \pm 430.2 \\ 88.48 \% \\ 0.024 \\ 462 \mathrm{ppm} \end{gathered}$ | $\begin{gathered} 678.2 \mathrm{ppm} \pm 56.7 \\ 8.36 \% \\ 0.12 \\ 69.1 \mathrm{ppm} \\ \hline \end{gathered}$ |

### 5.3.3.2 Barium results

Using the $\mathrm{Ba}-130(\mathrm{n}, \gamma) \mathrm{Ba}-131$ reaction for the analysis of barium results in great improvement in the precision of the analysis, especially at the low concentrations. Figure 5.4 shows the two comparison spectra for NIST SRM 278, with the peaks due to $\mathrm{Ba}-139$ and $\mathrm{Ba}-131$ highlighted. Whereas the peak due to $\mathrm{Ba}-139$ is hardly visible above background in the traditional irradiation, the peak from $\mathrm{Ba}-131$ ( 496.3 keV ) is strong.

Figure 5.4: Gamma spectra of SRM 278 (obsidian) depicting peaks for barium analysis


As for the titanium analysis, the limit of detection, precision, and signal-to-noise ratio were calculated for each standard for both the standard irradiation procedure and the cadmium-covered irradiation. The results of these calculations can be found in Table 5.4. The limit of detection for barium improves by a factor of ten, while the precision (\%RSD) improves by a factor of 100 . The accuracy of the measurement also improves greatly.

Table 5.4: Barium analysis results: comparison of thermal NAA and epithermal NAA
(Concentrations were calculated using NIST SRM 278 as a primary standard)

|  | Literature concentrations ${ }^{10}$ | $\begin{gathered} \text { Ba-139 } \\ \text { (thermal NAA) } \end{gathered}$ | $\begin{gathered} \mathrm{Ba}-131 \\ \text { (epithermal NAA) } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| SRM 278 <br> (primary standard) Signal to Noise Ratio LOD | $880 \pm 40$ | $\begin{gathered} \mathrm{N} / \mathrm{A} \\ 0.226 \\ 26.2 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{N} / \mathrm{A} \\ 5.07 \\ 4.26 \mathrm{ppm} \\ \hline \end{gathered}$ |
| SRM 1633a (n=3) <br> $[\mathrm{Ba}] \pm$ Std. Dev <br> \% RSD <br> Signal to Noise Ratio <br> LOD | $1320 \pm 40$ | $\begin{gathered} 1509 \mathrm{ppm} \pm 144 \\ 9.57 \% \\ 0.79 \\ 161 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 1279 \pm 19 \\ 1.49 \% \\ 2.68 \\ 124 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{aligned} & \text { AGV-1 }(\mathrm{n}=3) \\ & {[\mathrm{Ba}] \pm \text { Std. Dev }} \\ & \% \text { RSD } \end{aligned}$ <br> Signal to Noise Ratio <br> LOD | $1221 \pm 16$ | $\begin{gathered} 858 \mathrm{ppm} \pm 31 \\ 3.65 \% \\ 0.21 \\ 35.4 \mathrm{ppm} \end{gathered}$ | $\begin{gathered} 1194 \mathrm{ppm} \pm 8 \\ 0.65 \% \\ 5.26 \\ 4.70 \mathrm{ppm} \end{gathered}$ |
| $\begin{array}{\|r\|} \hline \text { JA-1 }(\mathrm{n}=3) \\ {[\mathrm{Ba}] \pm \text { Std. Dev }} \\ \text { \% RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \\ \hline \end{array}$ | 311 | $\begin{gathered} 161.1 \mathrm{ppm} \pm 21.7 \\ 13.46 \% \\ 0.049 \\ 54.2 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 293.6 \mathrm{ppm} \pm 2.4 \\ 0.80 \% \\ 1.68 \\ 4.74 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{\|r\|} \hline \text { JB-2 }(\mathrm{n}=3) \\ {[\mathrm{Ba}] \pm \text { Std. Dev }} \\ \% \text { RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \\ \hline \end{array}$ | 222 | $\begin{gathered} 96.7 \mathrm{ppm} \pm 23.4 \\ 24.44 \% \\ 0.031 \\ 70.8 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 216 \mathrm{ppm} \pm 6 \\ 2.72 \% \\ 0.85 \\ 6.41 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{r} \hline \text { JR-1 }(\mathrm{n}=3) \\ {[\mathrm{Ba}] \pm \text { Std. Dev }} \\ \% \text { RSD } \end{array}$ <br> Signal to Noise Ratio LOD | 50.3 | Not detected in spectrum; below detection limit | $\begin{gathered} 53.8 \mathrm{ppm} \pm 0.4 \\ 0.68 \% \\ 0.29 \\ 4.84 \mathrm{ppm} \\ \hline \end{gathered}$ |

### 5.4 Other elements of interest

In examining the spectra further, other isotopes of interest were identified. Two in particular, As-76 and Br-82, showed clear peaks in the spectra and have presented the opportunity to study the affect of these elements on the sourcing of obsidian. Arsenic has
never before been applied to obsidian studies, mostly because of the difficulty in measurement due to such low concentrations. Bromine has been occasionally examined, but has only been thought useful in the sourcing of obsidian in Turkey.

### 5.4.1 Arsenic Analysis

Arsenic is a trace element in obsidian that is usually only present in levels ranging from a few parts-per-million to low tens of parts-per-million. Because it has been difficult to measure via traditional NAA methods, where it usually is barely above background if the peak appears at all, it has not been applied to obsidian sourcing studies. Since the cadmium-covered irradiation appears to offer an alternative way of measuring arsenic, the presence of the arsenic peak in the spectrum has been examined further, both to confirm that the peak is due to arsenic in the sample and to determine the sensitivity of this method for its analysis.

### 5.4.1.1 Neutron capture of arsenic

Arsenic has only one naturally-abundant isotope, As-75. As-75 can undergo a neutron capture reaction with both thermal and epithermal neutrons to form As-76, which has a half-life of 26.3 hours. The typical reaction examined for NAA is the thermal neutron reaction. As-75 has a thermal neutron capture cross section of 4.5 barns. Even though the abundance of the target isotope is high and the thermal cross section is moderately high, very little activity due to As-76 is detectable during either the typical long or short irradiation analysis. However, As-75 also has a resonance cross section of 61 barns. With the use of the cadmium can to reduce the thermal neutron flux, the
background due to activity from other thermal neutron reactions is reduced and the probability of observing an epithermal neutron reaction increases. Thus, a strong peak at $559.1 \mathrm{keV}(44.60 \%)$ is visible and quantifiable. Figure 5.5 shows the comparison of the $599.1-\mathrm{keV}$ arsenic peak in both a standard irradiation and the cadmium-covered irradiation.

Figure 5.5: Gamma spectra of SRM 278 (obsidian) depicting peak for arsenic analysis


### 5.4.1.2 Is the peak really due to arsenic?

One concern about this peak was that the half-life of As-76 was short compared to the time the samples were allowed to decay (seven days for the initial experiments). By the time of counting, the As-76 had already been through about 6.5 half-lives. It seemed unlikely that such an intense peak would still be present after that period of time,
especially since the concentrations of arsenic are usually low in obsidian. To confirm that the peak was due to arsenic, two tests were performed. First, the theoretical activity due to As-76 was estimated using the measured thermal and epithermal neutron fluxes and compared to the activity seen in the spectrum for NIST SRM 278 (obsidian). These calculations showed that the intensity of the peak was consistent with the activity produced due to the As-75 (n, $\gamma$ ) As-76 reaction. A value of 3590 counts was calculated as a combination of the thermal and resonance neutron reactions and a value of 3827 counts was observed.

An additional test was to follow the decay of the peak over a period of several days and determine if the decay matched that of As-76. This would prove that the peak was due to As-76 and that it was due to As-76 produced during the irradiation (i.e., not being produced from any other decay after the irradiation). Several standards including the NIST SRM 278 (obsidian), SRM 1633a (fly ash), and JGS JR-1 (rhyolite) and two obsidian samples were irradiated for 48 hours, allowed to decay for about 4.5 days, and then counted several times over the next 5 days. Each count was for 30 minutes and each sample was counted about four times during the course of one day. A total of 17 counts were performed for each of the samples. The decay curve achieved from the analysis of the $559-\mathrm{keV}$ peak in SRM 1633a (fly ash) is shown in Figure 5.6 below.

Figure 5.6: Decay of the 559.1-keV gamma peak in SRM 1633a (fly ash)


The exponential fit gave an $R^{2}$ value of 0.9999 and a decay constant of $0.0262 \mathrm{~h}^{-}$, which gives a half-life of $26.408 \pm 0.053$ hours, a $0.41 \%$ difference from the literature value of 26.3 hours. The results and uncertainties for all standards are listed in Table 5.5. The results calculated from the SRM278 standard have a greater error than the other two, but this is most likely due to this standard containing the smallest concentration of arsenic. Since less activity due to arsenic was produced, the counting statistics were not as good for this sample. However, the calculated half-life is still very close to the true value. These results confirm that the $559-\mathrm{keV}$ peak is due to the decay of As-76 produced during the cadmium-covered irradiation.

Table 5.5 Results from the decay experiment for As-76

| Standard | Decay constant <br> $(\lambda)$ | $\sigma_{\lambda}$ | Half-life (ty/2) | $\sigma_{1 / 2}$ |
| :---: | :---: | :---: | :---: | :---: |
| SRM-278 <br> $(n=3)$ | $0.025560 \mathrm{~h}^{-}$ | $0.00036 \mathrm{~h}^{-}$ | 27.1 h | 0.38 h |
| SRM-1633a <br> $(n=3)$ | $0.026247 \mathrm{~h}^{-}$ | $0.000052 \mathrm{~h}^{-}$ | 26.4 h | 0.053 h |
| JR-1 <br> $(n=3)$ | $0.026043 \mathrm{~h}^{-}$ | $0.00017 \mathrm{~h}^{-}$ | 26.6 h | 0.17 h |

### 5.4.1.3 Arsenic results

As for titanium and barium, the limit of detection, relative standard deviation, and the signal-to-noise ratios were calculated for each standard analyzed in the initial experiment. The results of these calculations are found in Table 5.6 below. In addition, these were also calculated for several of these standards after an analysis involving a 4.5day decay instead of a 7-day decay (also shown in Table 5.6).

This procedure for the analysis of arsenic in geological samples proves to have good results with low limits of detection and good precision for the measurements. The 5-day decay allowed even better precision and a better detection limit to be achieved.

Several of the standards had not been analyzed with the second set of samples for the 5-day decay because the concentration of arsenic was so low in these materials. The main goal of this second set had been to follow the decay, and since space inside the cadmium can is limited, standards with larger concentrations, which would still have a signal after several days of counting for good counting statistics, were chosen for analysis.

Table 5.6: Arsenic analysis results: comparison of results after a 7-day or 5-day decay
(Concentrations were calculated using NIST SRM 278 as a primary standard)

|  | $\begin{gathered} \text { Literature } \\ \text { concentrations }{ }^{10} \\ \hline \end{gathered}$ | $\begin{gathered} \text { As-76 } \\ \text { (7-day decay) } \end{gathered}$ | $\begin{gathered} \text { As-76 } \\ \text { (5-day decay) } \end{gathered}$ |
| :---: | :---: | :---: | :---: |
| SRM 278 <br> (primary standard) Signal to Noise Ratio LOD | 4.7 ppm | $\begin{gathered} \mathrm{N} / \mathrm{A} \\ 0.75 \\ 0.2 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} \mathrm{N} / \mathrm{A} \\ 1.18 \\ 0.09 \mathrm{ppm} \end{gathered}$ |
| SRM 1633a ( $\mathrm{n}=3$ ) <br> $[\mathrm{As}] \pm$ Std. Dev <br> \% RSD <br> Signal to Noise Ratio <br> LOD | 145 ppm | $\begin{gathered} 150.4 \mathrm{ppm} \pm 1.8 \\ 1.20 \% \\ 9.70 \\ 0.497 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 141.8 \mathrm{ppm} \pm 1.2 \\ 0.83 \% \\ 34.25 \\ 0.109 \mathrm{ppm} \\ \hline \end{gathered}$ |
| $\begin{array}{\|l} \hline \text { AGV-1 }(\mathrm{n}=3) \\ {[\mathrm{As}] \pm \text { Std. Dev }} \\ \% \text { RSD } \end{array}$ <br> Signal to Noise Ratio LOD | 0.84 ppm | $\begin{gathered} 1.22 \mathrm{ppm} \pm 0.28 \\ 22.70 \% \\ 0.18 \\ 0.147 \mathrm{ppm} \\ \hline \end{gathered}$ | Not analyzed |
| $\begin{aligned} & \text { JA-1 }(\mathrm{n}=3) \\ & {[\mathrm{As}] \pm \text { Std. Dev }} \\ & \% \text { RSD } \end{aligned}$ <br> Signal to Noise Ratio <br> LOD | 2.78 ppm | $\begin{gathered} 3.09 \mathrm{ppm} \pm 0.17 \\ 5.60 \% \\ 0.48 \\ 0.214 \mathrm{ppm} \\ \hline \end{gathered}$ | Not analyzed |
| $\begin{array}{r} \hline \text { JB-2 }(\mathrm{n}=3) \\ {[\mathrm{As}] \pm \text { Std. Dev }} \\ \text { \% RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \\ \hline \end{array}$ | 2.87 ppm | $\begin{gathered} 3.63 \mathrm{ppm} \pm 0.37 \\ 10.15 \% \\ 0.28 \\ 0.252 \mathrm{ppm} \\ \hline \end{gathered}$ | Not analyzed |
| $\begin{array}{r} \hline \text { JR-1 }(\mathrm{n}=3) \\ {[\mathrm{As}] \pm \text { Std. Dev }} \\ \% \text { RSD } \\ \text { Signal to Noise Ratio } \\ \text { LOD } \end{array}$ | 16.3 ppm | $\begin{gathered} 15.9 \mathrm{ppm} \pm 0.34 \\ 2.13 \% \\ 2.08 \\ 0.253 \mathrm{ppm} \\ \hline \end{gathered}$ | $\begin{gathered} 16.43 \mathrm{ppm} \pm 0.21 \\ 1.28 \% \\ 4.82 \\ 0.0912 \mathrm{ppm} \\ \hline \end{gathered}$ |

### 5.4.2 Bromine analysis

Bromine is an element not often used for obsidian studies. However, it is thought to be useful in distinguishing between two Turkish obsidian sources, Bingol and Nemrut, that are otherwise almost identical. ${ }^{11}$ However, it is not usually determined during the
typical NAA procedure because of the low concentration in the standards and the assumed low concentration in the samples. Using epithermal neutron activation, however, the bromine can be more easily quantified as the signal is improved with the lower background.

### 5.4.2.1 Neutron capture reactions of bromine

For the analysis of bromine, there are two naturally abundant isotopes that can undergo reactions by capturing neutrons. The first, Br-79 (50.69\%), can react via an ( $\mathrm{n}, \gamma$ ) reaction to form $\mathrm{Br}-80$. $\mathrm{Br}-79$ has a thermal neutron cross section of 8.60 barns and $\mathrm{Br}-$ 80 has a half-life of 17.68 minutes, which would make it perfect for a short-irradiation analysis procedure. However, the gamma rays emitted during the decay of $\mathrm{Br}-80$ are so low in abundance ( 616.60 keV is the most abundant peak at $6.70 \%$ ), that it is difficult to see the peak above background unless the concentration of bromine in the sample is high. $\mathrm{Br}-79$ can also capture a neutron to form $\mathrm{Br}-80 \mathrm{~m}\left(\mathrm{t}_{1 / 2}=4.42\right.$ hours $)$ with a thermal cross section of 2.40 barns, but the only gamma ray emitted from the decay of $\mathrm{Br}-80 \mathrm{~m}$ to $\mathrm{Br}-80$ is 37.05 keV . Such a low energy peak can be difficult to distinguish in a complex gamma spectrum.

The other naturally-occurring isotope of bromine is $\mathrm{Br}-81$ ( $49.31 \%$ abundant). $\mathrm{Br}-81$ can undergo an ( $\mathrm{n}, \gamma$ ) reaction to form $\mathrm{Br}-82(\mathrm{t} 1 / 2=35.3$ hours $)$ with a thermal cross section of 0.26 barns and a resonance cross section of 17 barns. Br- 81 can also react to form $\mathrm{Br}-82 \mathrm{~m}(\mathrm{t} 1 / 2=6.13$ minutes $)$ with a thermal cross section of 2.43 barns and a resonance cross section of 34 barns. Br- 82 m then decays $99.6 \%$ of the time to $\mathrm{Br}-82$. When $\mathrm{Br}-82$ decays, there are several gamma rays emitted; the most abundant are 776.5
$\mathrm{keV}(83.54 \%)$ and $554.35 \mathrm{keV}(70.76 \%)$. When the cadmium can is used for the analysis, these peaks due to $\mathrm{Br}-82$ are clearly visible in the spectrum (Figure 5.7).

Figure 5.7: Gamma spectra of SRM 278 (obsidian) depicting peaks for bromine analysis


### 5.4.2.2 Br-82 decay test

Even though the half-life for $\mathrm{Br}-82$ is longer than that of the arsenic isotope tested before, the decay of the $776-\mathrm{keV}$ peak in the gamma spectrum was followed over the course of several days under the same conditions described for the arsenic peak above.

Figure 5.8 shows the decay curve of the peak.
Though the results are not quite as precise as those for the arsenic peak, an Rsquared value of 0.9918 indicates a good fit for the trendline. The half-life calculated
from the decay constant $\left(0.0197 \mathrm{~h}^{-}\right)$is 35.19 hours, a $0.32 \%$ difference from the literature value of 35.3 hours. This peak is definitely due to the decay of $\mathrm{Br}-82$ and it is unlikely that anything is significantly interfering with it.

Figure 5.8: Decay of the 776-keV gamma peak in SRM 278 (obsidian)


### 5.4.2.3 Bromine results

Because several of the standards analyzed during these irradiations did not have published values for bromine, and those that are known have large uncertainties, the accuracy and limits of detection could not be calculated. With the proper standards with known concentrations of bromine, the determination of bromine in obsidian could be quite accurate.

Two samples of obsidian from each of the two Turkish sources were analyzed as part of the second set of samples in the cadmium-covered irradiation. The samples from one source, Bingol, showed a strong bromine signal (over 10,000 counts for both samples); while the samples from the other source were much lower (only a few hundred counts in the peak). There seems to be a clear difference between the two sources, though more samples would need to be analyzed to confirm this.

### 5.5 Comparison of results from ENAA and x-ray methods

As stated in the introduction, one of the main goals of creating this new methodology was to verify previous results by Merrick and Brown for the sourcing of African obsidian. ${ }^{1-2}$ Since these results were analyzed using both x-ray fluorescence and electron microprobe analysis, it is necessary to be able to directly compare those results for the concentration of titanium to those achieved with this new method. This comparison is made more difficult by the lack of detailed information about the analysis procedure used for both the XRF and EMPA and the lack of a published table of data. Even though a perfect comparison cannot be made between these two data sets, trends in the data can be studied. To achieve this, the same set of Mesoamerican obsidian included in the initial experiments with the cadmium can was sent to be analyzed via EMPA. This allowed a more direct comparison of the two methods, since more information about the instrument and the analysis method was available. These experiments also provided a way to compare the results using x-ray methods from previous studies for the African obsidian and examine the trends between the data sets.

### 5.5.1 Results from the Mesoamerican obsidian samples

A selection of the Mesoamerican samples described above were sent to Ellery Frahm, Electron Microprobe Lab at the University of Minnesota-Twin Cities, to be analyzed via EMPA. The instrument used was a JEOL JXA-8900R "SuperProbe" electron probe microanalyzer, with a wavelength-dispersive detection system, giving this instrument a resolution of 5 eV . Use of a wavelength-dispersive instrument provided a comparison under the most ideal circumstances. The peaks due to barium and titanium are easily resolved with this instrument, allowing for more accurate results from the x-ray method. Figure 5.9 below shows the results from both methods for the concentration of titanium (in parts per million) plotted against each other.

Figure 5.9: Comparison of the results from EMPA versus those from ENAA for Mesoamerican obsidian


The relationship between the results from both methods is very close to one-to-one, and considering the relative standard deviation for the EMPA results (about $10 \%$ for each sample), the differences between the two data sets is minimal. To be sure that the differences were minimal statistically, an ANOVA analysis was performed on the two data sets. ${ }^{12}$ The results of that analysis proved that the two data sets were not significantly different $(\mathrm{p}=0.42)$.

From these results, it can be seen that, if a wavelength-dispersive detection system is used and the barium and titanium peaks can be resolved, then both the x-ray method and the ENAA method yield identical results. This hypothesis can then be extended to the African obsidian results: if the x-ray peaks for titanium and barium were wellresolved, then the results from the x-ray method and the ENAA method should be very similar.

### 5.5.2 Results from the African obsidian samples

Before any comparison between the ENAA method for titanium and the results from previous studies can be discussed, a few major challenges need to be addressed. There is the lack of published tables of the calculated values for the titanium concentrations for the previous studies. However, though tables of numbers are not given, the concentration of titanium is used in many plots to demonstrate its ability to distinguish source samples. The values used here, in contrast to those obtained by ENAA, have been estimated from these plots.

Secondly, the same exact samples as in previous studies were not available for analysis using the ENAA method. Instead, samples from the same or similar locations
were used for comparison. In most cases, three or four samples from each location (where available) were analyzed to account for any intra-source variation in the titanium concentration. Because of this limitation, only a subset of the localities listed in the published works was analyzed. The sample localities were chosen for the availability of samples from the same region for a better comparison.

Shown below in Figure 5.10 is the comparison of the results from each method, just as in Figure 5.9. Here, the results for the African obsidian samples have been added to the same plot for a more direct comparison. The slope of the trendline for the African obsidian data is greater than 1 , showing that the concentrations estimated for the previous studies are generally greater than those determined using the ENAA method. The fit of the data to the trendline, shown in the R-squared value, is also not as good, though this could mostly be due to intra-source variation and error introduced in the estimation of the concentrations. The larger concentrations from the x-ray methods used in previous works could be due to some contribution of barium to the peak intensity.

Figure 5.10: Comparison of the results from EMPA versus those from ENAA for Mesoamerican obsidian and African obsidian


### 5.6 Future work and conclusions

Here a method using epithermal neutron activation analysis for the determination of titanium, barium, arsenic, and bromine concentrations in obsidian has been presented. This method not only provides better limits of detection, precision, and signal-to-noise ratio for each of these elements in obsidian, but also has the potential to be applied to other materials. Future directions for this project include the application of this method to biological materials for the analysis of arsenic and evaluating the potential of this method as an alternative to the current long irradiation procedure.

### 5.6.1 Analysis of arsenic in biological samples

Arsenic is often an element of interest in biological materials, but it is difficult to measure using standard thermal INAA. ${ }^{13}$ Since NAA sample preparation procedures are usually minimal, sample preparation for analysis via NAA would greatly reduce the potential of contamination as compared to other analytical techniques. A small set of biological samples was analyzed using the ENAA method described here. Fifty milligrams of NIST SRM1577 bovine liver were weighed out into quartz vials and sealed under slight vacuum. Unfortunately, the heating of the samples caused the sealed quartz vials to burst, as the sample degraded and released carbon dioxide, oxygen, and other gases. Two potential solutions to the heating problem are reducing the irradiation time or by using larger vials. Reducing the irradiation time does decrease the amount of activity produced during the irradiation, but this would also allow for shorter decay times for the sample before counting and has the potential to prevent the pressure from released gasses reaching the point of rupturing the vial. Using larger vials simply gives the gases more room to expand, thus slowing the buildup of pressure. If the heating problem could be solved, then this method could be used to accurately and precisely measure the concentrations of arsenic in biological samples.

### 5.6.2 An alternative method for standard INAA procedures

ENAA can also be considered for the detection of other elements in the obsidian sample. Many of the rare earth elements that are of interest in sourcing studies have large resonance cross sections. These larger cross sections in addition to the reduction in thermal background can improve the signal-to-noise ratio significantly, leading to better precision and limits of detection for these elements. As described earlier, ENAA has
been demonstrated to increase the sensitivity for many trace elements of interest in geological samples. ${ }^{3,6-7}$ Table 5.7 below lists many of the elements currently analyzed during the long irradiation procedure for obsidian, along with their thermal and resonance cross sections. Theoretical advantage factors (F) were calculated using the Brune ${ }^{14}$ definition (Equation 5.1)
$F_{a}=\frac{R_{i}}{R}$
Equation 5.1
where R and $\mathrm{R}_{\mathrm{i}}$ are the cadmium ratios for the radionuclide and the interferent (in obsidian, usually ${ }^{30} \mathrm{Si}$ and ${ }^{24} \mathrm{Na}$ ). The cadmium ratios were calculated using Equation $5.2^{15}$
$R=\frac{\phi_{t h} \sigma_{t h}+\phi_{\text {epi }} I}{\phi_{\text {epi }} I}$
Equation 5.2
where $\varphi$ is the thermal (th) and epithermal (epi) neutron flux, $\sigma$ is the thermal cross section, and I is the resonance integral.

Table 5.7: Elements of interest for the ENAA procedure

| Target <br> Isotope | ${\text { Thermal } \boldsymbol{\sigma}(\mathbf{b})^{\mathbf{1 0}}}^{\text {Resonance } \mathbf{I}(\mathbf{b})^{\mathbf{1 0}}}$ | $\mathbf{F}$ <br> $(\mathbf{i}=\mathbf{N a - 2 4 )}$ | $\mathbf{F}$ <br> $\mathbf{( i = \mathbf { S i } - \mathbf { 3 0 } )}$ |  |
| :--- | :---: | :---: | :---: | :---: |
| Fe-58 | 1.28 | 1.7 | 2.25 | 1.99 |
| Co-59 | 17.1 | 34.3 | 3.37 | 2.98 |
| Zn-64 | 0.76 | 1.45 | 3.21 | 2.84 |
| Sr-84 | 0.35 | 6.72 | 27.61 | 24.44 |
| Rb-85 | 0.427 | 7.31 | 25.06 | 22.17 |
| Zr-94 | 0.05 | 0.23 | 7.54 | 6.67 |
| Sb-123 | 4.1 | 118 | 38.31 | 33.90 |
| Cs-133 | 26.4 | 390 | 22.06 | 19.53 |
| La-139 | 8.93 | 11.8 | 2.24 | 1.98 |
| Ce-140 | 0.57 | 0.47 | 1.40 | 1.24 |
| Nd-146 | 1.4 | 3.2 | 3.83 | 3.39 |
| Sm-152 | 206 | 2970 | 21.60 | 19.12 |
| Tb-159 | 23.4 | 418 | 25.98 | 22.99 |
| Hf-180 | 13.04 | 35 | 4.48 | 3.97 |
| Ta-181 | 20.5 | 660 | 41.75 | 36.95 |

These advantage factors provide a way of evaluating the potential increase in sensitivity for these elements. Other isotopes measured during the standard long irradiation procedure, such as $\mathrm{Sc}-45(\mathrm{Sc}-46)$, Eu-151 (Eu-152), Yb-174 (Yb-175), and Lu-176 (Lu177), show a decrease in the sensitivity because their resonance integrals are much smaller than the thermal cross section. However, for these same isotopes, the thermal cross section is so large (i.e., Eu-151 has a thermal cross section of 5900 b) that the reduction in thermal flux may only have a minor impact on the activity produced during the irradiation.

### 5.6.3 Conclusions

The results described in this section have shown the versatility of ENAA toward analyzing archeological samples, and also provide greater data per analysis for each sample. These methods allow for geologists and archeologists to obtain greater amounts of data, with better precision and accuracy, while only analyzing small samples. As described in Chapter 1, NAA methods are minimally destructive and would lend to application to samples of artifacts or for use in creation of standards for development of an x-ray based analysis.

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Appendix 1 - Limestone quarry information for the BNL Egyptian limestone samples.

| Anid | Quarry |
| :---: | :---: |
| EB0037 | Tura |
| EB0038 | Beni Hassan |
| EB0039 | Deir Abu Hennis |
| EB0040 | Deir Abu Hennis |
| EB0041 | Deir Abu Hennis |
| EB0042 | Deir Abu Hennis |
| EB0043 | Deir Abu Hennis |
| EB0044 | Deir Abu Hennis |
| EB0045 | Deir Abu Hennis |
| EB0046 | Deir Abu Hennis |
| EB0047 | Deir Abu Hennis |
| EB0048 | Deir Abu Hennis |
| EB0049 | Deir Abu Hennis |
| EB0050 | Deir Abu Hennis |
| EB0051 | Beni Hassan |
| EB0052 | Beni Hassan |
| EB0053 | Beni Hassan |
| EB0054 | Beni Hassan |
| EB0055 | Beni Hassan |
| EB0056 | Beni Hassan |
| EB0057 | Beni Hassan |
| EB0058 | Beni Hassan |
| EB0059 | Beni Hassan |
| EB0060 | Beni Hassan |
| EB0061 | Beni Hassan |
| EB0062 | Beni Hassan |
| EB0063 | Beni Hassan |
| EB0064 | Quseir el Amarna |
| EB0065 | Quseir el Amarna |
| EB0066 | Quseir el Amarna |
| EB0067 | Quseir el Amarna |
| EB0068 | Quseir el Amarna |
| EB0069 | Quseir el Amarna |
| EB0070 | Quseir el Amarna |
| EB0071 | Quseir el Amarna |
| EB0072 | Quseir el Amarna |
| EB0073 | Quseir el Amarna |
| EB0074 | Quseir el Amarna |
| EB0075 | Quseir el Amarna |
| EB0076 | Quseir el Amarna |
| EB0078 | Tura |
| EB0079 | Tura |

Appendix 1 - Limestone quarry information for the BNL Egyptian limestone samples, continued.

| Anid | Quarry |
| :---: | :---: |
| EB0080 | Tura |
| EB0081 | Tura |
| EB0082 | Tura |
| EB0083 | Tura |
| EB0084 | Tura |
| EB0085 | Tura |
| EB0086 | Tura |
| EB0087 | Quseir el Amarna |
| EB0088 | Quseir el Amarna |
| EB0089 | Quseir el Amarna |
| EB0090 | Quseir el Amarna |
| EB0091 | Quseir el Amarna |
| EB0092 | Quseir el Amarna |
| EB0093 | Quseir el Amarna |
| EB0094 | Quseir el Amarna |
| EB0095 | Quseir el Amarna |
| EB0096 | Quseir el Amarna |
| EB0097 | Quseir el Amarna |
| EB0098 | Quseir el Amarna |
| EB0099 | Quseir el Amarna |
| EB0100 | Maasara |
| EB0101 | Maasara |
| EB0102 | Maasara |
| EB0103 | Maasara |
| EB0104 | Maasara |
| EB0105 | Maasara |
| EB0106 | Maasara |
| EB0107 | Maasara |
| EB0108 | Maasara |
| EB0109 | Maasara |
| EB0110 | Maasara |
| EB0111 | Maasara |
|  | Maasara |

Appendix 2 - Concentration data from NAA for limestone from Egypt (BNL samples). All concentrations are in parts per million (ppm)

| Sample ID | Al | As | Ba | Ca | Ce | Co | Cr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0037 | 0.0 | 0.5455 | 12.3 | 418583.8 | 1.3434 | 2.6209 | 6.4077 |
| EB0038 | 0.0 | 0.2333 | 8.4 | 412720.7 | 0.3843 | 0.1550 | 6.5697 |
| EB0039 | 0.0 | 1.5955 | 0.0 | 415623.6 | 1.0475 | 0.4255 | 6.4056 |
| EB0040 | 0.0 | 0.1288 | 0.0 | 416673.1 | 0.6742 | 0.1300 | 8.5082 |
| EB0041 | 420.7 | 0.2155 | 0.0 | 414381.7 | 0.6485 | 0.1212 | 8.1787 |
| EB0042 | 0.0 | 0.1708 | 8.2 | 415773.8 | 0.5301 | 0.2197 | 7.0762 |
| EB0043 | 506.8 | 0.3500 | 0.0 | 415378.6 | 0.7300 | 0.3476 | 7.9735 |
| EB0044 | 0.0 | 0.1834 | 0.0 | 410863.2 | 0.4060 | 0.1015 | 9.4522 |
| EB0045 | 0.0 | 0.2249 | 6.3 | 409493.2 | 0.5205 | 0.3309 | 8.8425 |
| EB0046 | 0.0 | 0.3627 | 0.0 | 419790.8 | 0.7873 | 0.2548 | 5.1872 |
| EB0047 | 0.0 | 0.2306 | 0.0 | 414061.5 | 0.4985 | 0.1297 | 5.9179 |
| EB0048 | 0.0 | 0.1144 | 0.0 | 427300.2 | 0.4954 | 0.0652 | 11.0820 |
| EB0049 | 0.0 | 0.1885 | 0.0 | 416124.9 | 0.4955 | 0.0733 | 6.6187 |
| EB0050 | 0.0 | 0.2078 | 8.7 | 419983.0 | 0.5685 | 0.0711 | 12.4380 |
| EB0051 | 0.0 | 0.3490 | 18.5 | 421036.0 | 1.2756 | 0.3489 | 5.6239 |
| EB0052 | 0.0 | 0.2230 | 29.9 | 422701.8 | 0.8143 | 0.4683 | 4.9293 |
| EB0053 | 0.0 | 0.3917 | 17.3 | 419393.7 | 1.2765 | 1.4716 | 8.3478 |
| EB0054 | 0.0 | 0.0000 | 23.6 | 417692.0 | 1.1292 | 0.2967 | 6.7502 |
| EB0055 | 0.0 | 0.3975 | 9.1 | 419510.0 | 0.4598 | 0.1281 | 6.6555 |
| EB0056 | 0.0 | 0.0000 | 25.6 | 419910.9 | 0.6907 | 0.1059 | 6.1519 |
| EB0057 | 0.0 | 0.0000 | 40.8 | 421021.7 | 0.8343 | 0.1508 | 8.0770 |
| EB0058 | 0.0 | 0.2227 | 34.8 | 413957.1 | 2.0175 | 0.2066 | 6.6609 |
| EB0059 | 0.0 | 0.0000 | 13.3 | 431619.2 | 0.7433 | 0.2261 | 5.2167 |
| EB0060 | 0.0 | 0.0000 | 21.6 | 424844.3 | 0.8892 | 0.1179 | 5.9596 |
| EB0061 | 0.0 | 0.2885 | 19.2 | 423992.4 | 0.4810 | 0.1155 | 6.7758 |
| EB0062 | 0.0 | 0.0000 | 11.5 | 420263.3 | 0.5378 | 0.1126 | 4.9102 |
| EB0063 | 0.0 | 0.2247 | 25.1 | 416502.9 | 0.9440 | 0.1238 | 6.3768 |
| EB0064 | 903.5 | 0.2642 | 15.6 | 412956.8 | 0.7134 | 0.1740 | 9.5427 |
| EB0065 | 0.0 | 0.2446 | 26.2 | 411792.6 | 0.4699 | 0.1441 | 7.6373 |
| EB0066 | 0.0 | 0.1961 | 11.7 | 422219.4 | 0.6991 | 0.1156 | 10.7338 |
| EB0067 | 0.0 | 0.3618 | 0.0 | 411831.3 | 0.6619 | 0.0974 | 3.1013 |
| EB0068 | 0.0 | 0.0000 | 7.5 | 426331.6 | 0.3750 | 0.1455 | 3.7835 |
| EB0069 | 0.0 | 0.3348 | 12.8 | 420313.6 | 0.2123 | 0.1366 | 2.1585 |
| EB0070 | 0.0 | 0.3614 | 22.1 | 416647.2 | 0.3693 | 0.1119 | 4.0702 |
| EB0071 | 0.0 | 0.4705 | 0.0 | 413524.0 | 0.3503 | 0.1532 | 3.3583 |
| EB0072 | 0.0 | 0.0000 | 0.0 | 427257.3 | 0.3818 | 0.0889 | 3.6055 |
| EB0073 | 0.0 | 0.0000 | 0.0 | 417463.5 | 0.3464 | 0.0393 | 3.2798 |
| EB0074 | 1047.0 | 0.3009 | 0.0 | 419326.7 | 0.3881 | 0.0823 | 2.7328 |
| EB0075 | 0.0 | 0.0000 | 10.5 | 416278.0 | 0.3911 | 0.1567 | 3.5608 |
| EB0076 | 0.0 | 0.0000 | 10.0 | 416840.5 | 0.5929 | 0.1492 | 4.0360 |
| EB0078 | 3583.0 | 0.7841 | 6.6 | 388883.9 | 2.7446 | 0.1394 | 13.1269 |
| EB0079 | 0.0 | 0.0000 | 10.9 | 374084.0 | 2.1905 | 0.1616 | 12.6403 |
| EB0080 | 878.9 | 0.0000 | 10.7 | 385504.9 | 3.2841 | 0.0849 | 14.0466 |
| EB0081 | 7500.2 | 0.9092 | 0.0 | 316533.2 | 9.6137 | 0.3675 | 35.3593 |
| EB0082 | 4063.1 | 1.2900 | 127.6 | 398910.1 | 3.8534 | 0.2664 | 16.8637 |

Appendix 2, continued

| Sample ID | Al | As | Ba | Ca | Ce | Co | Cr |
| :--- | ---: | ---: | ---: | :---: | :---: | :---: | ---: |
| EB0083 | 4038.5 | 0.7987 | 12.1 | 381362.8 | 4.0728 | 0.3398 | 17.2407 |
| EB0084 | 0.0 | 0.9044 | 0.0 | 334532.1 | 5.8749 | 0.3678 | 27.5630 |
| EB0085 | 0.0 | 0.9240 | 13.1 | 406625.8 | 2.8517 | 0.1786 | 6.1525 |
| EB0086 | 0.0 | 0.0000 | 916.6 | 401046.4 | 1.1430 | 0.6990 | 5.9964 |
| EB0087 | 3071.0 | 0.1262 | 21.7 | 411073.7 | 0.5129 | 0.4038 | 5.9343 |
| EB0088 | 0.0 | 0.2465 | 0.0 | 417227.7 | 0.4259 | 0.4238 | 4.7781 |
| EB0089 | 0.0 | 0.2977 | 14.2 | 411580.3 | 0.3442 | 0.2936 | 5.2674 |
| EB0090 | 0.0 | 1.9382 | 14.5 | 405657.9 | 0.6172 | 0.1849 | 3.1207 |
| EB0091 | 0.0 | 0.1682 | 14.7 | 403082.3 | 0.3236 | 0.2595 | 3.9388 |
| EB0092 | 0.0 | 0.5488 | 15.1 | 409432.2 | 0.1724 | 0.1548 | 2.7042 |
| EB0093 | 0.0 | 0.3847 | 23.5 | 406409.1 | 0.3175 | 0.1829 | 3.2048 |
| EB0094 | 0.0 | 0.0000 | 9.0 | 412418.4 | 0.2551 | 0.1520 | 2.6628 |
| EB0095 | 0.0 | 0.1070 | 11.1 | 410423.5 | 0.2388 | 0.1922 | 2.4222 |
| EB0096 | 0.0 | 0.3281 | 11.3 | 414472.7 | 0.3016 | 0.1689 | 2.7161 |
| EB0097 | 0.0 | 0.4340 | 12.9 | 415494.5 | 0.4710 | 0.1530 | 4.4192 |
| EB0098 | 0.0 | 0.4275 | 13.2 | 407936.4 | 0.3964 | 0.1219 | 3.8574 |
| EB0099 | 0.0 | 0.1410 | 14.9 | 405628.3 | 0.2857 | 0.1015 | 3.3788 |
| EB0100 | 5616.4 | 3.0085 | 25.1 | 326187.1 | 9.8801 | 0.6885 | 26.9994 |
| EB0101 | 5298.6 | 1.1799 | 21.7 | 357753.9 | 6.0426 | 0.2638 | 25.7383 |
| EB0102 | 1574.8 | 0.5324 | 0.0 | 371152.4 | 2.7821 | 0.2047 | 14.7879 |
| EB0103 | 2760.6 | 0.5463 | 0.0 | 385744.1 | 3.2425 | 0.1791 | 14.9198 |
| EB0104 | 2003.5 | 0.0000 | 20.5 | 390019.9 | 3.4290 | 0.2088 | 14.4129 |
| EB0105 | 3255.3 | 0.8447 | 16.6 | 380537.9 | 5.0678 | 0.4850 | 21.7251 |
| EB0106 | 3460.0 | 1.4293 | 18.7 | 377277.2 | 6.1509 | 0.4128 | 21.2248 |
| EB0107 | 3191.3 | 2.1156 | 19.1 | 395556.6 | 4.0286 | 0.2561 | 19.7098 |
| EB0108 | 3905.6 | 0.9570 | 14.7 | 387576.4 | 3.5170 | 0.2698 | 14.7060 |
| EB0109 | 4310.5 | 2.5183 | 28.2 | 365227.3 | 7.7105 | 0.4060 | 25.8313 |
| EB0110 | 5556.8 | 3.8984 | 25.8 | 372107.7 | 7.2016 | 1.4240 | 28.3854 |
| EB0111 | 4495.3 | 0.3488 | 28.3 | 368428.1 | 6.8337 | 0.2191 | 28.8855 |
| EB0112 | 2955.7 | 1.1889 | 10.3 | 375009.0 | 3.4182 | 0.2247 | 15.4478 |

Appendix 2, continued

| Sample ID | Cs | Dy | Eu | Fe | Hf | K | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0037 | 0.0000 | 0.1556 | 0.0411 | 338.5 | 0.0101 | 0.0 | 0.6476 |
| EB0038 | 0.0092 | 0.0000 | 0.0044 | 235.7 | 0.0053 | 0.0 | 0.1713 |
| EB0039 | 0.0268 | 0.2372 | 0.0433 | 769.3 | 0.0242 | 0.0 | 1.1441 |
| EB0040 | 0.0000 | 0.0813 | 0.0122 | 64.4 | 0.0000 | 0.0 | 0.4115 |
| EB0041 | 0.0209 | 0.1071 | 0.0148 | 131.9 | 0.0119 | 0.0 | 0.4904 |
| EB0042 | 0.0300 | 0.0000 | 0.0136 | 131.8 | 0.0097 | 0.0 | 0.4396 |
| EB0043 | 0.0296 | 0.0802 | 0.0164 | 160.5 | 0.0157 | 0.0 | 0.5109 |
| EB0044 | 0.0135 | 0.0000 | 0.0108 | 101.0 | 0.0072 | 0.0 | 0.3631 |
| EB0045 | 0.0000 | 0.0000 | 0.0112 | 125.2 | 0.0080 | 0.0 | 0.4018 |
| EB0046 | 0.0176 | 0.1023 | 0.0258 | 246.5 | 0.0090 | 0.0 | 0.5264 |
| EB0047 | 0.0109 | 0.0000 | 0.0114 | 168.4 | 0.0097 | 0.0 | 0.3999 |
| EB0048 | 0.0207 | 0.0000 | 0.0113 | 101.3 | 0.0057 | 0.0 | 0.4099 |
| EB0049 | 0.0083 | 0.0000 | 0.0078 | 96.4 | 0.0083 | 0.0 | 0.3020 |
| EB0050 | 0.0169 | 0.0000 | 0.0078 | 111.3 | 0.0106 | 0.0 | 0.3379 |
| EB0051 | 0.0136 | 0.0000 | 0.0364 | 246.9 | 0.0109 | 0.0 | 0.5944 |
| EB0052 | 0.0000 | 0.0000 | 0.0052 | 178.8 | 0.0000 | 0.0 | 0.2804 |
| EB0053 | 0.0393 | 0.0656 | 0.0166 | 629.7 | 0.0794 | 0.0 | 0.5856 |
| EB0054 | 0.0150 | 0.0000 | 0.0086 | 114.1 | 0.0115 | 0.0 | 0.4313 |
| EB0055 | 0.0000 | 0.0000 | 0.0045 | 140.5 | 0.0000 | 0.0 | 0.1625 |
| EB0056 | 0.0000 | 0.0000 | 0.0023 | 73.4 | 0.0000 | 231.0 | 0.1839 |
| EB0057 | 0.0000 | 0.0000 | 0.0060 | 64.8 | 0.0067 | 0.0 | 0.3055 |
| EB0058 | 0.0000 | 0.0000 | 0.0039 | 130.6 | 0.0031 | 0.0 | 0.4607 |
| EB0059 | 0.0000 | 0.0000 | 0.0000 | 58.8 | 0.0000 | 0.0 | 0.2116 |
| EB0060 | 0.0000 | 0.0000 | 0.0062 | 53.9 | 0.0148 | 0.0 | 0.3235 |
| EB0061 | 0.0000 | 0.0000 | 0.0045 | 93.2 | 0.0000 | 0.0 | 0.1787 |
| EB0062 | 0.0000 | 0.0000 | 0.0035 | 47.8 | 0.0000 | 0.0 | 0.1766 |
| EB0063 | 0.0000 | 0.0000 | 0.0080 | 195.8 | 0.0070 | 0.0 | 0.2887 |
| EB0064 | 0.0389 | 0.0000 | 0.0112 | 388.5 | 0.0199 | 0.0 | 0.4677 |
| EB0065 | 0.0217 | 0.0000 | 0.0090 | 157.3 | 0.0167 | 0.0 | 0.3206 |
| EB0066 | 0.0335 | 0.0500 | 0.0155 | 265.0 | 0.0213 | 0.0 | 0.4821 |
| EB0067 | 0.0151 | 0.0000 | 0.0085 | 245.1 | 0.0916 | 0.0 | 0.2985 |
| EB0068 | 0.0089 | 0.0000 | 0.0113 | 164.4 | 0.0120 | 0.0 | 0.2767 |
| EB0069 | 0.0000 | 0.0000 | 0.0021 | 149.4 | 0.0119 | 0.0 | 0.0799 |
| EB0070 | 0.0124 | 0.0000 | 0.0052 | 127.2 | 0.0076 | 0.0 | 0.1539 |
| EB0071 | 0.0184 | 0.0000 | 0.0050 | 137.7 | 0.0154 | 0.0 | 0.1626 |
| EB0072 | 0.0138 | 0.0000 | 0.0051 | 116.4 | 0.0111 | 0.0 | 0.1624 |
| EB0073 | 0.0200 | 0.0000 | 0.0043 | 115.3 | 0.0140 | 0.0 | 0.1704 |
| EB0074 | 0.0174 | 0.0000 | 0.0051 | 87.8 | 0.0127 | 0.0 | 0.1555 |
| EB0075 | 0.0165 | 0.0000 | 0.0054 | 162.6 | 0.0110 | 0.0 | 0.1308 |
| EB0076 | 0.0335 | 0.0000 | 0.0090 | 238.0 | 0.0284 | 0.0 | 0.2329 |
| EB0078 | 0.1628 | 0.2925 | 0.0570 | 1358.2 | 0.2227 | 0.0 | 1.5065 |
| EB0079 | 0.1637 | 0.0000 | 0.0447 | 666.4 | 0.1933 | 0.0 | 1.1657 |
| EB0080 | 0.1278 | 0.2652 | 0.0779 | 380.0 | 0.2013 | 0.0 | 1.7598 |
| EB0081 | 0.5563 | 0.5851 | 0.2223 | 4464.1 | 0.6474 | 1949.4 | 5.2131 |
| EB0082 | 0.2129 | 0.2952 | 0.0736 | 1689.9 | 0.3003 | 168.2 | 1.9644 |

Appendix 2, continued

| Sample ID | Cs | Dy | Eu | Fe | Hf | K | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0083 | 0.2321 | 0.3679 | 0.0820 | 1073.8 | 0.2695 | 0.0 | 2.1546 |
| EB0084 | 0.3315 | 0.2126 | 0.1168 | 1948.2 | 0.3915 | 0.0 | 2.9454 |
| EB0085 | 0.0481 | 0.1965 | 0.0639 | 516.6 | 0.3125 | 0.0 | 1.2717 |
| EB0086 | 0.0510 | 0.0883 | 0.0286 | 320.6 | 0.1392 | 555.6 | 0.5361 |
| EB0087 | 0.0235 | 0.0000 | 0.0071 | 264.7 | 0.0143 | 510.5 | 0.1837 |
| EB0088 | 0.0000 | 0.0000 | 0.0081 | 193.9 | 0.0053 | 0.0 | 0.1319 |
| EB0089 | 0.0000 | 0.0000 | 0.0042 | 217.3 | 0.0000 | 0.0 | 0.0893 |
| EB0090 | 0.0192 | 0.0000 | 0.0079 | 196.7 | 0.0179 | 0.0 | 0.1830 |
| EB0091 | 0.0092 | 0.0000 | 0.0047 | 266.5 | 0.0000 | 0.0 | 0.0903 |
| EB0092 | 0.0000 | 0.0000 | 0.0039 | 95.9 | 0.0000 | 0.0 | 0.1137 |
| EB0093 | 0.0099 | 0.0000 | 0.0037 | 126.6 | 0.0000 | 1061.9 | 0.1396 |
| EB0094 | 0.0149 | 0.0000 | 0.0051 | 107.4 | 0.0124 | 391.1 | 0.1406 |
| EB0095 | 0.0152 | 0.0000 | 0.0049 | 109.0 | 0.0088 | 0.0 | 0.1400 |
| EB0096 | 0.0000 | 0.0000 | 0.0058 | 82.7 | 0.0058 | 457.5 | 0.1492 |
| EB0097 | 0.0118 | 0.0000 | 0.0038 | 83.3 | 0.0066 | 0.0 | 0.1757 |
| EB0098 | 0.0180 | 0.0000 | 0.0062 | 82.9 | 0.0000 | 0.0 | 0.1667 |
| EB0099 | 0.0000 | 0.0000 | 0.0035 | 79.1 | 0.0000 | 0.0 | 0.1611 |
| EB0100 | 0.4357 | 0.9859 | 0.2293 | 3867.2 | 0.5617 | 1612.9 | 5.1436 |
| EB0101 | 0.3759 | 0.3572 | 0.0959 | 1419.6 | 0.4542 | 880.5 | 2.8577 |
| EB0102 | 0.1539 | 0.1887 | 0.0472 | 1100.1 | 0.1890 | 247.7 | 1.3621 |
| EB0103 | 0.1791 | 0.2560 | 0.0615 | 1008.1 | 0.2364 | 0.0 | 1.6657 |
| EB0104 | 0.1760 | 0.1775 | 0.0562 | 560.1 | 0.2141 | 0.0 | 1.6123 |
| EB0105 | 0.2263 | 0.3786 | 0.1008 | 1966.1 | 0.2802 | 0.0 | 2.6163 |
| EB0106 | 0.2689 | 0.4798 | 0.1235 | 2305.8 | 0.3247 | 0.0 | 3.1575 |
| EB0107 | 0.1884 | 0.1993 | 0.0664 | 1323.0 | 0.2921 | 0.0 | 1.7701 |
| EB0108 | 0.1597 | 0.2411 | 0.0634 | 1413.5 | 0.2645 | 0.0 | 1.6249 |
| EB0109 | 0.3039 | 0.5946 | 0.1739 | 2413.2 | 0.3419 | 0.0 | 3.9933 |
| EB0110 | 0.3274 | 0.5760 | 0.1366 | 2678.1 | 0.3908 | 0.0 | 3.5189 |
| EB0111 | 0.3048 | 0.4444 | 0.1435 | 890.7 | 0.3574 | 0.0 | 3.5007 |
| EB0112 | 0.1860 | 0.2997 | 0.0588 | 1399.2 | 0.2670 | 0.0 | 1.6453 |

Appendix 2, continued

| Sample ID | Lu | Mn | Na | Nd | Ni | Rb | Sb |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0037 | 0.0367 | 88.37 | 414.5 | 0.9719 | 0.00 | 0.00 | 0.1518 |
| EB0038 | 0.0000 | 110.57 | 196.8 | 0.4344 | 0.00 | 0.00 | 0.0262 |
| EB0039 | 0.0221 | 20.14 | 164.6 | 1.1452 | 5.79 | 0.30 | 0.1836 |
| EB0040 | 0.0000 | 8.71 | 799.3 | 1.1090 | 4.70 | 0.00 | 0.0353 |
| EB0041 | 0.0211 | 18.85 | 445.4 | 0.0000 | 5.26 | 0.00 | 0.0769 |
| EB0042 | 0.0056 | 6.65 | 182.6 | 0.0000 | 2.71 | 0.00 | 0.0514 |
| EB0043 | 0.0069 | 25.84 | 459.0 | 0.5587 | 8.62 | 0.00 | 0.1201 |
| EB0044 | 0.0167 | 9.02 | 526.3 | 0.0000 | 0.00 | 0.00 | 0.0263 |
| EB0045 | 0.0146 | 13.02 | 192.8 | 0.0000 | 6.44 | 0.00 | 0.0518 |
| EB0046 | 0.0186 | 21.61 | 667.2 | 0.7584 | 5.51 | 0.00 | 0.1514 |
| EB0047 | 0.0056 | 13.60 | 240.9 | 0.4871 | 4.34 | 0.00 | 0.0867 |
| EB0048 | 0.0196 | 2.58 | 505.6 | 0.0000 | 0.00 | 0.00 | 0.0350 |
| EB0049 | 0.0047 | 7.03 | 1185.8 | 0.0000 | 0.00 | 0.00 | 0.0819 |
| EB0050 | 0.0250 | 5.01 | 427.3 | 0.0000 | 0.00 | 0.00 | 0.0084 |
| EB0051 | 0.0113 | 53.81 | 370.1 | 1.1584 | 4.61 | 0.00 | 0.1299 |
| EB0052 | 0.0000 | 230.38 | 824.6 | 0.5981 | 0.00 | 0.00 | 0.0429 |
| EB0053 | 0.0293 | 111.84 | 283.0 | 0.8905 | 0.00 | 0.00 | 0.0502 |
| EB0054 | 0.0000 | 52.27 | 349.0 | 0.7899 | 6.79 | 0.00 | 0.0259 |
| EB0055 | 0.0000 | 84.36 | 225.8 | 0.6269 | 4.58 | 0.00 | 0.0269 |
| EB0056 | 0.0000 | 125.54 | 337.1 | 0.0000 | 0.00 | 0.00 | 0.0095 |
| EB0057 | 0.0000 | 171.63 | 174.2 | 1.0679 | 0.00 | 0.00 | 0.0215 |
| EB0058 | 0.0000 | 98.07 | 510.0 | 1.7305 | 4.17 | 0.00 | 0.0473 |
| EB0059 | 0.0000 | 69.34 | 225.0 | 0.7260 | 0.00 | 0.00 | 0.0119 |
| EB0060 | 0.0000 | 102.36 | 470.0 | 0.9733 | 2.96 | 0.00 | 0.0184 |
| EB0061 | 0.0000 | 183.10 | 345.1 | 0.5279 | 3.31 | 0.00 | 0.0218 |
| EB0062 | 0.0000 | 116.35 | 348.3 | 0.0000 | 3.44 | 0.00 | 0.0089 |
| EB0063 | 0.0000 | 119.66 | 916.1 | 1.7987 | 0.00 | 0.00 | 0.0139 |
| EB0064 | 0.0195 | 46.50 | 401.0 | 0.6612 | 4.38 | 0.45 | 0.0351 |
| EB0065 | 0.0142 | 48.10 | 688.0 | 0.0000 | 3.48 | 0.00 | 0.0171 |
| EB0066 | 0.0060 | 5.86 | 434.3 | 1.0591 | 6.16 | 0.00 | 0.1768 |
| EB0067 | 0.0038 | 106.61 | 1694.4 | 0.0000 | 0.00 | 0.00 | 0.0580 |
| EB0068 | 0.0059 | 143.07 | 209.2 | 0.0000 | 3.00 | 0.00 | 0.0224 |
| EB0069 | 0.0000 | 119.04 | 348.6 | 0.0000 | 0.00 | 0.00 | 0.0256 |
| EB0070 | 0.0112 | 115.31 | 259.8 | 0.0000 | 2.78 | 0.00 | 0.0242 |
| EB0071 | 0.0000 | 62.84 | 508.5 | 0.0000 | 5.27 | 0.00 | 0.0178 |
| EB0072 | 0.0000 | 54.13 | 194.5 | 0.0000 | 6.06 | 0.21 | 0.0176 |
| EB0073 | 0.0000 | 51.77 | 384.9 | 0.0000 | 0.00 | 0.00 | 0.0177 |
| EB0074 | 0.0000 | 40.73 | 233.2 | 0.0000 | 2.87 | 0.00 | 0.0201 |
| EB0075 | 0.0000 | 85.78 | 257.8 | 0.6253 | 3.17 | 0.00 | 0.0374 |
| EB0076 | 0.0102 | 143.97 | 267.3 | 0.0000 | 4.36 | 0.54 | 0.0299 |
| EB0078 | 0.0471 | 12.85 | 9085.6 | 1.5552 | 0.00 | 2.60 | 0.0541 |
| EB0079 | 0.0294 | 14.38 | 10481.0 | 1.3701 | 0.00 | 2.25 | 0.0242 |
| EB0080 | 0.0286 | 13.01 | 1666.1 | 1.8233 | 0.00 | 2.32 | 0.0140 |
| EB0081 | 0.1435 | 17.42 | 4907.0 | 6.2906 | 8.43 | 7.23 | 0.0922 |
| EB0082 | 0.0609 | 13.82 | 1277.5 | 2.2272 | 7.58 | 3.49 | 0.0887 |

Appendix 2, continued

| Sample ID | Lu | Mn | Na | Nd | Ni | Rb | Sb |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0083 | 0.0641 | 23.85 | 2245.0 | 1.9171 | 0.00 | 3.67 | 0.1413 |
| EB0084 | 0.0657 | 12.77 | 22297.5 | 2.8973 | 0.00 | 4.31 | 0.0837 |
| EB0085 | 0.0365 | 16.72 | 383.5 | 1.4651 | 0.00 | 0.69 | 0.0713 |
| EB0086 | 0.0000 | 9.08 | 1001.9 | 0.0000 | 0.00 | 0.00 | 0.0161 |
| EB0087 | 0.0000 | 211.77 | 214.4 | 0.0000 | 3.66 | 0.00 | 0.0282 |
| EB0088 | 0.0000 | 47.08 | 184.2 | 0.0000 | 2.74 | 0.00 | 0.0256 |
| EB0089 | 0.0000 | 67.18 | 566.4 | 0.0000 | 2.67 | 0.12 | 0.0444 |
| EB0090 | 0.0000 | 87.79 | 630.0 | 0.0000 | 4.15 | 0.00 | 0.0355 |
| EB0091 | 0.0000 | 142.22 | 216.7 | 0.0000 | 6.17 | 0.00 | 0.0323 |
| EB0092 | 0.0000 | 42.42 | 198.0 | 0.0000 | 3.81 | 0.00 | 0.0154 |
| EB0093 | 0.0113 | 69.72 | 435.3 | 0.0000 | 0.00 | 0.54 | 0.0170 |
| EB0094 | 0.0000 | 40.55 | 273.4 | 0.0000 | 3.51 | 0.00 | 0.0117 |
| EB0095 | 0.0000 | 31.94 | 172.0 | 0.0000 | 2.43 | 0.00 | 0.0116 |
| EB0096 | 0.0085 | 28.10 | 352.3 | 0.0000 | 2.51 | 0.00 | 0.0181 |
| EB0097 | 0.0000 | 17.47 | 294.9 | 0.0000 | 4.21 | 0.00 | 0.0150 |
| EB0098 | 0.0000 | 36.56 | 219.8 | 0.0000 | 2.57 | 0.00 | 0.0086 |
| EB0099 | 0.0000 | 55.95 | 214.4 | 0.0000 | 2.89 | 0.00 | 0.0137 |
| EB0100 | 0.1386 | 27.65 | 1741.2 | 7.1610 | 20.08 | 6.36 | 0.1175 |
| EB0101 | 0.0780 | 12.49 | 2175.2 | 3.4926 | 5.06 | 5.83 | 0.0695 |
| EB0102 | 0.0387 | 11.31 | 872.7 | 1.7301 | 6.16 | 2.54 | 0.0325 |
| EB0103 | 0.0401 | 17.25 | 856.9 | 2.7155 | 10.01 | 2.69 | 0.0989 |
| EB0104 | 0.0497 | 58.04 | 742.0 | 2.6825 | 0.00 | 2.31 | 3.1663 |
| EB0105 | 0.0852 | 23.78 | 1089.3 | 4.7546 | 10.29 | 3.51 | 0.0927 |
| EB0106 | 0.1020 | 22.80 | 1597.9 | 4.2258 | 12.13 | 3.77 | 0.0755 |
| EB0107 | 0.0760 | 10.76 | 1263.6 | 2.5174 | 7.78 | 3.05 | 0.4289 |
| EB0108 | 0.0569 | 17.02 | 1161.7 | 2.2573 | 5.94 | 3.02 | 0.0465 |
| EB0109 | 0.1102 | 21.91 | 1243.8 | 5.7588 | 4.09 | 3.68 | 0.1427 |
| EB0110 | 0.1275 | 25.74 | 1872.7 | 5.5558 | 16.01 | 5.21 | 0.2753 |
| EB0111 | 0.1110 | 20.56 | 1592.2 | 4.4125 | 0.00 | 4.66 | 0.0331 |
| EB0112 | 0.0414 | 13.04 | 740.7 | 1.3530 | 0.00 | 2.70 | 0.0889 |

Appendix 2, continued

| Sample ID | Sc | Sm | Sr | Ta | Tb | Th | Ti |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0037 | 0.0610 | 0.3051 | 272.59 | 0.0220 | 0.0246 | 0.1058 | 0.0 |
| EB0038 | 0.0315 | 0.1152 | 238.68 | 0.0000 | 0.0000 | 0.0150 | 0.0 |
| EB0039 | 0.1874 | 0.2716 | 78.81 | 0.0087 | 0.0307 | 0.0791 | 0.0 |
| EB0040 | 0.0507 | 0.2094 | 202.08 | 0.0000 | 0.0106 | 0.0290 | 0.0 |
| EB0041 | 0.0466 | 0.1471 | 199.62 | 0.0000 | 0.0100 | 0.0359 | 0.0 |
| EB0042 | 0.0656 | 0.1171 | 180.91 | 0.0000 | 0.0104 | 0.0420 | 0.0 |
| EB0043 | 0.0654 | 0.1458 | 169.44 | 0.0035 | 0.0154 | 0.0526 | 0.0 |
| EB0044 | 0.0476 | 0.0880 | 220.67 | 0.0000 | 0.0108 | 0.0297 | 0.0 |
| EB0045 | 0.0473 | 0.1049 | 162.60 | 0.0000 | 0.0081 | 0.0347 | 0.0 |
| EB0046 | 0.0561 | 0.1519 | 146.23 | 0.0000 | 0.0198 | 0.0700 | 0.0 |
| EB0047 | 0.0450 | 0.1087 | 221.70 | 0.0000 | 0.0127 | 0.0301 | 0.0 |
| EB0048 | 0.0710 | 0.1312 | 236.84 | 0.0000 | 0.0096 | 0.0350 | 0.0 |
| EB0049 | 0.0403 | 0.1176 | 161.86 | 0.0000 | 0.0085 | 0.0289 | 0.0 |
| EB0050 | 0.0589 | 0.1457 | 399.86 | 0.0000 | 0.0043 | 0.0326 | 0.0 |
| EB0051 | 0.0599 | 0.2876 | 239.43 | 0.0000 | 0.0239 | 0.0971 | 0.0 |
| EB0052 | 0.0243 | 0.2591 | 306.30 | 0.0000 | 0.0000 | 0.0190 | 0.0 |
| EB0053 | 0.1134 | 0.2573 | 245.59 | 0.0102 | 0.0148 | 0.1140 | 0.0 |
| EB0054 | 0.0449 | 0.3391 | 271.78 | 0.0000 | 0.0048 | 0.0281 | 0.0 |
| EB0055 | 0.0165 | 0.1356 | 237.23 | 0.0000 | 0.0000 | 0.0114 | 0.0 |
| EB0056 | 0.0163 | 0.2162 | 228.44 | 0.0000 | 0.0000 | 0.0080 | 0.0 |
| EB0057 | 0.0267 | 0.2643 | 262.90 | 0.0000 | 0.0000 | 0.0143 | 0.0 |
| EB0058 | 0.0190 | 0.6130 | 245.28 | 0.0000 | 0.0000 | 0.0091 | 0.0 |
| EB0059 | 0.0215 | 0.2325 | 214.71 | 0.0000 | 0.0000 | 0.0133 | 111.3 |
| EB0060 | 0.0412 | 0.2795 | 159.59 | 0.0000 | 0.0000 | 0.0125 | 0.0 |
| EB0061 | 0.0202 | 0.1681 | 186.18 | 0.0000 | 0.0000 | 0.0113 | 0.0 |
| EB0062 | 0.0148 | 0.1837 | 210.82 | 0.0000 | 0.0000 | 0.0111 | 0.0 |
| EB0063 | 0.0342 | 0.2880 | 269.57 | 0.0000 | 0.0000 | 0.0139 | 0.0 |
| EB0064 | 0.0936 | 0.1340 | 308.09 | 0.0076 | 0.0090 | 0.0619 | 0.0 |
| EB0065 | 0.0572 | 0.0967 | 301.68 | 0.0000 | 0.0057 | 0.0351 | 0.0 |
| EB0066 | 0.1034 | 0.1839 | 326.05 | 0.0023 | 0.0106 | 0.0622 | 0.0 |
| EB0067 | 0.0589 | 0.1380 | 271.71 | 0.0051 | 0.0089 | 0.0545 | 0.0 |
| EB0068 | 0.0422 | 0.0987 | 225.04 | 0.0000 | 0.0084 | 0.0183 | 14.4 |
| EB0069 | 0.0160 | 0.0444 | 269.23 | 0.0000 | 0.0000 | 0.0112 | 0.0 |
| EB0070 | 0.0318 | 0.0925 | 389.34 | 0.0000 | 0.0000 | 0.0253 | 0.0 |
| EB0071 | 0.0362 | 0.0723 | 263.18 | 0.0000 | 0.0000 | 0.0355 | 0.0 |
| EB0072 | 0.0373 | 0.0817 | 258.96 | 0.0000 | 0.0058 | 0.0337 | 0.0 |
| EB0073 | 0.0341 | 0.0814 | 298.44 | 0.0000 | 0.0000 | 0.0326 | 0.0 |
| EB0074 | 0.0297 | 0.0963 | 230.94 | 0.0000 | 0.0000 | 0.0293 | 0.0 |
| EB0075 | 0.0301 | 0.0893 | 316.48 | 0.0032 | 0.0000 | 0.0256 | 0.0 |
| EB0076 | 0.0703 | 0.1389 | 140.98 | 0.0000 | 0.0048 | 0.0505 | 0.0 |
| EB0078 | 0.4055 | 0.3886 | 1980.76 | 0.0493 | 0.0447 | 0.2832 | 0.0 |
| EB0079 | 0.3402 | 0.3084 | 2437.61 | 0.0440 | 0.0237 | 0.2368 | 0.0 |
| EB0080 | 0.4368 | 0.4633 | 2493.54 | 0.0435 | 0.0518 | 0.2550 | 0.0 |
| EB0081 | 1.7643 | 1.4085 | 2461.96 | 0.1793 | 0.1554 | 1.0645 | 575.1 |
| EB0082 | 0.5384 | 0.6099 | 2513.42 | 0.0637 | 0.0506 | 0.3835 | 0.0 |

Appendix 2, continued

| Sample ID | Sc | Sm | Sr | Ta | Tb | Th | Ti |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EB0083 | 0.5711 | 0.5641 | 2345.70 | 0.0691 | 0.0673 | 0.3911 | 193.4 |
| EB0084 | 0.8708 | 0.7401 | 2257.66 | 0.0983 | 0.0777 | 0.5863 | 271.4 |
| EB0085 | 0.2183 | 0.4103 | 752.81 | 0.0330 | 0.0462 | 0.2683 | 0.0 |
| EB0086 | 0.1198 | 0.1905 | 2729.96 | 0.0165 | 0.0254 | 0.1169 | 0.0 |
| EB0087 | 0.0454 | 0.1263 | 178.97 | 0.0000 | 0.0000 | 0.0310 | 0.0 |
| EB0088 | 0.0327 | 0.1094 | 341.47 | 0.0000 | 0.0067 | 0.0282 | 0.0 |
| EB0089 | 0.0197 | 0.0882 | 297.54 | 0.0000 | 0.0000 | 0.0164 | 0.0 |
| EB0090 | 0.0423 | 0.1651 | 385.64 | 0.0000 | 0.0000 | 0.0317 | 0.0 |
| EB0091 | 0.0242 | 0.0908 | 342.00 | 0.0000 | 0.0000 | 0.0173 | 0.0 |
| EB0092 | 0.0190 | 0.0473 | 196.53 | 0.0000 | 0.0000 | 0.0081 | 0.0 |
| EB0093 | 0.0234 | 0.0778 | 208.92 | 0.0000 | 0.0000 | 0.0136 | 0.0 |
| EB0094 | 0.0371 | 0.0627 | 240.21 | 0.0000 | 0.0000 | 0.0242 | 0.0 |
| EB0095 | 0.0343 | 0.0656 | 235.72 | 0.0000 | 0.0048 | 0.0212 | 0.0 |
| EB0096 | 0.0210 | 0.0767 | 290.71 | 0.0000 | 0.0041 | 0.0147 | 0.0 |
| EB0097 | 0.0292 | 0.1395 | 311.93 | 0.0000 | 0.0000 | 0.0256 | 0.0 |
| EB0098 | 0.0250 | 0.1103 | 287.33 | 0.0000 | 0.0000 | 0.0152 | 0.0 |
| EB0099 | 0.0203 | 0.0654 | 276.46 | 0.0031 | 0.0000 | 0.0159 | 0.0 |
| EB0100 | 1.4699 | 1.4844 | 2063.63 | 0.1354 | 0.1547 | 1.1618 | 305.6 |
| EB0101 | 0.8876 | 0.7977 | 1943.60 | 0.1184 | 0.0789 | 0.6997 | 285.4 |
| EB0102 | 0.3647 | 0.4024 | 2239.04 | 0.0460 | 0.0361 | 0.2688 | 0.0 |
| EB0103 | 0.4495 | 0.4394 | 1658.31 | 0.0516 | 0.0414 | 0.3067 | 140.0 |
| EB0104 | 0.3991 | 0.4684 | 1789.00 | 0.0526 | 0.0472 | 0.2549 | 0.0 |
| EB0105 | 0.6849 | 0.8617 | 2037.57 | 0.0734 | 0.0759 | 0.4951 | 266.9 |
| EB0106 | 0.7767 | 1.0326 | 1691.17 | 0.0857 | 0.0898 | 0.6694 | 193.7 |
| EB0107 | 0.4919 | 0.7098 | 2091.72 | 0.0565 | 0.0383 | 0.3575 | 221.4 |
| EB0108 | 0.4609 | 0.6033 | 2267.89 | 0.0480 | 0.0466 | 0.3532 | 163.8 |
| EB0109 | 1.0176 | 1.2870 | 1891.93 | 0.0935 | 0.1227 | 0.8044 | 323.0 |
| EB0110 | 0.9605 | 1.2135 | 2743.73 | 0.1025 | 0.0931 | 0.6942 | 368.7 |
| EB0111 | 0.9606 | 1.1650 | 2118.64 | 0.1013 | 0.0911 | 0.6620 | 160.1 |
| EB0112 | 0.4632 | 0.4645 | 2406.64 | 0.0570 | 0.0351 | 0.3336 | 167.8 |

Appendix 2, continued

| Sample ID | $\mathbf{U}$ | $\mathbf{V}$ | $\mathbf{Y b}$ | $\mathbf{Z n}$ | $\mathbf{Z r}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| EB0037 | 1.7877 | 7.16 | 0.0900 | 19.67 | 12.59 |
| EB0038 | 1.0477 | 7.93 | 0.0163 | 10.94 | 8.20 |
| EB0039 | 1.0868 | 12.02 | 0.1936 | 16.79 | 9.83 |
| EB0040 | 1.6712 | 4.82 | 0.0582 | 12.23 | 8.79 |
| EB0041 | 1.0906 | 8.32 | 0.0480 | 16.35 | 11.20 |
| EB0042 | 0.7391 | 4.65 | 0.0494 | 15.11 | 7.84 |
| EB0043 | 1.0220 | 7.91 | 0.0766 | 21.69 | 7.33 |
| EB0044 | 0.5983 | 3.29 | 0.0445 | 18.39 | 6.71 |
| EB0045 | 0.7218 | 5.91 | 0.0620 | 15.89 | 5.66 |
| EB0046 | 0.7522 | 1.74 | 0.0566 | 18.78 | 6.01 |
| EB0047 | 0.7935 | 5.43 | 0.0606 | 14.85 | 6.96 |
| EB0048 | 0.8378 | 0.00 | 0.0651 | 8.79 | 6.29 |
| EB0049 | 1.0684 | 8.05 | 0.0304 | 8.36 | 10.97 |
| EB0050 | 1.1568 | 6.83 | 0.0388 | 9.64 | 7.04 |
| EB0051 | 1.8814 | 5.70 | 0.1092 | 14.13 | 12.72 |
| EB0052 | 2.7909 | 18.67 | 0.0213 | 7.58 | 18.05 |
| EB0053 | 2.0703 | 16.03 | 0.0392 | 10.27 | 15.40 |
| EB0054 | 3.6179 | 11.90 | 0.0444 | 8.70 | 26.38 |
| EB0055 | 1.5230 | 6.97 | 0.0000 | 10.24 | 11.30 |
| EB0056 | 2.3671 | 7.98 | 0.0117 | 9.40 | 16.75 |
| EB0057 | 2.7721 | 9.31 | 0.0515 | 12.74 | 19.99 |
| EB0058 | 7.2064 | 16.43 | 0.0217 | 8.43 | 47.41 |
| EB0059 | 2.5222 | 9.58 | 0.0270 | 9.22 | 16.64 |
| EB0060 | 2.9984 | 9.56 | 0.0584 | 5.36 | 20.74 |
| EB0061 | 1.5616 | 8.16 | 0.0174 | 12.31 | 12.26 |
| EB0062 | 1.8156 | 7.01 | 0.0208 | 9.36 | 13.97 |
| EB0063 | 3.0818 | 13.12 | 0.0335 | 9.11 | 21.15 |
| EB0064 | 0.9466 | 1.25 | 0.0391 | 13.55 | 11.02 |
| EB065 | 0.7539 | 2.97 | 0.0282 | 9.78 | 4.75 |
| EB0066 | 1.2604 | 8.43 | 0.0399 | 16.20 | 10.74 |
| EB0067 | 1.0316 | 9.20 | 0.0344 | 3.85 | 11.63 |
| EB0068 | 0.6457 | 11.39 | 0.0387 | 2.58 | 6.94 |
| EB0069 | 0.4443 | 0.00 | 0.0000 | 3.04 | 2.78 |
| EB0070 | 0.8544 | 0.00 | 0.0000 | 2.39 | 5.58 |
| EB0071 | 0.6297 | 0.00 | 0.0142 | 2.05 | 3.99 |
| EB0072 | 0.6919 | 0.00 | 0.0171 | 2.46 | 6.17 |
| EB0073 | 0.8866 | 6.74 | 0.0000 | 2.21 | 8.39 |
| EB0074 | 0.6735 | 5.07 | 0.0000 | 1.76 | 5.90 |
| EB0075 | 0.8543 | 5.08 | 0.0000 | 1.49 | 5.79 |
| EB0076 | 0.8665 | 9.47 | 0.0195 | 1.87 | 7.22 |
| EB0078 | 1.4831 | 0.00 | 0.2156 | 4.69 | 15.50 |
| EB0079 | 1.1536 | 0.00 | 0.1086 | 5.61 | 14.16 |
| EB0080 | 1.3919 | 8.56 | 0.1963 | 4.84 | 18.11 |
| EB0081 | 3.0537 | 13.14 | 0.5444 | 10.10 | 36.14 |
| EB0082 | 2.3058 | 9.47 | 0.1909 | 8.60 | 18.88 |
|  |  |  |  |  |  |

Appendix 2, continued

| Sample ID | $\mathbf{U}$ | $\mathbf{V}$ | $\mathbf{Y b}$ | $\mathbf{Z n}$ | $\mathbf{Z r}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| EB0083 | 1.8923 | 10.36 | 0.2054 | 11.93 | 17.29 |
| EB0084 | 2.7780 | 0.00 | 0.2932 | 7.03 | 19.96 |
| EB0085 | 1.1819 | 0.00 | 0.1476 | 2.30 | 15.02 |
| EB0086 | 0.8936 | 9.98 | 0.0910 | 3.35 | 7.33 |
| EB0087 | 1.1258 | 0.00 | 0.0203 | 1.63 | 6.94 |
| EB0088 | 1.0052 | 9.27 | 0.0150 | 3.42 | 4.67 |
| EB0089 | 0.8651 | 12.15 | 0.0000 | 2.60 | 4.00 |
| EB0090 | 1.5668 | 8.39 | 0.0000 | 2.15 | 9.93 |
| EB0091 | 0.9695 | 4.85 | 0.0000 | 2.72 | 5.49 |
| EB0092 | 0.4397 | 0.00 | 0.0129 | 5.30 | 2.57 |
| EB0093 | 0.7411 | 0.00 | 0.0246 | 6.69 | 3.98 |
| EB0094 | 0.4001 | 0.00 | 0.0118 | 6.49 | 2.22 |
| EB0095 | 0.5033 | 3.01 | 0.0201 | 6.20 | 0.00 |
| EB0096 | 0.6374 | 0.00 | 0.0339 | 8.07 | 0.00 |
| EB0097 | 1.4795 | 0.00 | 0.0000 | 8.06 | 7.78 |
| EB0098 | 0.8921 | 0.00 | 0.0157 | 6.34 | 5.63 |
| EB0099 | 0.5902 | 0.00 | 0.0302 | 6.15 | 4.22 |
| EB0100 | 3.4031 | 21.58 | 0.5030 | 16.33 | 36.63 |
| EB0101 | 2.8603 | 14.93 | 0.2291 | 7.03 | 29.49 |
| EB0102 | 1.8627 | 8.79 | 0.1108 | 5.08 | 11.89 |
| EB0103 | 1.6503 | 12.09 | 0.1593 | 6.06 | 11.10 |
| EB0104 | 2.2385 | 8.06 | 0.1427 | 2.42 | 15.33 |
| EB0105 | 3.5176 | 18.63 | 0.2908 | 9.96 | 26.24 |
| EB0106 | 3.8345 | 15.57 | 0.2891 | 10.63 | 30.30 |
| EB0107 | 3.9895 | 20.08 | 0.1727 | 6.68 | 26.36 |
| EB0108 | 2.6062 | 10.27 | 0.1672 | 5.54 | 18.36 |
| EB0109 | 4.0877 | 25.12 | 0.3929 | 7.53 | 29.71 |
| EB0110 | 5.2673 | 25.44 | 0.3436 | 11.71 | 35.32 |
| EB0111 | 4.4226 | 18.03 | 0.3417 | 5.46 | 31.74 |
| EB0112 | 1.7936 | 13.84 | 0.1341 | 5.37 | 23.07 |

Appendix 3 - Limestone source information for the Harrell Egyptian limestone samples.

| ID | Formation | Location | Coordinates |  |
| :---: | :---: | :---: | :---: | :---: |
|  |  |  | ${ }^{\circ} \mathrm{N}$ | ${ }^{\circ} \mathbf{E}$ |
| EL-01 | Alexandria | Abu Sir | 30.9467 | 29.5000 |
| EL-02 | Alexandria | Abu Sir | 30.9467 | 29.5000 |
| EL-03 | Alexandria | Mex village | 31.1542 | 29.8433 |
| EL-04 | Mokattam | Zawyet on Gebel Mokattam near Citadel | 30.0267 | 31.2700 |
| EL-05 | Mokattam | Gebel Tura near Tura village | 29.9333 | 31.2987 |
| EL-06 | Mokattam | Gebel Hof near el-Masara village | 29.9150 | 31.3200 |
| EL-07 | Mokattam | Gebel Hof near el-Masara village | 29.9150 | 31.3200 |
| EL-08 | Mokattam | Wadi Abu Mu'aymil near St. Antony Monastery | 28.8983 | 32.3250 |
| EL-09 | Mokattam | Wadi Abu Mu'aymil near St. Antony Monastery | 28.8983 | 32.3250 |
| EL-10 | Mokattam | Wadi Abu Mu'aymil near St. Antony Monastery | 28.8983 | 32.3250 |
| EL-11 | Mokattam | Wadi Abu Mu'aymil near St. Antony Monastery | 28.8983 | 32.3250 |
| EL-12 | Mokattam | Wadi Umm Zanatir near St. Antony Monastery | 28.9383 | 32.3950 |
| EL-13 | Mokattam | Wadi Umm Zanatir near St. Antony Monastery | 28.9383 | 32.3950 |
| EL-14 | Samalut | near el-Sawayta village | 28.3768 | 30.8010 |
| EL-15 | Samalut | el Babein tomb near Beni Khalid village | 28.3047 | 30.7507 |
| EL-16 | Samalut | Zawyet el-Amwat village in Zawyet Sultan district | 28.0550 | 30.8317 |
| EL-17 | Minia | Beni Hasan tombs | 27.9107 | 30.8717 |
| EL-18 | Minia | Beni Hasan tombs | 27.9107 | 30.8717 |
| EL-19 | Minia | el-Sheikh Timay village | 27.8617 | 30.8453 |
| EL-20 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-21 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-22 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-23 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-24 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-25 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-26 | Minia | Wadi el-Nakla near Deir el-Bersha village | 27.7512 | 30.9193 |
| EL-27 | Minia | el-Bersha village on Gebel Sheikh Said | 27.7207 | 30.8945 |
| EL-28 | Minia | eastern Wadi el-Zebeida (Queen Tiy Quarry) | 27.6835 | 30.9022 |
| EL-29 | Minia | eastern Wadi el-Zebeida (Queen Tiy Quarry) | 27.6835 | 30.9022 |
| EL-30 | Minia | eastern Wadi el-Zebeida (Queen Tiy Quarry) | 27.6835 | 30.9022 |
| EL-31 | Minia | eastern Wadi el-Zebeida (Queen Tiy Quarry) | 27.6835 | 30.9022 |
| EL-32 | Minia | central Wadi el-Zebeida (Abd el-Azziz Quarry) | 27.6895 | 30.9058 |
| EL-33 | Minia | central Wadi el-Zebeida (Abd el-Azziz Quarry) | 27.6895 | 30.9058 |
| EL-34 | Minia | western Wadi el-Zebeida | 27.6928 | 30.9017 |
| EL-35 | Minia | near Sheikh Said tomb on Gebel Sheikh Said | 27.6997 | 30.8890 |
| EL-36 | Minia | near Sheikh Said tomb on Gebel Sheikh Said | 27.6997 | 30.8890 |
| EL-37 | Minia | near Sheikh Said tomb on Gebel Sheikh Said | 27.6997 | 30.8890 |
| EL-38 | Minia | near Sheikh Said tomb on Gebel Sheikh Said | 27.6997 | 30.8890 |
| EL-39 | Minia | Northern Tombs at Amarna ruins | 27.6620 | 30.9297 |
| EL-40 | Minia | Northern Tombs at Amarna ruins | 27.6620 | 30.9297 |
| EL-41 | Minia | el-Maabda village on Gebel el-Harrana | 27.3453 | 31.0283 |
| EL-42 | Minia | Deir el-Gabrawi village on Gebel el-Tawila | 27.3388 | 31.1032 |
| EL-43 | Minia | Arab el-Atiat el-Bahariya village on Gebel el-Harrana | 27.3345 | 31.0662 |
| EL-44 | Drunka | el-Izam monastery near Assiut city | 27.1542 | 31.1483 |
| EL-45 | Drunka | Wadi Emu | 27.1195 | 31.3568 |

Appendix 3 - Limestone source information for the Harrell Egyptian limestone samples, continued.

| ID | Formation | Location | Coordinates |  |
| :---: | :---: | :--- | :---: | :---: |
|  |  |  | $\mathbf{N}$ | $\mathbf{E}$ |
| EL-46 | Drunka | el-Khawalid village | 27.0935 | 31.3870 |
| EL-47 | Drunka | Qaw el-Kebir/Antaeopolis ruins | 26.9218 | 31.5002 |
| EL-48 | Drunka | Qaw el-Kebir/Antaeopolis ruins | 26.9218 | 31.5002 |
| EL-49 | Drunka | Nazlet el-Haridi village on Gebel el-Haridi | 26.7772 | 31.5518 |
| EL-50 | Drunka | el-Salamuni village | 26.6178 | 31.7642 |
| EL-51 | Drunka | Nag Hamad village and Athribis ruins | 26.5093 | 31.6627 |
| EL-52 | Drunka | Nag Hamad village and Athribis ruins | 26.5093 | 31.6627 |
| EL-53 | Drunka | el-Salmuni village and Abydos ruins (mostly <br> destroyed) | 26.2042 | 31.8758 |
| EL-54 | Drunka | Wadi Naqb el-Salmuni near Abydos ruins | 26.1935 | 31.8658 |
| EL-55 | Drunka | Wadi Naqb el-Salmuni near Abydos ruins | 26.1935 | 31.8658 |
| EL-56 | Serai | Nag el-Buza village | 26.0950 | 32.3000 |
| EL-57 | Issawia | Gebel el-Gir near Tentyris/Dendara ruins | 26.1045 | 32.6950 |
| EL-58 | Serai | Wadi el-Muluk (Valley of Kings) and Qurna | 25.7463 | 32.6225 |
| EL-59 | Serai | Wadi el-Muluk (Valley of Kings) and Qurna | 25.7463 | 32.6225 |
| EL-60 | Serai | el-Ghrera village in el-Gebelein district (now <br> destroyed) | 25.4977 | 32.4790 |
| EL-61 | Tarawan | el-Dibabiya village | 25.5013 | 32.5183 |
| EL-62 | Tarawan | el-Dibabiya village | 25.5013 | 32.5183 |

Appendix 4 - Concentration data from NAA for limestone from Egypt (Harrell samples). All concentrations are in parts per million (ppm)

| Sample ID | Al | As | Ba | Ca | Ce | Co | Cr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL01 | 723.4373 | 12.082 | 38.73553 | 380778.8 | 3.3874 | 0.2964 | 2.3437 |
| EL02 | 1632.057 | 15.0009 | 23.28542 | 382987.3 | 3.7104 | 0.3575 | 2.6266 |
| EL03 | 0 | 2.0288 | 17.90644 | 375440.6 | 2.0614 | 0.1109 | 1.8339 |
| EL04 | 9204.907 | 5.2011 | 24.72878 | 354192.8 | 8.9902 | 0.5688 | 30.9747 |
| EL05 | 5482.053 | 1.8776 | 43.52981 | 343062.6 | 4.8384 | 0.2821 | 23.3047 |
| EL06 | 0 | 0.9003 | 91.05511 | 276624.6 | 0.8977 | 0.0829 | 9.5759 |
| EL07 | 0 | 0.28 | 13.51148 | 384113.3 | 1.1704 | 0.0772 | 12.0844 |
| EL08 | 4191.739 | 0.3489 | 40.35509 | 379221.8 | 3.8859 | 0.2349 | 23.2527 |
| EL09 | 3794.588 | 0.8741 | 17.77534 | 371380.5 | 3.8971 | 0.2453 | 18.8732 |
| EL10 | 2060.192 | 0.8426 | 10.53595 | 383732.3 | 2.6199 | 0.1029 | 14.8273 |
| EL11 | 6185.242 | 2.2953 | 25.413 | 363021.1 | 6.8083 | 0.6473 | 30.6139 |
| EL12 | 3507.994 | 1.074 | 0 | 231764.8 | 1.3711 | 0.5827 | 6.0781 |
| EL13 | 4906.892 | 1.2019 | 0 | 252326.8 | 1.2592 | 1.1459 | 11.0204 |
| EL14 | 4828.893 |  | 8.03519 | 398464 | 0.4031 | 0.0665 | 6.3108 |
| EL15 | 6041.666 | 0 | 0 | 401336.9 | 0.4738 | 0.0265 | 6.9172 |
| EL16 | 4697.163 | 0 | 0 | 408532.9 | 0.3962 | 0.0471 | 4.1542 |
| EL17 | 10260.58 | 0.113 | 16.98495 | 400398 | 0.7205 | 0.0777 | 7.5618 |
| EL18 | 3300.669 | 0.2921 | 19.82584 | 401567.8 | 0.7636 | 0.2297 | 6.7607 |
| EL19 | 1828.264 | 0.4307 | 0 | 398113.3 | 1.5801 | 0.1906 | 10.1975 |
| EL20 | 1420.299 | 0.227 | 0 | 399441.7 | 0.515 | 0.1199 | 7.8936 |
| EL21 | 3463.952 | 0 | 0 | 401099.4 | 0.4809 | 0.0373 | 11.4205 |
| EL22 | 508.7635 | 0.3126 | 0 | 409365.5 | 0.4122 | 0.0268 | 9.5965 |
| EL23 | 0 | 0.1877 | 0 | 405182.7 | 0.4379 | 0.0236 | 15.8539 |
| EL24 | 665.5685 | 0 | 0 | 400512.3 | 0.5215 | 0.0488 | 8.8138 |
| EL25 | 0 | 0 | 0 | 396609.8 | 0.476 | 0.0421 | 12.2639 |
| EL26 | 0 | 0.298 | 8.87214 | 405606.4 | 0.4892 | 0.0816 | 9.3463 |
| EL27 | 0 | 0.5053 | 162.0849 | 396315.5 | 0.3118 | 0.2907 | 4.8339 |
| EL28 | 1350.58 | 0 | 0 | 401009.1 | 0.7505 | 0.0707 | 9.4043 |
| EL29 | 0 | 0 | 0 | 405085.7 | 0.4243 | 0.062 | 6.699 |
| EL30 | 588.2661 | 0 | 0 | 395176.3 | 0.307 | 0.0314 | 5.9396 |
| EL31 | 4831.842 | 0 | 0 | 404898.5 | 0.4354 | 0.0636 | 7.1738 |
| EL32 | 0 | 0 | 0 | 402996.8 | 0.2404 | 0.1445 | 7.14 |
| EL33 | 0 | 0 | 0 | 395812.5 | 0.4202 | 0.1556 | 4.9878 |
| EL34 | 0 | 0 | 0 | 395425.9 | 0.2571 | 0.0211 | 5.5698 |
| EL35 | 0 | 0.5831 | 21.56146 | 389214.6 | 1.4283 | 0.0824 | 7.9978 |
| EL36 | 0 | 0 | 7.70596 | 392607.5 | 0.6953 | 0.0728 | 5.3599 |
| EL37 | 1657.535 | 0.4273 | 0 | 359715.3 | 1.8887 | 0.2458 | 38.1255 |
| EL38 | 0 | 0 | 10.75944 | 386910 | 0.4404 | 0.073 | 6.5269 |
| EL39 | 0 | 0 | 0 | 402256.8 | 0.4737 | 0.1087 | 5.138 |
| EL40 | 0 | 0 | 0 | 394662.6 | 0.4383 | 0.0522 | 5.6434 |
| EL41 | 0 | 0 | 0 | 393214.5 | 0.2793 | 0.1534 | 5.0608 |
| EL42 | 0 | 0 | 13.65086 | 396275.2 | 0.8084 | 0.1809 | 9.4878 |
| EL43 | 602.6922 | 0 | 13.03092 | 398985.5 | 0.5134 | 0.1955 | 2.3445 |
| EL44 | 1453.334 | 0 | 13.34118 | 391381.8 | 0.271 | 0.1364 | 4.6014 |
| EL45 |  |  |  |  |  |  |  |

Appendix 4, continued

| Sample ID | Al | As | Ba | Ca | Ce | Co | Cr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL46 | 3854.677 | 0 | 0 | 395907.8 | 0.6739 | 0.1316 | 9.2355 |
| EL47 | 2219.237 | 0 | 0 | 399337.4 | 0.3482 | 0.1105 | 5.6537 |
| EL48 | 5256.07 | 0 | 0 | 402816.7 | 0.5301 | 0.1825 | 4.1893 |
| EL49 | 2388.83 | 0 | 15.02989 | 411851 | 0.2832 | 0.1311 | 5.5741 |
| EL50 | 1222.226 | 0 | 0 | 406193.9 | 0.2164 | 0.125 | 8.285 |
| EL51 | 0 | 0.307 | 44.19081 | 394167.1 | 0.2921 | 0.1201 | 4.0962 |
| EL52 | 0 | 0.2869 | 39.12222 | 408888.1 | 0.5962 | 0.1748 | 7.5556 |
| EL53 | 1079.175 | 0 | 27.67321 | 400048.8 | 1.052 | 0.1079 | 6.6188 |
| EL54 | 0 | 0.1522 | 15.13041 | 403665.5 | 1.3008 | 0.0707 | 10.8048 |
| EL55 | 0 | 0.4066 | 17.71337 | 400098 | 2.1467 | 0.3903 | 16.0166 |
| EL56 | 1800.865 | 0.4217 | 31.1739 | 360483.9 | 1.8718 | 0.1406 | 33.8924 |
| EL57 | 6748.384 | 0.9706 | 40.93084 | 239055.3 | 9.2895 | 1.9224 | 71.8134 |
| EL58 | 8046.468 | 1.8536 | 22.032 | 306716.1 | 10.8423 | 2.6984 | 44.0464 |
| EL59 | 10069.42 | 1.2169 | 52.99543 | 276179.5 | 11.4837 | 1.8664 | 69.4308 |
| EL60 | 990.9293 | 0.4342 | 17.80272 | 330374.4 | 1.9206 | 0.3544 | 44.7385 |
| EL61 | 8826.216 | 1.2488 | 158.2274 | 274685.9 | 15.0968 | 4.5816 | 55.134 |
| EL62 | 10485.34 | 1.2516 | 327.1948 | 270151.5 | 18.9766 | 5.0219 | 57.6346 |

Appendix 4, continued

| Sample ID | Cs | Dy | Eu | Fe | Hf | K | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL01 | 0.0374 | 0.5396 | 0.0779 | 1450.398 | 0.1835 | 0 | 1.3923 |
| EL02 | 0.0398 | 0.3387 | 0.0942 | 1798.032 | 0.0979 | 0 | 1.6877 |
| EL03 | 0.0257 | 0.3751 | 0.0364 | 400.3628 | 0.1322 | 0 | 0.776 |
| EL04 | 0.5289 | 0.4956 | 0.1637 | 4948.203 | 0.8452 | 727.2986 | 4.4866 |
| EL05 | 0.2894 | 0.3236 | 0.1081 | 2211.933 | 0.4325 | 0 | 2.5929 |
| EL06 | 0 | 0.0459 | 0.0232 | 479.8501 | 0.3495 | 0 | 0.4953 |
| EL07 | 0.0258 | 0.2194 | 0.0333 | 249.444 | 0.4269 | 0 | 0.674 |
| EL08 | 0.1009 | 0.3608 | 0.1047 | 1592.769 | 0.3797 | 1179.483 | 2.6663 |
| EL09 | 0.105 | 0.229 | 0.096 | 1990.069 | 0.3966 | 1086.756 | 2.2495 |
| EL10 | 0.0418 | 0.3328 | 0.0666 | 1157.213 | 0.2506 | 0 | 1.6691 |
| EL11 | 0.192 | 0.5701 | 0.1662 | 3937.007 | 0.6078 | 1489.509 | 3.9003 |
| EL12 | 0.0552 | 0 | 0.0198 | 33698.04 | 0.02 | 0 | 0.5562 |
| EL13 | 0.0389 | 0 | 0.0248 | 39265.82 | 0.0245 | 0 | 0.6437 |
| EL14 | 0 | 0.137 | 0.0169 | 49.2617 | 0.0338 | 0 | 0.5636 |
| EL15 | 0 | 0 | 0.0176 | 30.6542 | 0.0595 | 0 | 0.5447 |
| EL16 | 0 | 0 | 0.0114 | 27.5386 | 0.0436 | 0 | 0.481 |
| EL17 | 0 | 0 | 0.0049 | 62.897 | 0.0659 | 0 | 0.218 |
| EL18 | 0 | 0 | 0.0049 | 304.612 | 0.0662 | 0 | 0.1896 |
| EL19 | 0.0162 | 0.1719 | 0.0552 | 345.1442 | 0.3745 | 0 | 1.3838 |
| EL20 | 0.0126 | 0 | 0.009 | 150.7457 | 0.0973 | 0 | 0.2975 |
| EL21 | 0.0268 | 0 | 0.0079 | 91.7291 | 0.0171 | 0 | 0.3581 |
| EL22 | 0.0211 | 0 | 0.0073 | 78.1986 | 0.0112 | 0 | 0.2983 |
| EL23 | 0.0163 | 0 | 0.0096 | 65.4546 | 0.0095 | 77.33 | 0.4294 |
| EL24 | 0.013 | 0.2338 | 0.01 | 85.3663 | 0.0129 | 0 | 0.3872 |
| EL25 | 0.0208 | 0 | 0.0092 | 147.9299 | 0.0163 | 0 | 0.3055 |
| EL26 | 0.0237 | 0 | 0.009 | 166.9162 | 0.0169 | 0 | 0.3624 |
| EL27 | 0 | 0 | 0.0041 | 186.1024 | 0 | 0 | 0.1198 |
| EL28 | 0.027 | 0 | 0.0111 | 150.3192 | 0.1376 | 0 | 0.4736 |
| EL29 | 0.0147 | 0.087 | 0.0076 | 56.4655 | 0.0787 | 0 | 0.2983 |
| EL30 | 0 | 0 | 0.0063 | 44.3811 | 0.0092 | 0 | 0.2134 |
| EL31 | 0.0139 | 0 | 0.0082 | 83.4844 | 0.0115 | 0 | 0.3743 |
| EL32 | 0 | 0 | 0.0045 | 51.252 | 0.0043 | 0 | 0.1656 |
| EL33 | 0.0262 | 0 | 0.0069 | 144.9231 | 0.0073 | 0 | 0.2419 |
| EL34 | 0 | 0 | 0.0055 | 61.8807 | 0.0066 | 0 | 0.1324 |
| EL35 | 0.0126 | 0.1094 | 0.008 | 116.8608 | 0 | 0 | 0.4666 |
| EL36 | 0.0238 | 0 | 0.0096 | 241.2478 | 0.0142 | 1233.601 | 0.3292 |
| EL37 | 0.1551 | 0.2451 | 0.041 | 1166.068 | 0.1076 | 697.9789 | 1.2591 |
| EL38 | 0.0158 | 0 | 0.008 | 135.8086 | 0.012 | 0 | 0.2203 |
| EL39 | 0.0142 | 0 | 0.0058 | 89.3379 | 0.0091 | 0 | 0.2033 |
| EL40 | 0 | 0 | 0.0032 | 37.435 | 0.0072 | 0 | 0.1725 |
| EL41 | 0 | 0 | 0.0043 | 82.7982 | 0.18 | 0 | 0.1698 |
| EL42 | 0.0201 | 0 | 0.008 | 157.249 | 0.1469 | 0 | 0.3391 |
| EL43 | 0.0148 | 0 | 0.0075 | 135.1587 | 0.1153 | 0 | 0.221 |
| EL44 | 0.009 | 0 | 0.0046 | 71.3161 | 0.0294 | 173.2735 | 0.1336 |
| EL45 | 0.0264 | 0 | 0.0104 | 202.0459 | 0.1396 | 0 | 0.4223 |

Appendix 4, continued

| Sample ID | Cs | Dy | Eu | Fe | Hf | K | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL46 | 0.0347 | 0 | 0.0135 | 189.9125 | 0.1258 | 0 | 0.4246 |
| EL47 | 0.0101 | 0 | 0.0053 | 131.3949 | 0.0269 | 0 | 0.1685 |
| EL48 | 0.0145 | 0 | 0.0082 | 195.4129 | 0.0234 | 0 | 0.2537 |
| EL49 | 0 | 0 | 0.0026 | 53.22 | 0.059 | 0 | 0.1458 |
| EL50 | 0 | 0 | 0.0047 | 53.9828 | 0.0479 | 0 | 0.1811 |
| EL51 | 0.0178 | 0 | 0.004 | 458.5437 | 0.0696 | 0 | 0.1263 |
| EL52 | 0.0137 | 0 | 0.0074 | 401.6594 | 0.183 | 0 | 0.2404 |
| EL53 | 0.044 | 0 | 0.007 | 186.1854 | 0.0492 | 0 | 0.3028 |
| EL54 | 0.0218 | 0.1013 | 0.0302 | 275.3276 | 0.067 | 0 | 0.6393 |
| EL55 | 0.0474 | 0.2325 | 0.0516 | 741.6682 | 0.6514 | 0 | 1.1757 |
| EL56 | 0.0845 | 0.042 | 0.0272 | 585.0671 | 0.0529 | 0 | 0.9368 |
| EL57 | 0.4681 | 0.8316 | 0.1684 | 5694.186 | 0.8675 | 0 | 4.6481 |
| EL58 | 0.4577 | 0.575 | 0.1959 | 6407.481 | 1.4812 | 1158.705 | 5.3008 |
| EL59 | 0.5395 | 0.7299 | 0.2058 | 7424.528 | 1.0799 | 1609.493 | 5.4622 |
| EL60 | 0.0937 | 0.1402 | 0.0325 | 983.686 | 0.0781 | 1143.589 | 1.18 |
| EL61 | 0.708 | 2.7671 | 0.685 | 7669.845 | 0.5234 | 1951.388 | 15.049 |
| EL62 | 0.8627 | 3.7595 | 0.8822 | 8556.633 | 0.6152 | 1451.53 | 19.3173 |

Appendix 4, continued

| Sample ID | Lu | $\mathbf{M n}$ | $\mathbf{N a}$ | $\mathbf{N d}$ | $\mathbf{N i}$ | $\mathbf{R b}$ | $\mathbf{\text { Sb }}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL01 | 0.0301 | 34.3109 | 1064.067 | 1.7869 | 0 | 0.9209 | 0.0825 |
| EL02 | 0.035 | 43.5692 | 1003.023 | 2.7748 | 0 | 0.7298 | 0.1168 |
| EL03 | 0.0159 | 13.015 | 6456.238 | 3.5621 | 0 | 0.896 | 0.0309 |
| EL04 | 0.0544 | 27.0273 | 1112.977 | 3.676 | 0 | 7.5503 | 0.2019 |
| EL05 | 0.0348 | 14.1102 | 3559.249 | 2.3654 | 10.2143 | 4.5842 | 0.1056 |
| EL06 | 0.0102 | 10.7988 | 1599.525 | 0 | 2.3651 | 0 | 0.0617 |
| EL07 | 0.0168 | 16.9071 | 848.515 | 0.9552 | 0 | 0.8688 | 0.0363 |
| EL08 | 0.0377 | 57.1585 | 459.3869 | 2.8752 | 6.307 | 2.3893 | 0.0477 |
| EL09 | 0.033 | 29.0611 | 434.3038 | 2.126 | 3.3955 | 2.5947 | 0.1632 |
| EL10 | 0.0258 | 29.0441 | 245.9055 | 1.3345 | 0 | 0.7673 | 0.0997 |
| EL11 | 0.0651 | 40.316 | 720.0685 | 4.1882 | 16.818 | 4.1427 | 0.25 |
| EL12 | 0.0121 | 1339.166 | 670.811 | 1.009 | 0 | 0 | 0.0278 |
| EL13 | 0 | 1495.365 | 696.4515 | 1.9215 | 0 | 2.4554 | 0.064 |
| EL14 | 0.0091 | 8.1181 | 306.6308 | 0 | 0 | 0 | 0.011 |
| EL15 | 0.0136 | 2.788 | 344.1302 | 0 | 0 | 0 | 0.0061 |
| EL16 | 0.0046 | 52.6155 | 457.3216 | 0.8041 | 0 | 0 | 0 |
| EL17 | 0.0034 | 51.2094 | 361.7356 | 0.9619 | 1.6612 | 0 | 0.0096 |
| EL18 | 0.0034 | 161.3169 | 1312.553 | 1.1929 | 0 | 0 | 0.0469 |
| EL19 | 0.0252 | 51.7697 | 3177.694 | 1.6061 | 0 | 0 | 0.1093 |
| EL20 | 0.008 | 8.9752 | 1176.862 | 0 | 3.1257 | 0 | 0.1008 |
| EL21 | 0 | 2.5637 | 3058.591 | 0 | 0 | 0 | 0.0217 |
| EL22 | 0 | 2.0483 | 175.8513 | 0 | 2.2136 | 0 | 0.0218 |
| EL23 | 0 | 2.5053 | 191.0553 | 0 | 0 | 0 | 0.0046 |
| EL24 | 0.006 | 5.1695 | 160.2459 | 0 | 0 | 0 | 0.0449 |
| EL25 | 0 | 12.6175 | 220.3936 | 0 | 0 | 0 | 0.0637 |
| EL26 | 0.0158 | 6.538 | 345.404 | 0 | 0 | 0 | 0.3041 |
| EL27 | 0 | 1142.789 | 2634.11 | 0 | 0.6276 | 0 | 0.0263 |
| EL28 | 0.0042 | 7.3827 | 402.2969 | 0 | 0 | 0 | 0.0354 |
| EL29 | 0.0046 | 5.8379 | 133.4854 | 0 | 0 | 0 | 0.0056 |
| EL30 | 0 | 3.7516 | 123.0436 | 0 | 0 | 0 | 0.0149 |
| EL31 | 0.0167 | 6.6473 | 141.5274 | 0 | 0 | 0 | 0.0212 |
| EL32 | 0.013 | 2.9104 | 235.5294 | 0 | 1.605 | 0 | 0 |
| EL33 | 0 | 5.9085 | 143.226 | 0 | 0 | 0 | 0.1402 |
| EL34 | 0.0017 | 3.0269 | 111.214 | 0 | 0 | 0 | 0.0088 |
| EL35 | 0.0064 | 11.6866 | 176.8533 | 1.698 | 0 | 0 | 0.2018 |
| EL36 | 0 | 63.4345 | 354.399 | 0 | 2.1257 | 0.6697 | 0.0483 |
| EL37 | 0.0171 | 30.7654 | 552.128 | 1.7245 | 6.8641 | 2.3267 | 0.0504 |
| EL38 | 0 | 8.9243 | 4000.347 | 0 | 0 | 0 | 0.0049 |
| EL39 | 0 | 4.3713 | 139.7246 | 0 | 0 | 0 | 0.0574 |
| EL40 | 0 | 2.0882 | 108.7462 | 0 | 0 | 0 | 0.0118 |
| EL41 | 0.0028 | 36.0973 | 756.6639 | 0 | 2.103 | 0 | 0.0183 |
| EL42 | 0.0046 | 7.0797 | 247.3509 | 0 | 6.4983 | 0 | 0.0348 |
| EL43 | 0 | 68.0256 | 192.9405 | 0 | 3.4825 | 0 | 0.0317 |
| EL44 | 0.0015 | 24.5823 | 217.638 | 0 | 4.2704 | 0 | 0 |
| EL45 | 0.0058 | 11.9245 | 373.7765 | 0 | 0 | 0 | 0.0139 |

Appendix 4, continued

| Sample ID | $\mathbf{L u}$ | $\mathbf{M n}$ | $\mathbf{N a}$ | $\mathbf{N d}$ | $\mathbf{N i}$ | $\mathbf{R b}$ | $\mathbf{S b}$ |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL46 | 0.0033 | 9.5907 | 158.0595 | 0 | 0 | 0 | 0.0083 |
| EL47 | 0 | 4.2884 | 175.9028 | 0 | 1.809 | 0 | 0.0757 |
| EL48 | 0 | 11.0106 | 534.4362 | 0 | 4.3701 | 0.4496 | 0.2277 |
| EL49 | 0.0032 | 52.9297 | 321.0579 | 0 | 3.7417 | 0 | 0.0333 |
| EL50 | 0.0059 | 5.9865 | 190.9769 | 0 | 2.2595 | 0 | 0.0091 |
| EL51 | 0.0088 | 126.3166 | 415.8568 | 0 | 0 | 0 | 0.013 |
| EL52 | 0 | 140.4131 | 467.0687 | 0 | 0 | 0 | 0.0189 |
| EL53 | 0 | 4.1455 | 236.125 | 0 | 0 | 0.2931 | 0.038 |
| EL54 | 0.0212 | 16.1386 | 142.008 | 0 | 0 | 0 | 0.0121 |
| EL55 | 0.0394 | 58.2782 | 274.9984 | 2.1373 | 0 | 0 | 0.0464 |
| EL56 | 0 | 16.0943 | 449.1566 | 1.2677 | 0 | 1.6721 | 0.0523 |
| EL57 | 0.0472 | 77.0488 | 19189.6 | 6.467 | 13.1045 | 7.9769 | 0.1065 |
| EL58 | 0.0956 | 112.0825 | 1526.814 | 5.5222 | 13.0176 | 6.5001 | 0.2366 |
| EL59 | 0.0741 | 67.9393 | 2966.074 | 8.1652 | 13.3228 | 10.3911 | 0.1741 |
| EL60 | 0.0394 | 35.1365 | 2965.087 | 2.0817 | 4.4029 | 1.7252 | 0.0393 |
| EL61 | 0.2904 | 111.385 | 3460.12 | 14.7363 | 46.752 | 7.8564 | 0.2725 |
| EL62 | 0.3843 | 134.0708 | 3772.372 | 18.155 | 53.8096 | 9.2892 | 0.2524 |

Appendix 4, continued

| Sample ID | Sc | Sm | Sr | Ta | Tb | Th | Ti |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL01 | 0.2433 | 0.5554 | 11290.6 | 0.0193 | 0.0544 | 0.176 | 34.8701 |
| EL02 | 0.2777 | 0.6135 | 10933.83 | 0.0192 | 0.0682 | 0.1948 | 0 |
| EL03 | 0.1746 | 0.3685 | 7474.351 | 0.0165 | 0.0257 | 0.1094 | 0 |
| EL04 | 1.4133 | 0.9746 | 1828.375 | 0.1975 | 0.1257 | 1.0202 | 572.3628 |
| EL05 | 0.775 | 0.6597 | 3563.617 | 0.0919 | 0.0738 | 0.5334 | 383.6012 |
| EL06 | 0.118 | 0.1853 | 3638.516 | 0.009 | 0.0205 | 0.0871 | 0 |
| EL07 | 0.1558 | 0.2149 | 2093.064 | 0.0126 | 0.0323 | 0.0956 | 96.9171 |
| EL08 | 0.6648 | 0.5222 | 2493.24 | 0.047 | 0.0726 | 0.3125 | 212.6439 |
| EL09 | 0.67 | 0.5995 | 3211.235 | 0.0422 | 0.0667 | 0.4126 | 319.8443 |
| EL10 | 0.4781 | 0.4117 | 2298.087 | 0.0285 | 0.0391 | 0.1934 | 339.5098 |
| EL11 | 1.4061 | 0.9476 | 2277.717 | 0.1023 | 0.1136 | 0.6302 | 621.2758 |
| EL12 | 0.1427 | 0.2666 | 231.459 | 0 | 0.0201 | 0.0798 | 0 |
| EL13 | 0.1796 | 0.2386 | 246.1732 | 0 | 0 | 0.1136 | 0 |
| EL14 | 0.0354 | 0.0967 | 260.743 | 0 | 0.0109 | 0.0224 | 0 |
| EL15 | 0.0342 | 0.0925 | 269.2094 | 0 | 0.013 | 0.0238 | 0 |
| EL16 | 0.0219 | 0.0896 | 208.0103 | 0 | 0.0092 | 0.012 | 0 |
| EL17 | 0.0234 | 0.2002 | 229.6062 | 0 | 0 | 0.0137 | 0 |
| EL18 | 0.0438 | 0.2271 | 280.4822 | 0 | 0.0024 | 0.0212 | 0 |
| EL19 | 0.1305 | 0.2747 | 292.9138 | 0.0058 | 0.051 | 0.0723 | 0 |
| EL20 | 0.0494 | 0.1176 | 296.5218 | 0 | 0.0092 | 0.0251 | 0 |
| EL21 | 0.0794 | 0.123 | 353.4978 | 0 | 0.0092 | 0.034 | 0 |
| EL22 | 0.0583 | 0.0876 | 201.764 | 0 | 0.0091 | 0.0278 | 0 |
| EL23 | 0.0688 | 0.1056 | 294.1511 | 0 | 0.0127 | 0.0409 | 0 |
| EL24 | 0.0561 | 0.11 | 244.1501 | 0 | 0.0088 | 0.0381 | 0 |
| EL25 | 0.066 | 0.1032 | 321.1066 | 0 | 0.0054 | 0.0331 | 0 |
| EL26 | 0.0956 | 0.0963 | 256.7806 | 0.004 | 0.0078 | 0.0512 | 0 |
| EL27 | 0.0211 | 0.0602 | 203.2115 | 0 | 0 | 0.0101 | 0 |
| EL28 | 0.0688 | 0.152 | 230.861 | 0.0042 | 0.0091 | 0.0448 | 0 |
| EL29 | 0.0374 | 0.1145 | 189.2319 | 0.003 | 0.0063 | 0.0221 | 0 |
| EL30 | 0.0338 | 0.0599 | 209.5441 | 0 | 0 | 0.0138 | 0 |
| EL31 | 0.0395 | 0.1273 | 186.3802 | 0 | 0.0092 | 0.0219 | 0 |
| EL32 | 0.0289 | 0.0671 | 180.5539 | 0 | 0.0057 | 0.0122 | 0 |
| EL33 | 0.0504 | 0.0958 | 196.711 | 0 | 0.0054 | 0.0314 | 0 |
| EL34 | 0.0186 | 0.101 | 235.0531 | 0 | 0 | 0.0098 | 0 |
| EL35 | 0.0301 | 0.3679 | 442.0373 | 0 | 0.0099 | 0.0187 | 0 |
| EL36 | 0.0785 | 0.1421 | 313.5757 | 0 | 0.008 | 0.062 | 0 |
| EL37 | 0.3937 | 0.2973 | 479.3769 | 0.0304 | 0.0433 | 0.2549 | 0 |
| EL38 | 0.0461 | 0.0839 | 522.1274 | 0 | 0 | 0.0355 | 0 |
| EL39 | 0.0406 | 0.0864 | 255.1243 | 0 | 0.0057 | 0.026 | 0 |
| EL40 | 0.0211 | 0.1249 | 238.6739 | 0 | 0 | 0.013 | 0 |
| EL41 | 0.0172 | 0.0751 | 176.3389 | 0 | 0 | 0.0092 | 0 |
| EL42 | 0.0457 | 0.2141 | 295.343 | 0 | 0 | 0.0474 | 124.018 |
| EL43 | 0.0506 | 0.1149 | 155.4796 | 0.0053 | 0 | 0.0378 | 0 |
| EL44 | 0.0274 | 0.0766 | 128.7726 | 0 | 0 | 0.023 | 0 |
| EL45 | 0.0714 | 0.2061 | 461.2819 | 0.0051 | 0 | 0.0666 | 0 |

Appendix 4, continued

| Sample ID | Sc | Sm | Sr | Ta | Tb | Th | Ti |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| EL46 | 0.0933 | 0.1418 | 466.2859 | 0.0074 | 0.0161 | 0.0607 | 0 |
| EL47 | 0.0303 | 0.0809 | 255.434 | 0 | 0 | 0.0184 | 0 |
| EL48 | 0.0616 | 0.1373 | 183.3022 | 0 | 0 | 0.0453 | 0 |
| EL49 | 0.0105 | 0.0102 | 264.9862 | 0 | 0 | 0.0085 | 0 |
| EL50 | 0.0197 | 0.0521 | 186.202 | 0 | 0.0073 | 0.0182 | 0 |
| EL51 | 0.0356 | 0.0686 | 371.202 | 0 | 0 | 0.0397 | 0 |
| EL52 | 0.0458 | 0.1529 | 226.5613 | 0 | 0.0053 | 0.0329 | 0 |
| EL53 | 0.0711 | 0.2423 | 515.4393 | 0 | 0.0053 | 0.0661 | 0 |
| EL54 | 0.0991 | 0.2517 | 291.4455 | 0 | 0.0186 | 0.1114 | 0 |
| EL55 | 0.235 | 0.3519 | 205.0295 | 0.0177 | 0.0404 | 0.2945 | 0 |
| EL56 | 0.1736 | 0.4344 | 4278.684 | 0.0144 | 0.0155 | 0.1286 | 0 |
| EL57 | 1.7811 | 0.8381 | 1450.459 | 0.1541 | 0.0994 | 1.2051 | 367.3408 |
| EL58 | 1.9266 | 1.0331 | 1858.577 | 0.1788 | 0.1205 | 1.2621 | 781.2933 |
| EL59 | 2.0828 | 1.1204 | 1950.057 | 0.1751 | 0.1252 | 1.4198 | 699.3959 |
| EL60 | 0.2744 | 0.3335 | 2124.61 | 0.0197 | 0.0249 | 0.1467 | 0 |
| EL61 | 4.0379 | 2.8899 | 1248.565 | 0.1283 | 0.4862 | 1.5436 | 349.8646 |
| EL62 | 4.9028 | 3.6965 | 1120.609 | 0.1389 | 0.638 | 1.9149 | 224.1217 |

Appendix 4, continued

| Sample ID | $\mathbf{U}$ | $\mathbf{V}$ | $\mathbf{Y b}$ | Zn | $\mathbf{Z r}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| EL01 | 3.032 | 7.9294 | 0.2054 | 1.3756 | 22.5537 |
| EL02 | 2.8863 | 7.5073 | 0.2319 | 1.4084 | 21.9198 |
| EL03 | 2.4951 | 0 | 0.125 | 0.8558 | 21.8603 |
| EL04 | 2.9 | 18.7925 | 0.358 | 12.2408 | 41.1845 |
| EL05 | 2.3815 | 12.8581 | 0.2671 | 8.9789 | 27.4947 |
| EL06 | 0.9999 | 2.8176 | 0.1064 | 3.0109 | 17.4974 |
| EL07 | 1.0064 | 6.4402 | 0.1285 | 2.5326 | 19.0856 |
| EL08 | 1.7332 | 9.3488 | 0.2516 | 13.8195 | 17.4705 |
| EL09 | 2.1205 | 13.7889 | 0.3115 | 10.4437 | 24.2473 |
| EL10 | 1.8724 | 16.6078 | 0.1416 | 8.1415 | 18.6158 |
| EL11 | 3.2001 | 30.8785 | 0.3673 | 19.5619 | 35.1588 |
| EL12 | 2.2403 | 0 | 0.0502 | 8.6349 | 11.9428 |
| EL13 | 1.5971 | 12.1595 | 0.038 | 7.4144 | 6.6969 |
| EL14 | 0.332 | 0 | 0.0415 | 2.4006 | 3.279 |
| EL15 | 0.338 | 0.8985 | 0.0565 | 3.7945 | 3.3155 |
| EL16 | 0.578 | 0 | 0.0334 | 3.6516 | 5.0081 |
| EL17 | 2.0094 | 11.3434 | 0 | 8.1023 | 14.3838 |
| EL18 | 2.163 | 7.5182 | 0 | 7.8103 | 16.5354 |
| EL19 | 0.7837 | 7.5216 | 0.1735 | 13.7863 | 16.5669 |
| EL20 | 0.8816 | 4.0829 | 0.0495 | 10.6507 | 7.8324 |
| EL21 | 1.0257 | 0 | 0.0478 | 10.6218 | 6.5836 |
| EL22 | 0.7445 | 4.0603 | 0.0251 | 9.5381 | 5.1326 |
| EL23 | 0.6838 | 9.3305 | 0.0441 | 10.4029 | 5.3454 |
| EL24 | 0.838 | 8.4779 | 0.0411 | 10.0226 | 7.6276 |
| EL25 | 0.7143 | 5.7707 | 0.0276 | 8.8814 | 4.8567 |
| EL26 | 0.6064 | 7.8829 | 0.0342 | 11.7557 | 5.7535 |
| EL27 | 0.3747 | 17.6844 | 0 | 7.3011 | 3.3687 |
| EL28 | 1.1327 | 3.5604 | 0.0328 | 6.7768 | 11.0694 |
| EL29 | 1.0068 | 0 | 0.0366 | 5.8483 | 7.8417 |
| EL30 | 0.5529 | 3.0444 | 0.0253 | 3.6863 | 2.8913 |
| EL31 | 0.8298 | 1.9009 | 0.0224 | 5.3279 | 7.4458 |
| EL32 | 0.5908 | 0 | 0.0211 | 5.2849 | 3.2315 |
| EL33 | 0.7985 | 7.7772 | 0.0315 | 11.5935 | 5.7069 |
| EL34 | 0.7805 | 0 | 0 | 6.5919 | 5.1082 |
| EL35 | 3.7759 | 12.0885 | 0.0492 | 8.7755 | 22.4288 |
| EL36 | 1.1998 | 9.8503 | 0.0256 | 8.7215 | 7.4088 |
| EL37 | 1.3496 | 17.1491 | 0.1141 | 19.5972 | 9.4212 |
| EL38 | 0.5715 | 0 | 0 | 4.2853 | 5.6691 |
| EL39 | 0.8771 | 7.0999 | 0.0208 | 7.4022 | 6.0646 |
| EL40 | 1.2748 | 11.964 | 0 | 6.9998 | 7.9397 |
| EL41 | 0.6618 | 0 | 0 | 9.4886 | 7.6919 |
| EL42 | 1.6895 | 14.0219 | 0.0244 | 9.2345 | 15.5206 |
| EL43 | 1.0132 | 6.0041 | 0.0222 | 3.5119 | 9.6047 |
| EL44 | 0.7392 | 4.7171 | 0.0157 | 3.0409 | 4.6871 |
| EL45 | 1.5446 | 10.1572 | 0.0413 | 5.6344 | 12.4842 |
|  |  |  |  |  |  |

Appendix 4, continued

| Sample ID | $\mathbf{U}$ | $\mathbf{V}$ | $\mathbf{Y b}$ | $\mathbf{Z n}$ | $\mathbf{Z r}$ |
| :---: | :---: | :---: | :---: | :---: | :---: |
| EL46 | 0.9576 | 23.3888 | 0.037 | 6.1448 | 9.8634 |
| EL47 | 0.6377 | 0 | 0 | 3.0822 | 4.2509 |
| EL48 | 0.9706 | 9.3313 | 0 | 14.0141 | 7.142 |
| EL49 | 0.8685 | 4.3765 | 0 | 3.0171 | 6.808 |
| EL50 | 0.3305 | 0 | 0.0365 | 6.8323 | 3.8955 |
| EL51 | 0.4956 | 6.499 | 0 | 2.1383 | 5.9543 |
| EL52 | 1.2156 | 7.2502 | 0.0231 | 6.5661 | 13.4194 |
| EL53 | 2.2697 | 13.3964 | 0 | 2.2548 | 15.5998 |
| EL54 | 1.0953 | 4.5943 | 0.0838 | 14.4077 | 9.6091 |
| EL55 | 0.9622 | 6.6202 | 0.1437 | 15.1188 | 24.6919 |
| EL56 | 3.6851 | 25.4008 | 0.0902 | 9.3571 | 25.9875 |
| EL57 | 1.768 | 54.6314 | 0.3654 | 31.0637 | 32.2842 |
| EL58 | 1.7584 | 55.2784 | 0.4872 | 33.4192 | 53.2964 |
| EL59 | 2.5504 | 62.9306 | 0.3317 | 40.1934 | 37.5222 |
| EL60 | 2.0311 | 33.2006 | 0.1042 | 14.253 | 15.7593 |
| EL61 | 2.6861 | 42.5034 | 1.9354 | 67.3745 | 49.4551 |
| EL62 | 2.702 | 45.6051 | 2.48 | 63.6789 | 41.7665 |

Appendix 5 - Geographical data for African obsidian samples from Kenya

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  | ${ }^{\circ}$ N or ${ }^{\circ}$ S | ${ }^{\circ}$ E |
| KES001 | Kinangop \#1 | -0.63350 | 36.48714 |
| KES002 | Kinangop \#2 | -0.57367 | 36.49033 |
| KES003 | Kongoni area Mundui road cut | -0.81272 | 36.26106 |
| KES004 | Sonachi | -0.78047 | 36.26683 |
| KES005 | N. Lake Rd quarry pit | -0.66817 | 36.32500 |
| KES006 | Eburru GsJj53/52 area | -0.62014 | 36.31361 |
| KES007 | Masai Gorge rd quarry | -0.64647 | 36.33419 |
| KES008 | Eburu/GilGil Elmenteita Junction GsJj82 | -0.58619 | 36.26408 |
| KES009 | Eburu/GilGil Elmenteita Junction GsJj82 | -0.58619 | 36.26408 |
| KES010 | N. Eburu Rd. scree | -0.58619 | 36.27667 |
| KES011 | N. Eburu Rd. scree | -0.58619 | 36.27667 |
| KES012 | Gilgil-Eburu Rd pumice bed | -0.54461 | 36.29850 |
| KES013 | Gilgil-Eburu Rd pumice bed | -0.63450 | 36.25594 |
| KES014 | Lukenya quarry fiame | -1.48006 | 37.08367 |
| KES015 | Lukenya quarry bomb | -1.48006 | 37.08367 |
| KES016 | Upper Kedong rd | -1.23286 | 36.54603 |
| KES017 | Upper Kedong rd | -1.23286 | 36.54603 |
| KES018 | Gicheru | -1.18706 | 36.54944 |
| KES019 | Gicheru | -1.18706 | 36.54944 |
| KES020 | Mid-Kedong Valley | -1.10297 | 36.46361 |
| KES021 | Dawson's Camp | -0.98572 | 36.31286 |
| KES022 | Dawson's Camp | -0.98572 | 36.31286 |
| KES023 | Hell's Gate South Entrance | -0.97489 | 36.30689 |
| KES024 | Hell's Gate South Entrance | -0.97489 | 36.30689 |
| KES025 | Hell's Gate S Upper Hill | -0.97181 | 36.30992 |
| KES026 | Hell's Gate S Upper Hill | -0.97181 | 36.30992 |
| KES027 | Hell's Gate South East Wall | -0.95928 | 36.31014 |
| KES028 | Hell's Gate South East Wall | -0.95928 | 36.31014 |
| KES029 | Obsidian Cave, Hell's Gate Nat. Pk. | -0.88781 | 36.38375 |
| KES030 | Obsidian Cave, Hell's Gate Nat. Pk. | -0.88781 | 36.38375 |
| KES031 | Obsidian Cave, Hell's Gate Nat. Pk. | -0.88781 | 36.38375 |
| KES032 | Obsidian Cave, Upper | -0.88781 | 36.38375 |
| KES033 | Eburu top 100m E Kinogono school | -0.63044 | 36.25431 |
| KES034 | Eburu top 100m E Kinogono school | -0.63044 | 36.25431 |
| KES035 | Eburu top 100m E Kinogono sch | -0.63044 | 36.25431 |
| KES036 | Eburu top 100m E Kinogono sch | -0.63044 | 36.25431 |
| KES037 | Eburu top Kinogono sch | -0.63450 | 36.25594 |
| KES038 | Eburu top Kinogono sch | -0.63450 | 36.25594 |
| KES039 | Eburu top Kinogono sch | -0.63450 | 36.25594 |
| KES040 | Eburu top Kinogono sch | -0.63450 | 36.25594 |
| KES042 | Old Tepesi Rock Shelter | -0.69192 | 36.20728 |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES043 | Eburu Station Rd. | -0.58108 | 36.24889 |
| KES044 | Eburu Station Rd. | -0.58108 | 36.24889 |
| KES045 | Eburu Station Rd. 150 m S | -0.58192 | 36.24972 |
| KES046 | Eburu Station Rd. 150 m S | -0.58192 | 36.24972 |
| KES047 | Hell's Gate main Middle | -0.87978 | 36.34272 |
| KES048 | Hell's Gate Central Tower | -0.89667 | 36.32436 |
| KES049 | Hell's Gate Central Tower | -0.89667 | 36.32436 |
| KES050 | Ol Karia | -0.89819 | 36.30919 |
| KES051 | Ol Karia | -0.89819 | 36.30919 |
| KES052 | Ol Karia II | -0.87436 | 36.29433 |
| KES053 | Ol Karia II | -0.87436 | 36.29433 |
| KES054 | Ol Jorai Quarry base | -0.59700 | 36.21092 |
| KES055 | Ol Jorai Quarry west | -0.60008 | 36.21108 |
| KES056 | Ol Jorai Quarry GsJi59 | -0.60008 | 36.21108 |
| KES057 | Menengai lookout | -0.22656 | 36.09567 |
| KES058 | Prospect Farm Upper | -0.61803 | 36.18994 |
| KES059 | GsJj53 Acheulean/MSA | -0.61603 | 36.30694 |
| KES060 | Marula Estate (upper) | -0.62211 | 36.31333 |
| KES061 | Marula Estate (upper) | -0.62211 | 36.31333 |
| KES062 | GsJj84 Gema 1 | -0.54556 | 36.31406 |
| KES063 | Gema 2 planar exposure | -0.56433 | 36.30875 |
| KES064 | Gema 2 planar exposure | -0.56433 | 36.30875 |
| KES065 | Marula Valley NE | -0.58756 | 36.32039 |
| KES066 | Ol Doinyo Nyokie (Magadi) | -1.80556 | 36.37500 |
| KES067 | GsJj85 Nagum | -0.57686 | 36.30044 |
| KES068 | GsJj85 Nagum | -0.57686 | 36.30044 |
| KES069 | GsJj85 Nagum | -0.57686 | 36.30044 |
| KES070 | GsJj53 Southwest | -0.61608 | 36.30856 |
| KES071 | GsJj53 Southwest | -0.61608 | 36.30856 |
| KES072 | GsJj53 Southwest | -0.61608 | 36.30856 |
| KES073 | GsJj53 Southwest | -0.61608 | 36.30856 |
| KES074 | GsJj53 Southwest | -0.61608 | 36.30856 |
| KES075 | Fisherman's Camp | -0.82833 | 36.33589 |
| KES076 | Mundui at student outcrop | -0.80911 | 36.25739 |
| KES077 | Mundui at student outcrop | -0.80911 | 36.25739 |
| KES078 | Sonachi Green Crater Lake | -0.78047 | 36.26683 |
| KES079 | N. Lake Rd. Quarry | -0.66817 | 36.32500 |
| KES080 | Eburu Rd N. | -0.66817 | 36.32500 |
| KES081 | GsJj50 Area A | -0.62947 | 36.25372 |
| KES082 | GsJj50 Area A | -0.62947 | 36.25372 |
| KES083 | GsJj50 Area B | 36.25372 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES084 | GsJj50 Area B | -0.62947 | 36.25372 |
| KES085 | GsJj50 Area B | -0.62947 | 36.25372 |
| KES086 | GsJj50 Area B | -0.62947 | 36.25372 |
| KES087 | Rd. Quarry above Masai Gorge | -0.64644 | 36.33411 |
| KES088 | Rd. Quarry above Masai Gorge | -0.64644 | 36.33411 |
| KES089 | Marula Valley NW | -0.60522 | 36.30456 |
| KES090 | Marula Valley NW | -0.60522 | 36.30456 |
| KES091 | GilGil Eburu pumice outcrop | -0.56475 | 36.30022 |
| KES092 | GilGil Eburu pumice outcrop | -0.56475 | 36.30022 |
| KES093 | Gilgil Eburu Rd. 2 | -0.56769 | 36.29733 |
| KES094 | Gilgil Eburu Rd. 2 | -0.56769 | 36.29733 |
| KES095 | Ol Doinyo Njeru Ofisini | -0.56558 | 36.28917 |
| KES096 | Ol Doinyo Njeru Ofisini | -0.56558 | 36.28917 |
| KES097 | Ol Doinyo Njeru Ofisini | -0.56558 | 36.28917 |
| KES098 | Lukenya Rd. | -1.46267 | 37.11436 |
| KES099 | Lukenya Rd. | -1.46267 | 37.11436 |
| KES100 | Lukenya Rd. | -1.46267 | 37.11436 |
| KES101 | Ol Doinyo Alasho | -2.15297 | 36.15686 |
| KES102 | Ol Doinyo Alasho | -2.15297 | 36.15686 |
| KES103 | Ol Doinyo Alasho | -2.15297 | 36.15686 |
| KES104 | Ol Doinyo Alasho | -2.15297 | 36.15686 |
| KES124 | Ol Doinyo Nyokie Loc 1 | -1.80508 | 36.37708 |
| KES125 | Ol Doinyo Nyokie Loc 1 | -1.80508 | 36.37708 |
| KES126 | Ol Doinyo Nyokie Loc 2 | -1.80269 | 36.37631 |
| KES127 | Ol Doinyo Nyokie Loc 3 | -1.79767 | 36.37203 |
| KES128 | Ol Doinyo Nyokie Loc 3 | -1.79725 | 36.37147 |
| KES130 | Ol Doinyo Nyokie Loc 5 | -1.78850 | 36.37369 |
| KES131 | Ol Doinyo Nyokie Loc 6 | -1.78875 | 36.37083 |
| KES132 | Ol Doinyo Nyokie Loc 6 | -1.78933 | 36.37131 |
| KES133 | Ewuaso-Suswa Road Loc 1 | -1.10275 | 36.46403 |
| KES134 | Ewuaso-Suswa Road Loc 1 | -1.10283 | 36.46336 |
| KES135 | Ewuaso-Suswa Road Loc 1 | -1.10283 | 36.46336 |
| KES136 | Ewuaso-Suswa Road Loc 1 | -1.10275 | 36.46264 |
| KES137 | Salasun | -1.11119 | 36.39047 |
| KES138 | Salasun | -1.11119 | 36.39047 |
| KES139 | Salasun | -1.11083 | 36.38975 |
| KES140 | Ewuaso-Suswa Road Loc 2 | -1.10047 | 36.45847 |
| KES141 | Ewuaso-Suswa Road Loc 2 | -1.10053 | 36.45811 |
| KES147 | Naivasha Top Camp Loc 1 | -0.83383 | 36.33228 |
| KES148 | Naivasha Top Camp Loc 1 | -0.83392 | 36.33228 |
| KES149 | Naivasha Top Camp Loc 1 | 36.33228 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  | ${ }^{\circ}$ E |  |
| KES150 | Naivasha Top Camp Loc 1 | -0.83417 | 36.33228 |
| KES151 | Naivasha Top Camp Loc 1 | -0.83417 | 36.33228 |
| KES152 | Ol Doinyo Oserian | -0.81808 | 36.31817 |
| KES153 | Ol Doinyo Oserian | -0.83211 | 36.31842 |
| KES154 | Ol Doinyo Oserian | -0.83186 | 36.31853 |
| KES155 | Ol Doinyo Oserian | -0.83186 | 36.31853 |
| KES156 | Ol Doinyo Oserian Quarry | -0.83244 | 36.31972 |
| KES157 | Ol Doinyo Oserian Quarry | -0.83244 | 36.31972 |
| KES158 | Ol Doinyo Oserian Quarry | -0.83244 | 36.31972 |
| KES159 | Naivasha Top Camp Loc 2 | -0.83206 | 36.33211 |
| KES160 | Naivasha Top Camp Loc 2 | -0.83192 | 36.33194 |
| KES161 | Naivasha Top Camp Loc 2 | -0.83175 | 36.33183 |
| KES162 | Naivasha Top Camp Loc 3 | -0.83014 | 36.33333 |
| KES163 | Naivasha Top Camp Loc 3 | -0.83103 | 36.33339 |
| KES164 | Naivasha Top Camp Loc 3 | -0.83103 | 36.33339 |
| KES165 | GsJi53 Kiteko Loc 1 | -0.69175 | 36.20717 |
| KES166 | GsJi53 Kiteko Loc 1 | -0.69175 | 36.20717 |
| KES167 | GsJi53 Kiteko Loc 2 | -0.69269 | 36.20508 |
| KES168 | Delamere Dam Loc 1 | -0.68811 | 36.20342 |
| KES169 | Delamere Dam Loc 2 | -0.69189 | 36.20211 |
| KES170 | Ololerai Loc 1 | -0.79294 | 36.31275 |
| KES171 | Ololerai Loc 2 | -0.78769 | 36.30931 |
| KES172 | Ololerai Loc 3 | -0.79133 | 36.31256 |
| KES173 | Ololerai Loc 4 | -0.79428 | 36.30783 |
| KES174 | Ololerai Loc 5 | -0.79275 | 36.30408 |
| KES175 | Ololerai Loc 6 | -0.79572 | 36.30339 |
| KES176 | Mundui Loc 1 | -0.81322 | 36.26144 |
| KES177 | Mundui Loc 2 | -0.80722 | 36.25661 |
| KES178 | Mundui Loc 2 | -0.80722 | 36.25661 |
| KES179 | Mundui Loc 3 | -0.80350 | 36.25278 |
| KES180 | Crater Lake Loc 1 | -0.78161 | 36.25894 |
| KES181 | Crater Lake Loc 1 | -0.78161 | 36.25894 |
| KES182 | Ilkek Loc 1 | -0.59431 | 36.36619 |
| KES183 | Ilkek Loc 2 | -0.59386 | 36.36536 |
| KES184 | Ilkek Loc 3 | -0.59356 | 36.36433 |
| KES185 | Ilkek Loc 4 | -0.59411 | 36.36300 |
| KES186 | Ilkek Loc 5 | -0.59842 | 36.36589 |
| KES187 | Ilkek Drift Hill Loc 1 | -0.59650 | 36.35333 |
| KES188 | Ilkek Drift Hill Loc 2 | -0.59672 | 36.35236 |
| KES189 | Ilkek Drift Hill Loc 3 | -0.59925 | 36.34600 |
| KES190 | Ilkek Drift Hill Loc 3 | -0.59925 | 36.34600 |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES191 | Waterloo Ridge Loc 1 | -0.60519 | 36.34178 |
| KES192 | Waterloo Ridge Loc 2a | -0.58489 | 36.33775 |
| KES193 | Waterloo Ridge Loc 2b | -0.58447 | 36.33822 |
| KES194 | Waterloo Ridge Loc 2c | -0.58553 | 36.33664 |
| KES195 | Waterloo Ridge Loc 3a | -0.60792 | 36.33542 |
| KES196 | Waterloo Ridge Loc 3b | -0.60856 | 36.33547 |
| KES197 | Waterloo Ridge Loc 3c | -0.60844 | 36.33589 |
| KES198 | Waterloo Ridge Loc 3d | -0.60958 | 36.33647 |
| KES199 | Waterloo Ridge Loc 4a | -0.61217 | 36.33647 |
| KES200 | Waterloo Ridge Loc 4b | -0.61264 | 36.33658 |
| KES201 | Waterloo Ridge Loc 4c | -0.61294 | 36.33661 |
| KES202 | Waterloo Ridge Loc 4d | -0.61250 | 36.33650 |
| KES203 | Waterloo Ridge Loc 5a | -0.61558 | 36.33744 |
| KES204 | Waterloo Ridge Loc 5b | -0.61594 | 36.33753 |
| KES205 | Waterloo Ridge Loc 5c | -0.61619 | 36.33711 |
| KES206 | Waterloo Ridge Loc 6a | -0.61883 | 36.33714 |
| KES207 | Waterloo Ridge Loc 6b | -0.61792 | 36.33719 |
| KES208 | Waterloo Ridge Loc 6c | -0.61731 | 36.33756 |
| KES209 | Waterloo Ridge Loc 7a | -0.62050 | 36.34206 |
| KES210 | Waterloo Ridge Loc 7b | -0.62067 | 36.34153 |
| KES211 | Waterloo Ridge Loc 7c | -0.62114 | 36.34108 |
| KES212 | Waterloo Ridge Loc 7d | -0.62111 | 36.34306 |
| KES213 | Waterloo Ridge Loc 8a | -0.61861 | 36.34406 |
| KES214 | Waterloo Ridge Loc 8b | -0.61869 | 36.34369 |
| KES215 | Waterloo Ridge Loc 8c | -0.61886 | 36.34319 |
| KES216 | Waterloo Ridge Loc 9 | -0.61936 | 36.34700 |
| KES217 | Waterloo Ridge Loc 10 | -0.62172 | 36.34519 |
| KES218 | Elsa Loc 1a | -0.81578 | 36.31628 |
| KES219 | Elsa Loc 1b | -0.81575 | 36.31564 |
| KES220 | Waterloo Ridge Loc 11a | -0.62258 | 36.35450 |
| KES221 | Waterloo Ridge Loc 11b | -0.62297 | 36.35483 |
| KES222 | Waterloo Ridge Loc 12a | -0.62625 | 36.34875 |
| KES223 | Waterloo Ridge Loc 12b | -0.62622 | 36.34856 |
| KES224 | Waterloo Ridge Loc 13a | -0.62822 | 36.35133 |
| KES225 | Waterloo Ridge Loc 13b | -0.62714 | 36.35147 |
| KES226 | Waterloo Ridge Loc 13c | -0.62814 | 36.34833 |
| KES227 | Waterloo Ridge Loc 13d | -0.62731 | 36.34819 |
| KES228 | Waterloo Ridge Loc 13e | -0.62900 | 36.34825 |
| KES229 | Waterloo Ridge Loc 14a | -0.61994 | 36.35347 |
| KES230 | Waterloo Ridge Loc 14b | -0.62003 | 36.35317 |
| KES231 | Waterloo Ridge Loc 15 | 36.35433 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES232 | Waterloo Ridge Loc 16 | -0.57711 | 36.33433 |
| KES233 | Waterloo Ridge Loc 16 | -0.57711 | 36.33433 |
| KES234 | Waterloo Ridge Loc 17a | -0.57211 | 36.33400 |
| KES235 | Waterloo Ridge Loc 17b | -0.56972 | 36.33322 |
| KES236 | Waterloo Ridge Loc 18 | -0.56675 | 36.32972 |
| KES237 | Waterloo Ridge Loc 19a | -0.57739 | 36.32672 |
| KES238 | Waterloo Ridge Loc 19b | -0.57819 | 36.32681 |
| KES239 | Waterloo Ridge Loc 20a | -0.58219 | 36.32806 |
| KES240 | Waterloo Ridge Loc 20b | -0.58339 | 36.32833 |
| KES241 | Waterloo Ridge Loc 21a | -0.58936 | 36.33036 |
| KES242 | Waterloo Ridge Loc 21b | -0.58972 | 36.33047 |
| KES243 | Njorowa Loc 1a | -0.96578 | 36.30619 |
| KES244 | Njorowa Loc 1b | -0.96578 | 36.30619 |
| KES245 | Njorowa Loc 2a | -0.95925 | 36.31008 |
| KES246 | Njorowa Loc 2b | -0.95936 | 36.31036 |
| KES247 | Njorowa Loc 2c | -0.95925 | 36.31008 |
| KES248 | Njorowa Loc 3a | -0.95758 | 36.31142 |
| KES249 | Njorowa Loc 3b | -0.95758 | 36.31142 |
| KES250 | Njorowa Loc 3c | -0.95778 | 36.31158 |
| KES251 | Njorowa Loc 4 | -0.97169 | 36.30425 |
| KES252 | Njorowa Loc 5a | -0.96903 | 36.30819 |
| KES253 | Njorowa Loc 5b | -0.96903 | 36.30819 |
| KES254 | Njorowa Loc 6 | -0.97411 | 36.30631 |
| KES255 | Gicheru Loc 1a | -1.18694 | 36.54944 |
| KES256 | Gicheru Loc 1b | -1.18722 | 36.54931 |
| KES257 | Gicheru Loc 1c | -1.18719 | 36.54961 |
| KES258 | Gicheru Loc 1d | -1.18719 | 36.55008 |
| KES259 | Gicheru Loc 1e | -1.18644 | 36.54931 |
| KES260 | Gicheru Loc 2a | -1.18475 | 36.55072 |
| KES261 | Gicheru Loc 2b | -1.18475 | 36.55072 |
| KES262 | Gicheru Loc 2c | -1.18475 | 36.55072 |
| KES263 | Gicheru Loc 2d | -1.18475 | 36.55072 |
| KES264 | Gicheru Loc 3a | -1.19336 | 36.54500 |
| KES265 | Gicheru Loc 3b | -1.19336 | 36.54500 |
| KES266 | Gicheru Loc 3c | -1.19281 | 36.54506 |
| KES267 | Gicheru Loc 3d | -1.19281 | 36.54506 |
| KES268 | Kedong Road Loc 1a | -1.22589 | 36.54636 |
| KES269 | Kedong Road Loc 1b | -1.22589 | 36.54636 |
| KES270 | Kedong Road Loc 1c | -1.22589 | 36.54636 |
| KES271 | Kedong Road Loc 2a | -1.22669 | 36.54914 |
| KES272 | Kedong Road Loc 2b | 36.22669 | 36.54914 |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES273 | Kedong Road Loc 2c | -1.22669 | 36.54914 |
| KES274 | Kedong Road Loc 3a | -1.22850 | 36.54889 |
| KES275 | Kedong Road Loc 3b | -1.22850 | 36.54889 |
| KES276 | Kedong Road Loc 4 | -1.23142 | 36.54836 |
| KES277 | Kedong Road Loc 5 | -1.23311 | 36.54669 |
| KES278 | Njorowa Loc 7 | -0.95528 | 36.31403 |
| KES279 | Njorowa Loc 8 | -0.95461 | 36.31469 |
| KES281 | Njorowa Loc 7b | -0.95561 | 36.31406 |
| KES282 | Dawson's Camp Loc 1 | -0.98719 | 36.30922 |
| KES283 | Dawson's Camp Loc 2 | -0.98578 | 36.30864 |
| KES284 | Dawson's Camp Loc 3 | -0.98531 | 36.31008 |
| KES285 | Menengai Loc 1 | -0.22583 | 36.09681 |
| KES286 | Menengai Loc 2 | -0.22903 | 36.09483 |
| KES287 | Lion Hill Loc 1 | -0.31153 | 36.12078 |
| KES288 | Lion Hill Loc 2 | -0.31183 | 36.12389 |
| KES289 | Bahati Loc 1a | -0.16661 | 36.12117 |
| KES290 | Bahati Loc 1b | -0.16661 | 36.12117 |
| KES291 | Banwala Loc 1 | -0.09403 | 36.02192 |
| KES292 | Lasibil Loc 1 | -0.10064 | 35.99069 |
| KES293 | Kampi Ya Moto Loc 3 | -0.11953 | 35.95022 |
| KES294 | Kampi Ya Moto Loc 1a | -0.11614 | 35.95389 |
| KES295 | Kampi Ya Moto Loc 1b | -0.11631 | 35.95381 |
| KES296 | Kampi Ya Moto Loc 1c | -0.11631 | 35.95381 |
| KES297 | Naivasha Top Camp Loc 4a | -0.82839 | 36.33519 |
| KES298 | Naivasha Top Camp Loc 4b | -0.82839 | 36.33519 |
| KES299 | Naivasha Top Camp Loc 4c | -0.82833 | 36.33500 |
| KES300 | Orengenai Loc 1a | -0.68425 | 36.33297 |
| KES301 | Orengenai Loc 1b | -0.68489 | 36.33294 |
| KES302 | Orengenai Loc 2a | -0.68608 | 36.33147 |
| KES303 | Orengenai Loc 2b | -0.68550 | 36.33142 |
| KES304 | Orengenai Loc 3a | -0.68842 | 36.32889 |
| KES305 | Orengenai Loc 3b | -0.68853 | 36.32892 |
| KES306 | Orengenai Loc 4a | -0.68539 | 36.32742 |
| KES307 | Orengenai Loc 4b | -0.68469 | 36.32700 |
| KES308 | Orengenai Loc 4c | -0.68581 | 36.32728 |
| KES309 | Orengenai Loc 4d | -0.68322 | 36.32675 |
| KES310 | Orengenai Loc 4e | -0.68289 | 36.32725 |
| KES311 | Orengenai Loc 4f | -0.68161 | 36.32769 |
| KES312 | Orengenai Loc 4g | -0.68067 | 36.32764 |
| KES313 | Orengenai Log 4h | -0.68458 | 36.32567 |
| KES314 | Orengenai Loc 5a | 36.32761 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  | ${ }^{\circ}$ E |  |
| KES315 | Orengenai Loc 5b | -0.68442 | 36.32536 |
| KES316 | Orengenai Loc 6a | -0.68181 | 36.32558 |
| KES317 | Orengenai Loc 6b | -0.68181 | 36.32558 |
| KES318 | Baixia Loc 1 | -0.68122 | 36.31814 |
| KES319 | Green Park Loc 1 | -0.66808 | 36.29164 |
| KES320 | Murwa Loc 1 | -0.65472 | 36.21419 |
| KES321 | Murwa Loc 2 | -0.65583 | 36.23617 |
| KES322 | Eburru Forest Station Loc 1 | -0.64619 | 36.24703 |
| KES323 | Eburru Forest Station Loc 2a | -0.63997 | 36.25292 |
| KES324 | Eburru Condensor Loc 1 | -0.64014 | 36.25336 |
| KES325 | Eburru Condensor Loc 2a | -0.63975 | 36.25353 |
| KES326 | Eburru Condensor Loc 2b | -0.63956 | 36.25356 |
| KES327 | Eburru Condensor Loc 2c | -0.63947 | 36.25353 |
| KES328 | Eburru Condensor Loc 3a | -0.63903 | 36.25361 |
| KES329 | Eburru Condensor Loc 3b | -0.63861 | 36.25381 |
| KES330 | Eburru Condensor Loc 3c | -0.63839 | 36.25350 |
| KES331 | Eburru Condensor Loc 3d | -0.63800 | 36.25375 |
| KES335 | GsJj50 Area A | -0.62961 | 36.25378 |
| KES336 | GsJj50 Area B1 | -0.63031 | 36.25431 |
| KES337 | GsJj50 Area B2 | -0.63019 | 36.25406 |
| KES338 | GsJj50 Area C1 | -0.63419 | 36.25594 |
| KES339 | GsJj50 Area C2 | -0.63467 | 36.25589 |
| KES340 | GsJj50 Area C3 | -0.63436 | 36.25536 |
| KES341 | Masai Gorge Loc 1a | -0.65992 | 36.33072 |
| KES342 | Masai Gorge Loc 1b | -0.65992 | 36.33072 |
| KES343 | Masai Gorge Loc 2a | -0.65975 | 36.33281 |
| KES344 | Masai Gorge Loc 2b | -0.65983 | 36.33286 |
| KES345 | Eburru North Road Loc 1a | -0.64558 | 36.30847 |
| KES346 | Eburru North Road Loc 1b | -0.64569 | 36.30733 |
| KES347 | GsJj52 Loc 2a | -0.62211 | 36.31308 |
| KES348 | GsJj52 Loc 2b | -0.62203 | 36.31336 |
| KES349 | GsJj52 Loc 2c | -0.62208 | 36.31386 |
| KES350 | GsJj52 Loc 3a | -0.62100 | 36.31419 |
| KES351 | GsJj52 Loc 3b | -0.62089 | 36.31450 |
| KES352 | GsJj52 Loc 3c | -0.62142 | 36.31464 |
| KES353 | GsJj52 Loc 3d | -0.62014 | 36.31353 |
| KES354 | GsJj52 Loc 4a | -0.61764 | 36.31183 |
| KES355 | GsJj52 Loc 4b | -0.61778 | 36.31192 |
| KES356 | GsJj52 Loc 4c | -0.61736 | 36.31153 |
| KES357 | GsJj52 Loc 4d | -0.61725 | 36.30997 |
| KES358 | GsJj52 Loc 1c | -0.61636 | 36.30933 |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  | ${ }^{\circ}$ E |  |
| KES359 | GsJj52 Loc 1c | -0.61636 | 36.30933 |
| KES360 | GsJj52 Loc 1b | -0.61594 | 36.30922 |
| KES361 | GsJj52 Loc 1a | -0.61606 | 36.30850 |
| KES362 | GsJj53 Loc 1a | -0.61419 | 36.30725 |
| KES363 | GsJj53 Loc 1b | -0.61428 | 36.30722 |
| KES364 | Masai Gorge Loc 3a | -0.64639 | 36.33431 |
| KES365 | Masai Gorge Loc 3b | -0.64592 | 36.33406 |
| KES366 | Masai Gorge Loc 3b | -0.64592 | 36.33406 |
| KES367 | Masai Gorge Loc 4a | -0.64928 | 36.33453 |
| KES368 | Masai Gorge Loc 4b | -0.64928 | 36.33453 |
| KES369 | Masai Gorge Loc 4c | -0.64928 | 36.33453 |
| KES370 | Entorobonni Loc 1 | -0.91739 | 35.69475 |
| KES371 | Entorobonni Loc 1 | -0.91739 | 35.69475 |
| KES372 | Entorobonni Loc 2 | -0.91844 | 35.69511 |
| KES373 | Entorobonni Loc 2 | -0.91844 | 35.69511 |
| KES374 | Hell's Gate Loc 1a (Obsidian Caves track) | -0.86961 | 36.37453 |
| KES375 | Hell's Gate Loc 1b | -0.86961 | 36.37453 |
| KES376 | Hell's Gate Loc 2a | -0.87150 | 36.37856 |
| KES377 | Hell's Gate Loc 2b | -0.87150 | 36.37856 |
| KES378 | Hell's Gate Loc 2b | -0.87150 | 36.37856 |
| KES379 | Hell's Gate Loc 2c | -0.87150 | 36.37856 |
| KES380 | Hell's Gate Loc 3a | -0.88653 | 36.38347 |
| KES381 | Hell's Gate Loc 3b | -0.88653 | 36.38347 |
| KES382 | Hell's Gate Loc 3c | -0.88653 | 36.38347 |
| KES383 | Hell's Gate Loc 4a, upper flow (Obsidian Caves) | -0.88769 | 36.38378 |
| KES384 | Hell's Gate Loc 4a, upper flow (Obsidian Caves) | -0.88769 | 36.38378 |
| KES385 | Hell's Gate Loc 4a, lower flow | -0.88769 | 36.38378 |
| KES386 | Hell's Gate Loc 4b, lower flow | -0.88800 | 36.38392 |
| KES387 | Hell's Gate Loc 4b, middle flow | -0.88800 | 36.38392 |
| KES388 | Hell's Gate Loc 4b, upper flow | -0.88800 | 36.38392 |
| KES389 | Hell's Gate Loc 5a | -0.90192 | 36.38028 |
| KES390 | Hell's Gate Loc 5b | -0.90192 | 36.38028 |
| KES391 | Hell's Gate Loc 5c | -0.90192 | 36.38028 |
| KES392 | Hell's Gate Loc 6 (Fischer's Tower Drift) | -0.89667 | 36.32433 |
| KES393 | Hell's Gate Loc 7 (Fischer's Tower Gorge) | -0.89544 | 36.32067 |
| KES394 | Hell's Gate Loc 8 (Olkaria) | -0.89828 | 36.30717 |
| KES403 | Kalusha Loc 1c (Lukenya) | -1.44781 | 37.10525 |
| KES404 | Lukenya Road Loc 1 | -1.46258 | 37.11436 |
| KES405 | Lukenya Road Loc 2a | -1.45539 | 37.12481 |
| KES406 | Lukenya Road Loc 3 | -1.45392 | 37.12608 |
| KES407 | Lukenya Road Loc 2b | -1.45528 | 37.12481 |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES408 | Kisanana Loc 1a | 0.03417 | 36.06406 |
| KES409 | Kisanana Loc 1b | 0.03417 | 36.06406 |
| KES410 | Kisanana Loc 1c | 0.03414 | 36.06433 |
| KES411 | Kabazi | -0.06900 | 36.14786 |
| KES412 | Olongai Loc 1 | -0.12706 | 36.01317 |
| KES413 | Olongai Loc 2 | -0.12669 | 36.01203 |
| KES414 | Olongai Loc 3 | -0.12706 | 36.01317 |
| KES415 | Arahuka | -0.15808 | 36.06094 |
| KES416 | NE Menengai Crater Loc 1a | -0.17681 | 36.10975 |
| KES417 | NE Menengai Crater Loc 1b | -0.17681 | 36.10975 |
| KES418 | Lokil | 0.84472 | 36.20244 |
| KES419 | Karau Loc 1 | 0.60475 | 36.18783 |
| KES420 | Karau Loc 2 | 0.60633 | 36.18544 |
| KES421 | Karau Loc 3 | 0.61217 | 36.18772 |
| KES422 | Karau Loc 4 | 0.61708 | 36.18756 |
| KES423 | Karau Loc 5 | 0.62150 | 36.18864 |
| KES424 | Chepungus (Paka) Loc 1a | 0.87758 | 36.18489 |
| KES425 | Chepungus (Paka) Loc 1b | 0.87758 | 36.18489 |
| KES426 | Chepungus (Paka) Loc 1c | 0.87758 | 36.18489 |
| KES427 | Chepungus (Paka) Loc 1d | 0.87758 | 36.18489 |
| KES428 | Chepungus (Paka) Loc 2 | 0.87819 | 36.17611 |
| KES429 | Lokoritabim | 0.68933 | 36.01178 |
| KES430 | Shin | 3.91656 | 36.47650 |
| KES431 | Naiyenareng Loc 1 | 2.41158 | 36.71647 |
| KES432 | Naiyenareng Loc 1 | 2.41158 | 36.71647 |
| KES433 | Naiyenareng Loc 1 | 2.41158 | 36.71647 |
| KES434 | Naiyenareng Loc 2 | 2.39636 | 36.72939 |
| KES435 | Naiyenareng Loc 2 | 2.39636 | 36.72939 |
| KES436 | Kalossi Loc 1 | 0.73669 | 36.13150 |
| KES437 | Kalossi Loc 2 | 0.73867 | 36.13647 |
| KES438 | Kalossi Loc 2 | 0.73867 | 36.13647 |
| KES439 | Cheptumkelek Loc 1 | 0.80239 | 36.18664 |
| KES440 | Cheptumkelek Loc 1 | 0.80239 | 36.18664 |
| KES441 | Cheptumkelek Loc 2 | 0.79564 | 36.18625 |
| KES442 | Silali Loc 1 | 1.16747 | 36.15197 |
| KES443 | Silali Loc 2 | 1.16647 | 36.15581 |
| KES444 | Kibelbel Loc 1 | 1.47778 | 36.14431 |
| KES445 | Kibelbel Loc 1 | 1.47778 | 36.14431 |
| KES446 | Kibelbel Loc 1 | 1.47778 | 36.14431 |
| KES447 | Kibelbel Loc 2 | 1.48311 | 36.14272 |
| KES448 | Kibelbel Loc 2 | 36.14272 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES449 | Lomi Loc 1a | 1.84692 | 36.28567 |
| KES450 | Lomi Loc 1a | 1.84692 | 36.28567 |
| KES451 | Lomi Loc 1b | 1.84697 | 36.28564 |
| KES452 | Lomi Loc 1b | 1.84697 | 36.28564 |
| KES453 | Lomi Loc 1b | 1.84697 | 36.28564 |
| KES454 | Lomi Loc 1c | 1.84714 | 36.28561 |
| KES455 | Lomi Loc 1c | 1.84714 | 36.28561 |
| KES456 | Lomi Loc 1c | 1.84714 | 36.28561 |
| KES457 | Lomi Loc 2 | 1.84736 | 36.28353 |
| KES458 | Lomi Loc 2 | 1.84736 | 36.28353 |
| KES459 | Kachalakwen | 1.09603 | 36.19736 |
| KES460 | Kachalakwen | 1.09603 | 36.19736 |
| KES461 | Kachalakwen | 1.09603 | 36.19736 |
| KES462 | Kachalakwen | 1.09603 | 36.19736 |
| KES463 | Alale Loc 1 | 1.13050 | 36.27958 |
| KES464 | Alale Loc 1 | 1.13050 | 36.27958 |
| KES465 | Alale Loc 1 | 1.13050 | 36.27958 |
| KES466 | Alale Loc 2 | 1.13253 | 36.27797 |
| KES467 | Nyakinywa Loc 1a | -0.58425 | 36.27667 |
| KES468 | Nyakinywa Loc 1b | -0.58425 | 36.27667 |
| KES469 | Nyakinywa Loc 2 | -0.58597 | 36.27483 |
| KES470 | Nyakinywa Loc 3 | -0.58647 | 36.27489 |
| KES471 | Jaika (Eburru Station) Loc 1a | -0.58628 | 36.26411 |
| KES472 | Jaika (Eburru Station) Loc 1b | -0.58628 | 36.26411 |
| KES473 | Jaika Loc 1c | -0.58650 | 36.26411 |
| KES474 | Jaika Loc 2 | -0.58733 | 36.26394 |
| KES475 | Jaika South Loc 1 | -0.58000 | 36.27322 |
| KES476 | Jaika South Loc 1 | -0.58000 | 36.27322 |
| KES477 | Jaika South Loc 2 | -0.57992 | 36.27292 |
| KES478 | Jaika South Loc 2 | -0.57992 | 36.27292 |
| KES479 | Ole Lorkumani (ole Enkapune) | -0.61794 | 36.18994 |
| KES480 | Ole Polos Loc. 1 (Ol Jorai) | -0.59972 | 36.21133 |
| KES481 | GsJj46 Rockshelter Loc 1 | -0.64764 | 36.32633 |
| KES482 | GsJj46 Rockshelter Loc 2 | -0.64836 | 36.32686 |
| KES483 | GsJj46 Rockshelter Loc 3 | -0.64817 | 36.32883 |
| KES484 | GsJj46 Rockshelter Loc 4 | -0.64644 | 36.32961 |
| KES485 | Marula Valley NE | -0.58756 | 36.32050 |
| KES486 | Marula Valley NE | -0.58756 | 36.32050 |
| KES487 | Nagum Loc 2a | -0.57722 | 36.30086 |
| KES488 | Nagum Loc 2b | -0.56533 | 36.30058 |
| KES489 | Nagum Loc 3 | 36.30942 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :--- | :--- | :---: | :---: |
|  |  |  |  |
| KES490 | GsJj84 Loc 1 | -0.54561 | 36.31392 |
| KES491 | Nagum pipeline Loc 1 | -0.54708 | 36.30608 |
| KES492 | Nagum pipeline Loc 2 | -0.54675 | 36.30575 |
| KES493 | Murai loc 1 | -0.54408 | 36.30128 |
| KES494 | Njeru Ofisini East | -0.56769 | 36.29978 |
| KES495 | Njeru Ofisini North | -0.57928 | 36.28989 |
| KES496 | Kilima Loc 1a (Kinangop) | -0.63358 | 36.48719 |
| KES497 | Kilima Loc 1a (Kinangop) | -0.63358 | 36.48719 |
| KES498 | Kilima Loc 1b (Kinangop) | -0.63353 | 36.48708 |
| KES499 | Kilima Loc 1b (Kinangop) | -0.63353 | 36.48708 |
| KES500 | Kilima Loc 1c (Kinangop) | -0.63333 | 36.48711 |
| KES501 | Kilima Loc 1c (Kinangop) | -0.63333 | 36.48711 |
| KES502 | Kilima Loc 1c (Kinangop) | -0.63333 | 36.48711 |
| KES503 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES504 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES505 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES506 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES507 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES508 | Lemudongo Loc 1 | -1.30433 | 35.97869 |
| KES509 | Ol Doinyo Nyokie West Loc 1 | -1.81092 | 36.35694 |
| KES510 | Ol Doinyo Nyokie West Loc 1 | -1.81092 | 36.35694 |
| KES511 | Ol Doinyo Nyokie West Loc 2a | -1.81506 | 36.35892 |
| KES512 | Ol Doinyo Nyokie West Loc 2a | -1.81506 | 36.35892 |
| KES513 | Ol Doinyo Nyokie West Loc 2b | -1.81553 | 36.35911 |
| KES514 | Ogata Rongai Loc 1 | -1.39075 | 36.72436 |
| KES515 | Ogata Rongai Loc 1 | -1.39075 | 36.72436 |
| KES516 | North Island Loc 1 | 4.05336 | 36.05781 |
| KES517 | North Island Loc 2 | 4.06472 | 36.04553 |
| KES518 | North Island Loc 2 | 4.06472 | 36.04553 |
| KES519 | North Island Loc 3 | 4.06542 | 36.04447 |
| KES520 | North Island Loc 4 | 4.06697 | 36.04331 |
| KES521 | North Island Loc 4 | 4.06697 | 36.04331 |
| KES522 | North Island Loc 5 | 4.06775 | 36.04458 |
| KES523 | North Island Loc 5 | 4.06775 | 36.04458 |
| KES524 | North Island Loc 6 | 4.06775 | 36.04542 |
| KES525 | North Island Loc 6 | 4.06775 | 36.04542 |
| KES526 | North Island Loc 7 | 4.06936 | 36.04569 |
| KES527 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES528 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES529 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES530 | Ol Doinyo Nyiru | 36.62275 |  |

Appendix 5 - continued

| ANID | Source Name | Coordinates |  |
| :---: | :--- | :---: | :---: |
|  |  | ${ }^{\circ} \mathbf{N}^{\circ}$ or ${ }^{\circ}$ S | ${ }^{\circ} \mathbf{E}$ |
| KES531 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES532 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES533 | Ol Doinyo Nyiru | -1.72064 | 36.62275 |
| KES534 | Museum Hill Bus Stage | -1.27325 | 36.81383 |
| KES535 | Museum Hill Bus Stage | -1.27325 | 36.81383 |
| KES536 | Kalusha Loc 1c (Lukenya) | -1.44781 | 37.10525 |
| KES537 | Lukenya Road Loc 1 | -1.46258 | 37.11436 |
| KES538 | Lukenya Road Loc 1 | -1.46258 | 37.11436 |
| KES539 | Lukenya Road Loc 2b | -1.45528 | 37.12481 |
| KES540 | Kabazi | -0.06900 | 36.14786 |
| KES541 | Chepungus (Paka) Loc 1b | 0.87758 | 36.18489 |
| KES542 | Chepungus (Paka) Loc 1b | 0.87758 | 36.18489 |

Appendix 6 - Concentration data from NAA for obsidian from Kenya. All concentrations are in parts per million ( ppm )

| Sample ID | Al | Cl | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES001 | 45718.3 | 1732.0 | 31.514 | 32253.8 | 1469.42 | 47281.4 | 0.0 |
| KES002 | 46702.8 | 1671.8 | 31.641 | 30100.0 | 1463.09 | 46994.0 | 0.0 |
| KES003 | 65231.8 | 856.9 | 16.288 | 40291.7 | 308.48 | 33476.0 | 0.0 |
| KES004 | 58297.1 | 808.9 | 15.609 | 41032.3 | 329.37 | 33323.8 | 46.9 |
| KES005 | 39670.8 | 2419.9 | 61.424 | 40559.7 | 1641.97 | 50158.8 | 0.0 |
| KES006 | 43875.0 | 2631.3 | 72.075 | 34024.9 | 1841.18 | 56668.0 | 143.3 |
| KES007 | 51415.2 | 2555.6 | 66.552 | 37538.3 | 1735.53 | 51982.8 | 0.0 |
| KES008 | 32159.0 | 1564.4 | 42.702 | 37305.6 | 2118.47 | 52009.6 | 77.0 |
| KES009 | 41756.4 | 1700.3 | 43.911 | 38894.5 | 2129.80 | 52392.2 | 0.0 |
| KES010 | 41613.4 | 2620.1 | 68.255 | 44636.1 | 1729.27 | 51845.2 | 0.0 |
| KES011 | 41241.4 | 2715.8 | 71.538 | 34333.6 | 1817.00 | 53605.4 | 0.0 |
| KES012 | 42359.4 | 2020.4 | 54.670 | 40423.0 | 1710.06 | 46633.9 | 0.0 |
| KES013 | 44079.7 | 1868.1 | 42.318 | 34033.8 | 1912.42 | 49868.7 | 152.3 |
| KES014 | 37539.0 | 1135.7 | 27.787 | 39046.1 | 2169.04 | 38244.8 | 0.0 |
| KES015 | 38937.4 | 1262.8 | 28.227 | 42509.6 | 2269.72 | 47645.1 | 0.0 |
| KES016 | 75639.7 | 634.9 | 14.523 | 36943.9 | 1044.30 | 36808.4 | 238.6 |
| KES017 | 71351.4 | 618.9 | 14.115 | 42795.4 | 1074.79 | 37551.1 | 288.5 |
| KES018 | 69790.0 | 536.8 | 13.651 | 40691.0 | 1113.40 | 37122.9 | 253.4 |
| KES019 | 73470.5 | 582.5 | 14.727 | 41953.8 | 1077.95 | 37700.7 | 266.7 |
| KES020 | 84152.2 | 813.9 | 16.850 | 46228.3 | 2613.64 | 63592.6 | 331.2 |
| KES021 | 54726.4 | 1866.0 | 31.596 | 36272.1 | 492.64 | 39978.4 | 79.3 |
| KES022 | 53729.2 | 1734.2 | 32.685 | 39852.6 | 519.28 | 39751.2 | 72.2 |
| KES023 | 52903.3 | 1808.7 | 33.490 | 33491.4 | 478.62 | 40299.2 | 84.9 |
| KES024 | 52448.0 | 1804.4 | 31.784 | 34476.5 | 479.47 | 39846.8 | 37.2 |
| KES025 | 55696.1 | 1956.8 | 34.424 | 42345.4 | 481.21 | 40717.7 | 62.4 |
| KES026 | 58010.1 | 1888.6 | 34.732 | 37620.8 | 471.61 | 40023.6 | 43.9 |
| KES027 | 63432.3 | 1884.9 | 34.048 | 37097.8 | 481.74 | 40200.2 | 86.5 |
| KES028 | 56365.7 | 1850.0 | 35.048 | 36824.6 | 480.87 | 40372.4 | 53.3 |
| KES029 | 57422.0 | 1855.4 | 34.493 | 39806.3 | 472.92 | 40420.7 | 37.5 |
| KES030 | 59229.5 | 1809.2 | 33.803 | 35880.0 | 460.56 | 39027.5 | 63.4 |
| KES031 | 51614.6 | 1905.4 | 34.835 | 36966.6 | 466.97 | 40035.1 | 72.0 |
| KES032 | 59077.0 | 1861.3 | 34.384 | 35940.7 | 471.59 | 40592.4 | 66.4 |
| KES033 | 41869.3 | 2097.6 | 44.798 | 35745.6 | 1899.03 | 49342.0 | 0.0 |
| KES034 | 51611.8 | 1897.2 | 44.948 | 32984.8 | 1903.06 | 49802.3 | 0.0 |
| KES035 | 39651.0 | 1929.5 | 46.765 | 36918.2 | 1888.54 | 48599.9 | 123.7 |
| KES036 | 36731.0 | 1986.1 | 45.331 | 36599.0 | 1904.55 | 49068.3 | 142.7 |
| KES037 | 39110.1 | 2010.4 | 45.329 | 40888.0 | 1870.07 | 48887.5 | 167.6 |
| KES038 | 49170.4 | 1987.1 | 44.348 | 37817.6 | 1882.25 | 48726.3 | 64.1 |
| KES039 | 44806.6 | 1972.7 | 44.551 | 36917.6 | 1888.25 | 48965.2 | 100.7 |
| KES040 | 37003.9 | 2080.1 | 44.717 | 34232.7 | 1897.57 | 49637.3 | 116.9 |
| KES042 | 43439.8 | 2329.0 | 44.916 | 35107.3 | 2339.91 | 53468.8 | 0.0 |
| KES043 | 37605.1 | 1704.5 | 47.541 | 32052.2 | 2122.24 | 51716.9 | 0.0 |
| KES044 | 34636.8 | 1666.6 | 45.995 | 35858.2 | 2150.08 | 52517.9 | 52.9 |
| KES045 | 38781.7 | 1962.4 | 49.283 | 37950.6 | 2151.78 | 52482.2 | 109.1 |
| KES046 | 34054.6 | 1679.0 | 45.692 | 42284.9 | 1985.61 | 50412.3 | 97.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES047 | 52526.5 | 1935.9 | 35.008 | 36734.6 | 471.96 | 39893.4 | 56.0 |
| KES048 | 53937.9 | 1810.1 | 32.006 | 38865.2 | 458.56 | 38941.0 | 53.4 |
| KES049 | 53158.3 | 1859.8 | 33.089 | 36856.9 | 465.62 | 39767.4 | 54.3 |
| KES050 | 50389.3 | 1704.7 | 34.937 | 36063.8 | 439.15 | 38911.6 | 67.4 |
| KES051 | 53849.1 | 1739.0 | 36.216 | 37814.5 | 443.15 | 39063.3 | 67.4 |
| KES052 | 50299.0 | 1900.0 | 40.025 | 34242.3 | 519.53 | 40983.6 | 74.3 |
| KES053 | 53291.3 | 1962.0 | 40.954 | 31740.6 | 521.33 | 40741.3 | 62.9 |
| KES054 | 46669.3 | 1543.4 | 44.626 | 37611.8 | 2228.87 | 55853.8 | 0.0 |
| KES055 | 44578.5 | 1457.6 | 44.116 | 39050.4 | 2226.31 | 56193.1 | 0.0 |
| KES056 | 40082.5 | 1445.9 | 42.727 | 33544.5 | 2203.83 | 54424.3 | 0.0 |
| KES057 | 60975.2 | 843.3 | 23.064 | 38287.1 | 2794.18 | 57212.7 | 0.0 |
| KES058 | 43803.8 | 2175.8 | 56.755 | 37326.0 | 1586.83 | 47992.7 | 83.5 |
| KES059 | 37459.9 | 2371.8 | 70.272 | 32766.2 | 1853.10 | 56358.9 | 135.8 |
| KES060 | 48217.1 | 1474.1 | 47.477 | 38385.5 | 1319.50 | 44907.5 | 74.3 |
| KES061 | 48097.5 | 1513.4 | 47.312 | 36185.6 | 1321.21 | 44837.8 | 71.2 |
| KES062 | 37661.9 | 2107.8 | 51.933 | 36865.7 | 1685.13 | 45967.0 | 105.3 |
| KES063 | 41879.5 | 2095.7 | 53.500 | 36544.9 | 1675.14 | 47911.6 | 0.0 |
| KES064 | 57104.7 | 2036.6 | 54.219 | 39950.3 | 1767.06 | 48015.2 | 0.0 |
| KES065 | 39805.4 | 2024.2 | 55.851 | 33814.3 | 1644.44 | 47780.4 | 0.0 |
| KES066 | 67186.3 | 694.0 | 18.264 | 35529.9 | 2441.42 | 51363.9 | 62.6 |
| KES067 | 49440.3 | 1954.1 | 54.598 | 36765.9 | 1555.49 | 47251.9 | 0.0 |
| KES068 | 42113.9 | 2014.1 | 54.467 | 34216.5 | 1545.82 | 47126.8 | 0.0 |
| KES069 | 45716.7 | 2062.2 | 54.730 | 39083.9 | 1530.45 | 46506.8 | 0.0 |
| KES070 | 41395.5 | 2357.9 | 67.541 | 34490.5 | 1842.17 | 55586.2 | 0.0 |
| KES071 | 37160.8 | 2445.7 | 68.083 | 29866.3 | 1827.47 | 55657.5 | 0.0 |
| KES072 | 42654.3 | 2436.6 | 71.221 | 31538.3 | 1859.69 | 57218.4 | 0.0 |
| KES073 | 44779.8 | 2435.7 | 66.079 | 34700.9 | 1798.68 | 54835.8 | 90.3 |
| KES074 | 42779.3 | 2355.8 | 69.519 | 29086.9 | 1825.13 | 55783.0 | 0.0 |
| KES075 | 56615.8 | 1790.6 | 38.528 | 32418.5 | 397.83 | 39818.2 | 85.1 |
| KES076 | 69857.7 | 854.0 | 15.870 | 36689.0 | 498.04 | 32866.9 | 0.0 |
| KES077 | 55401.9 | 805.3 | 15.561 | 36066.6 | 293.91 | 32157.5 | 35.4 |
| KES078 | 58981.8 | 820.8 | 14.766 | 36439.5 | 302.64 | 32351.2 | 43.4 |
| KES079 | 40276.3 | 2232.9 | 58.938 | 33294.8 | 1623.58 | 48866.9 | 0.0 |
| KES080 | 59891.2 | 757.1 | 20.883 | 31512.9 | 2547.39 | 51863.3 | 0.0 |
| KES081 | 47956.3 | 1632.1 | 45.055 | 35060.7 | 1891.95 | 49214.5 | 138.4 |
| KES082 | 48842.7 | 1755.9 | 44.569 | 41107.3 | 1913.35 | 50405.2 | 133.4 |
| KES083 | 53242.4 | 1856.4 | 43.826 | 42029.4 | 1908.89 | 49815.5 | 147.2 |
| KES084 | 47083.3 | 1834.3 | 44.339 | 35852.1 | 1882.32 | 49438.7 | 83.5 |
| KES085 | 50163.8 | 1716.1 | 45.294 | 35521.3 | 1914.23 | 50494.6 | 99.3 |
| KES086 | 50421.0 | 1888.6 | 45.375 | 34858.2 | 1910.26 | 49816.5 | 100.5 |
| KES087 | 43861.1 | 2090.1 | 66.180 | 30407.3 | 1654.77 | 49762.2 | 102.4 |
| KES088 | 43539.5 | 2143.6 | 66.720 | 32675.2 | 1815.83 | 49952.9 | 0.0 |
| KES089 | 46922.6 | 1265.5 | 28.526 | 37602.0 | 2144.13 | 55764.2 | 0.0 |
| KES090 | 43213.8 | 1173.8 | 30.807 | 37136.3 | 2174.57 | 56669.1 | 0.0 |
| KES091 | 44762.2 | 1803.5 | 53.198 | 32713.1 | 1679.57 | 45474.9 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES092 | 41491.3 | 1845.7 | 54.524 | 35581.5 | 1691.04 | 46302.8 | 0.0 |
| KES093 | 49952.7 | 1828.5 | 54.086 | 38458.6 | 1709.09 | 46439.3 | 0.0 |
| KES094 | 44732.4 | 1699.1 | 54.283 | 38117.8 | 1696.49 | 45687.4 | 0.0 |
| KES095 | 45100.4 | 1806.0 | 53.319 | 38696.9 | 1522.74 | 46712.5 | 0.0 |
| KES096 | 47213.6 | 1830.7 | 54.816 | 34402.9 | 1513.72 | 46025.5 | 0.0 |
| KES097 | 44180.4 | 1798.4 | 54.428 | 38261.0 | 1526.15 | 46217.8 | 0.0 |
| KES098 | 43418.9 | 1044.8 | 29.783 | 34104.5 | 2228.96 | 46342.0 | 0.0 |
| KES099 | 48208.5 | 992.9 | 28.753 | 36790.7 | 2232.40 | 46844.1 | 0.0 |
| KES100 | 42487.2 | 1139.8 | 27.519 | 33403.7 | 2221.26 | 46475.2 | 0.0 |
| KES101 | 68631.5 | 675.7 | 19.100 | 38045.6 | 2512.42 | 48846.7 | 131.1 |
| KES102 | 68669.7 | 720.5 | 18.741 | 37716.7 | 2171.96 | 47829.4 | 0.0 |
| KES103 | 79001.4 | 704.9 | 19.090 | 37497.9 | 2106.95 | 47674.1 | 144.4 |
| KES104 | 70034.2 | 1055.2 | 18.078 | 36773.4 | 2264.08 | 47773.1 | 0.0 |
| KES124 | 63006.4 | 643.3 | 17.406 | 40877.8 | 2276.74 | 39976.3 | 91.1 |
| KES125 | 63545.0 | 397.0 | 19.730 | 50558.1 | 2233.73 | 31566.5 | 104.8 |
| KES126 | 71872.1 | 686.8 | 18.698 | 39770.3 | 2450.26 | 52459.5 | 99.4 |
| KES127 | 64036.1 | 481.8 | 18.613 | 38809.3 | 2392.27 | 40717.5 | 0.0 |
| KES128 | 62212.1 | 457.1 | 16.160 | 48218.0 | 2360.32 | 27059.3 | 0.0 |
| KES130 | 69824.6 | 669.8 | 18.004 | 38652.6 | 2421.60 | 51460.0 | 102.6 |
| KES131 | 68159.7 | 686.1 | 20.393 | 41915.9 | 2442.67 | 51515.7 | 0.0 |
| KES132 | 65064.9 | 610.1 | 17.793 | 40461.3 | 2396.69 | 51837.0 | 100.3 |
| KES133 | 92386.2 | 924.3 | 16.222 | 38592.6 | 2667.44 | 63948.9 | 329.4 |
| KES134 | 84898.7 | 825.7 | 15.646 | 41159.4 | 2660.60 | 60886.3 | 310.3 |
| KES135 | 75720.8 | 773.0 | 15.880 | 39446.8 | 2648.83 | 64148.8 | 388.9 |
| KES136 | 84437.1 | 879.5 | 16.219 | 38547.2 | 2585.28 | 63392.6 | 338.7 |
| KES137 | 75582.0 | 793.7 | 18.395 | 36584.2 | 2745.60 | 63823.2 | 0.0 |
| KES138 | 75140.0 | 748.3 | 21.407 | 43781.9 | 2763.67 | 64145.8 | 67.2 |
| KES139 | 83035.5 | 848.9 | 16.127 | 37094.8 | 2649.46 | 63709.3 | 304.3 |
| KES140 | 83564.3 | 923.5 | 15.180 | 41037.5 | 2610.74 | 63927.6 | 313.2 |
| KES141 | 71310.9 | 978.1 | 18.134 | 39186.5 | 2651.54 | 65377.3 | 302.5 |
| KES147 | 52203.3 | 1780.9 | 39.865 | 32867.9 | 398.65 | 39953.5 | 76.4 |
| KES148 | 57228.0 | 1663.1 | 39.741 | 34702.8 | 397.97 | 39770.7 | 80.8 |
| KES149 | 56564.4 | 1667.4 | 39.110 | 34994.4 | 395.28 | 39930.5 | 58.7 |
| KES150 | 54147.9 | 1761.8 | 39.221 | 40466.9 | 406.29 | 40482.7 | 67.6 |
| KES151 | 56216.8 | 1742.6 | 38.925 | 36157.4 | 397.03 | 39994.3 | 65.6 |
| KES152 | 53377.3 | 1663.6 | 38.757 | 39381.6 | 393.87 | 39560.5 | 85.1 |
| KES153 | 52381.6 | 1645.8 | 40.472 | 33682.1 | 392.96 | 39697.0 | 0.0 |
| KES154 | 55818.7 | 1684.3 | 37.721 | 35125.7 | 392.69 | 39633.5 | 64.3 |
| KES155 | 55377.5 | 1661.1 | 39.382 | 39530.3 | 393.95 | 39623.4 | 86.8 |
| KES156 | 59217.8 | 1693.3 | 38.648 | 34656.9 | 390.42 | 39205.4 | 0.0 |
| KES157 | 53790.7 | 1711.5 | 39.783 | 33522.2 | 401.61 | 40413.9 | 0.0 |
| KES158 | 55324.8 | 1648.0 | 37.500 | 36794.9 | 392.67 | 39580.3 | 67.3 |
| KES159 | 58449.5 | 1594.2 | 38.747 | 38603.4 | 397.20 | 39811.1 | 0.0 |
| KES160 | 51833.2 | 1758.7 | 39.389 | 35550.5 | 401.94 | 40489.0 | 45.4 |
| KES161 | 54458.5 | 1699.6 | 40.812 | 36400.0 | 407.50 | 40745.3 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES162 | 56663.8 | 1643.1 | 37.919 | 34663.1 | 394.80 | 39483.0 | 96.7 |
| KES163 | 58566.1 | 1747.5 | 39.282 | 35265.5 | 396.12 | 39735.0 | 98.1 |
| KES164 | 61296.1 | 1723.1 | 40.923 | 36948.0 | 405.50 | 40585.9 | 81.7 |
| KES165 | 34000.1 | 2234.5 | 41.606 | 31899.4 | 2241.82 | 53731.3 | 0.0 |
| KES166 | 37804.8 | 2302.3 | 42.566 | 31651.6 | 2195.80 | 52399.6 | 0.0 |
| KES167 | 34362.9 | 2302.1 | 44.795 | 31064.3 | 2276.75 | 53339.4 | 0.0 |
| KES168 | 39158.4 | 1212.3 | 35.761 | 29331.6 | 2100.46 | 51068.7 | 0.0 |
| KES169 | 58975.3 | 1138.8 | 26.436 | 36518.3 | 1322.91 | 46506.9 | 0.0 |
| KES170 | 54507.6 | 833.3 | 28.166 | 38061.3 | 1157.84 | 42456.7 | 0.0 |
| KES171 | 54380.1 | 904.1 | 28.344 | 39552.4 | 1239.71 | 42695.2 | 0.0 |
| KES172 | 54515.4 | 899.5 | 28.344 | 36610.1 | 1162.33 | 42421.6 | 0.0 |
| KES173 | 59547.9 | 825.8 | 28.904 | 42635.7 | 1199.49 | 43249.5 | 0.0 |
| KES174 | 55322.0 | 832.5 | 28.945 | 41715.8 | 1149.14 | 42736.9 | 0.0 |
| KES175 | 53549.3 | 846.2 | 28.101 | 37471.5 | 1096.43 | 41621.5 | 0.0 |
| KES176 | 62597.4 | 899.1 | 15.737 | 39100.3 | 305.79 | 34203.7 | 0.0 |
| KES177 | 66206.4 | 945.8 | 16.659 | 36913.8 | 305.84 | 33714.6 | 0.0 |
| KES178 | 64894.2 | 911.4 | 16.378 | 39468.5 | 311.93 | 35143.7 | 70.0 |
| KES179 | 66551.8 | 879.3 | 15.770 | 40329.7 | 298.96 | 33308.9 | 0.0 |
| KES180 | 58940.0 | 840.9 | 15.767 | 41602.9 | 308.99 | 33781.9 | 0.0 |
| KES181 | 62186.8 | 810.0 | 16.319 | 40650.9 | 314.63 | 34471.5 | 81.1 |
| KES182 | 41957.2 | 1993.6 | 56.620 | 40091.9 | 1637.21 | 48557.3 | 156.0 |
| KES183 | 49448.2 | 2058.2 | 57.592 | 37808.3 | 1641.38 | 48728.4 | 0.0 |
| KES184 | 42188.8 | 1957.9 | 55.698 | 35260.2 | 1601.11 | 47148.6 | 0.0 |
| KES185 | 40317.8 | 1985.8 | 56.977 | 43949.6 | 1630.11 | 48288.0 | 0.0 |
| KES186 | 40244.9 | 1929.3 | 54.548 | 36227.8 | 1571.23 | 46507.4 | 0.0 |
| KES187 | 40981.0 | 1974.5 | 54.310 | 35464.3 | 1577.76 | 47000.9 | 0.0 |
| KES188 | 38498.9 | 1938.1 | 51.937 | 38833.9 | 1602.34 | 47433.6 | 0.0 |
| KES189 | 41572.3 | 2037.2 | 54.514 | 40533.2 | 1593.09 | 47500.1 | 0.0 |
| KES190 | 43199.7 | 2005.5 | 57.200 | 38492.3 | 1619.29 | 47726.9 | 115.6 |
| KES191 | 41279.1 | 2005.7 | 55.785 | 39928.3 | 1603.07 | 47398.5 | 0.0 |
| KES192 | 38737.1 | 1988.4 | 56.271 | 37612.6 | 1629.41 | 47333.3 | 0.0 |
| KES193 | 42167.4 | 2081.0 | 56.293 | 42228.7 | 1638.27 | 47641.1 | 0.0 |
| KES194 | 42391.2 | 1988.3 | 55.705 | 37640.7 | 1612.69 | 46963.3 | 0.0 |
| KES195 | 47586.1 | 2076.1 | 57.666 | 45612.8 | 1649.98 | 48964.1 | 0.0 |
| KES196 | 45727.6 | 1999.9 | 56.383 | 40076.4 | 1607.06 | 47415.5 | 0.0 |
| KES197 | 43538.1 | 1961.4 | 56.152 | 39527.8 | 1592.99 | 47119.2 | 0.0 |
| KES198 | 45622.0 | 1961.5 | 55.242 | 36561.0 | 1573.73 | 46580.4 | 0.0 |
| KES199 | 48771.8 | 1944.3 | 54.607 | 38999.3 | 1568.71 | 46529.9 | 0.0 |
| KES200 | 38036.6 | 2012.1 | 56.454 | 37532.5 | 1616.74 | 47684.5 | 0.0 |
| KES201 | 37502.7 | 2040.7 | 55.925 | 38674.8 | 1593.66 | 47187.5 | 0.0 |
| KES202 | 38672.1 | 2128.9 | 56.684 | 40935.9 | 1638.32 | 48432.9 | 0.0 |
| KES203 | 45055.1 | 1977.9 | 54.956 | 36104.6 | 1577.57 | 46934.5 | 0.0 |
| KES204 | 38605.4 | 1979.2 | 54.914 | 41868.2 | 1595.86 | 47286.3 | 0.0 |
| KES205 | 40703.6 | 2003.4 | 56.269 | 44278.2 | 1629.20 | 47538.1 | 0.0 |
| KES206 | 44960.8 | 1960.9 | 57.719 | 38206.5 | 1617.31 | 47889.1 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES207 | 43601.2 | 1944.4 | 54.483 | 35984.7 | 1577.41 | 46876.8 | 0.0 |
| KES208 | 47597.6 | 1967.1 | 55.130 | 31270.3 | 1581.28 | 46395.5 | 0.0 |
| KES209 | 45127.0 | 1985.7 | 50.211 | 36931.4 | 1609.28 | 47372.0 | 0.0 |
| KES210 | 47285.4 | 1931.4 | 56.235 | 35226.1 | 1637.90 | 48055.7 | 0.0 |
| KES211 | 43423.2 | 1829.7 | 55.182 | 36469.3 | 1594.65 | 46959.7 | 0.0 |
| KES212 | 44862.9 | 2048.8 | 56.631 | 33985.9 | 1618.64 | 47881.2 | 0.0 |
| KES213 | 41699.3 | 1985.7 | 56.131 | 39721.4 | 1632.97 | 47603.8 | 0.0 |
| KES214 | 43520.8 | 1936.0 | 54.792 | 38027.8 | 1608.72 | 47336.3 | 0.0 |
| KES215 | 44485.0 | 2034.5 | 56.080 | 44520.7 | 1625.65 | 47637.9 | 0.0 |
| KES216 | 36486.7 | 2136.2 | 56.709 | 38285.9 | 1610.19 | 48172.4 | 0.0 |
| KES217 | 43787.7 | 2096.9 | 57.623 | 40583.6 | 1601.49 | 48310.2 | 0.0 |
| KES218 | 57605.3 | 1798.2 | 40.249 | 39049.6 | 396.30 | 39861.4 | 0.0 |
| KES219 | 59171.7 | 1792.8 | 38.696 | 33961.6 | 386.79 | 39458.6 | 84.3 |
| KES220 | 45426.7 | 2108.3 | 56.599 | 31650.2 | 1575.41 | 47668.2 | 0.0 |
| KES221 | 43757.1 | 2106.3 | 56.672 | 35573.9 | 1606.17 | 47723.2 | 0.0 |
| KES222 | 40323.0 | 2136.4 | 56.039 | 39978.7 | 1600.75 | 48158.7 | 0.0 |
| KES223 | 42179.5 | 2125.2 | 53.077 | 32935.1 | 1572.76 | 46825.8 | 0.0 |
| KES224 | 42589.4 | 2152.1 | 54.566 | 34412.0 | 1583.99 | 47463.5 | 0.0 |
| KES225 | 43484.9 | 2107.2 | 54.829 | 36900.7 | 1584.64 | 47134.2 | 0.0 |
| KES226 | 46907.1 | 2091.7 | 55.343 | 34173.9 | 1575.24 | 47334.4 | 0.0 |
| KES227 | 45841.6 | 2094.6 | 53.651 | 42465.3 | 1560.47 | 46590.8 | 0.0 |
| KES228 | 42118.3 | 2150.1 | 56.697 | 36504.8 | 1584.99 | 47704.7 | 0.0 |
| KES229 | 42261.4 | 1964.8 | 54.295 | 36748.1 | 1535.24 | 46453.7 | 0.0 |
| KES230 | 44116.6 | 2029.1 | 55.550 | 35642.7 | 1567.94 | 46601.0 | 0.0 |
| KES231 | 52171.3 | 2315.6 | 59.000 | 42497.2 | 1660.56 | 50117.1 | 0.0 |
| KES232 | 40533.0 | 2195.7 | 56.077 | 32971.9 | 1610.58 | 47729.2 | 0.0 |
| KES233 | 39966.0 | 2093.4 | 56.399 | 34336.0 | 1608.37 | 48019.4 | 99.8 |
| KES234 | 41935.5 | 1993.0 | 55.200 | 38379.7 | 1620.96 | 47609.2 | 0.0 |
| KES235 | 46672.4 | 2118.0 | 56.761 | 34233.3 | 1602.55 | 47530.1 | 0.0 |
| KES236 | 46255.4 | 2073.1 | 56.076 | 35456.0 | 1622.26 | 47537.1 | 0.0 |
| KES237 | 47078.5 | 2035.6 | 55.419 | 41446.1 | 1614.21 | 47253.4 | 102.6 |
| KES238 | 44527.3 | 2020.3 | 55.894 | 37022.3 | 1605.04 | 47554.2 | 0.0 |
| KES239 | 42869.6 | 2095.0 | 56.182 | 37925.2 | 1593.75 | 47033.9 | 0.0 |
| KES240 | 47192.2 | 2045.0 | 55.323 | 32522.0 | 1739.90 | 46914.9 | 0.0 |
| KES241 | 41717.5 | 1945.7 | 55.480 | 35451.5 | 1600.92 | 47233.3 | 0.0 |
| KES242 | 42682.9 | 2000.3 | 56.526 | 35643.0 | 1642.69 | 47969.6 | 0.0 |
| KES243 | 61426.5 | 1884.1 | 34.123 | 40157.4 | 478.08 | 40784.1 | 57.5 |
| KES244 | 53316.1 | 1853.9 | 33.269 | 36576.2 | 465.55 | 40037.8 | 118.7 |
| KES245 | 57156.4 | 1789.9 | 30.668 | 41447.6 | 487.11 | 40500.5 | 110.3 |
| KES246 | 56539.4 | 1927.1 | 33.359 | 38948.1 | 482.24 | 40468.6 | 70.0 |
| KES247 | 53455.1 | 1749.9 | 33.021 | 35033.9 | 497.30 | 39850.5 | 78.7 |
| KES248 | 55727.9 | 1765.8 | 34.320 | 38092.7 | 509.68 | 41127.4 | 0.0 |
| KES249 | 56684.1 | 1864.8 | 33.704 | 37334.0 | 546.29 | 41071.3 | 138.7 |
| KES250 | 52925.2 | 1775.9 | 32.131 | 38879.8 | 479.67 | 40285.0 | 0.0 |
| KES251 | 56232.1 | 1650.8 | 30.572 | 38057.5 | 473.35 | 38478.3 | 82.6 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES252 | 49696.1 | 1899.3 | 33.048 | 40946.8 | 475.67 | 40268.5 | 100.6 |
| KES253 | 51625.1 | 1894.4 | 33.656 | 37295.1 | 469.65 | 40680.1 | 77.5 |
| KES254 | 55265.2 | 1845.0 | 34.078 | 39758.0 | 479.42 | 40943.9 | 0.0 |
| KES255 | 67307.8 | 617.4 | 15.421 | 40931.5 | 1099.78 | 37841.6 | 235.2 |
| KES256 | 74116.6 | 641.1 | 13.252 | 40378.0 | 1101.89 | 38187.6 | 225.7 |
| KES257 | 71345.9 | 643.8 | 15.099 | 39927.4 | 1102.12 | 38597.5 | 219.5 |
| KES258 | 59770.9 | 572.4 | 14.803 | 40691.2 | 1142.78 | 37530.7 | 261.6 |
| KES259 | 72332.0 | 652.9 | 15.200 | 38560.8 | 1079.86 | 37875.6 | 300.7 |
| KES260 | 66802.9 | 648.3 | 15.407 | 41658.1 | 1087.34 | 38079.4 | 254.5 |
| KES261 | 67182.2 | 680.1 | 15.832 | 47214.0 | 1107.38 | 37818.6 | 287.3 |
| KES262 | 66078.1 | 672.2 | 15.254 | 43491.5 | 1117.05 | 37903.7 | 275.9 |
| KES263 | 75185.4 | 691.6 | 15.853 | 48581.4 | 1169.21 | 40306.4 | 286.4 |
| KES264 | 72690.7 | 720.5 | 15.662 | 40990.9 | 1155.46 | 38512.2 | 243.0 |
| KES265 | 71742.7 | 540.2 | 15.603 | 53240.9 | 1124.67 | 38750.6 | 286.5 |
| KES266 | 68507.3 | 612.6 | 16.253 | 41898.4 | 1114.20 | 38801.3 | 305.4 |
| KES267 | 68549.2 | 579.3 | 15.923 | 48003.9 | 1146.30 | 39154.1 | 295.3 |
| KES268 | 74068.9 | 648.2 | 15.223 | 42682.3 | 1096.01 | 38596.5 | 249.6 |
| KES269 | 66945.6 | 567.0 | 16.147 | 42065.2 | 1121.47 | 37257.2 | 264.7 |
| KES270 | 65093.1 | 640.4 | 15.795 | 44705.8 | 1092.57 | 38063.4 | 290.4 |
| KES271 | 74482.2 | 574.5 | 15.086 | 41585.6 | 1083.91 | 37381.8 | 322.7 |
| KES272 | 74893.6 | 665.6 | 14.926 | 39924.5 | 1085.39 | 37683.8 | 276.0 |
| KES273 | 69540.3 | 631.1 | 15.542 | 46779.5 | 1087.24 | 38240.8 | 261.4 |
| KES274 | 70987.6 | 587.7 | 15.726 | 41807.8 | 1118.93 | 37979.9 | 313.4 |
| KES275 | 69561.2 | 690.6 | 14.865 | 41111.6 | 1093.27 | 37215.0 | 245.7 |
| KES276 | 71382.2 | 661.4 | 15.244 | 41427.4 | 1066.95 | 37565.8 | 217.8 |
| KES277 | 72450.5 | 671.2 | 15.351 | 43650.8 | 1091.71 | 37199.2 | 201.4 |
| KES278 | 55323.9 | 1898.6 | 31.948 | 38356.4 | 475.79 | 39573.5 | 107.6 |
| KES279 | 59653.3 | 1832.4 | 33.410 | 38373.4 | 464.08 | 39680.9 | 0.0 |
| KES281 | 54519.8 | 1714.1 | 31.947 | 34189.4 | 515.60 | 38657.0 | 131.6 |
| KES282 | 56651.3 | 1781.7 | 31.382 | 37863.9 | 478.95 | 38810.6 | 0.0 |
| KES283 | 60305.7 | 1770.1 | 30.709 | 35427.1 | 497.43 | 39430.0 | 0.0 |
| KES284 | 54531.4 | 1740.1 | 30.457 | 37485.1 | 500.76 | 38709.5 | 0.0 |
| KES285 | 72045.5 | 854.0 | 24.672 | 40306.7 | 2803.98 | 55296.1 | 0.0 |
| KES286 | 67442.3 | 824.7 | 23.161 | 34560.1 | 2833.69 | 55469.6 | 0.0 |
| KES287 | 69618.8 | 1161.4 | 21.722 | 33737.8 | 2919.62 | 61687.5 | 0.0 |
| KES288 | 59920.7 | 1175.9 | 22.158 | 35496.3 | 2850.04 | 56718.7 | 0.0 |
| KES289 | 61956.3 | 1331.6 | 21.411 | 40013.8 | 2942.94 | 60832.0 | 0.0 |
| KES290 | 66838.8 | 1228.9 | 21.411 | 37915.8 | 2945.03 | 61785.6 | 0.0 |
| KES291 | 67278.0 | 1405.7 | 19.443 | 40345.9 | 2858.08 | 59986.7 | 0.0 |
| KES292 | 58752.1 | 1255.7 | 19.126 | 43155.1 | 2880.12 | 60464.3 | 0.0 |
| KES293 | 62556.1 | 1336.1 | 20.980 | 39011.7 | 2918.54 | 61573.9 | 0.0 |
| KES294 | 66043.0 | 1235.8 | 22.183 | 33491.0 | 2928.91 | 61693.0 | 0.0 |
| KES295 | 63877.3 | 1360.0 | 21.979 | 38749.8 | 2913.10 | 61453.4 | 46.6 |
| KES296 | 74039.7 | 1269.1 | 21.145 | 40606.3 | 2928.09 | 61047.6 | 0.0 |
| KES297 | 55775.3 | 1872.8 | 41.147 | 38693.7 | 407.41 | 39920.0 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES298 | 59869.7 | 1808.4 | 40.858 | 37138.3 | 397.27 | 39360.5 | 60.8 |
| KES299 | 56607.9 | 1839.1 | 39.764 | 35781.4 | 397.77 | 39953.1 | 89.4 |
| KES300 | 55437.8 | 2477.0 | 68.205 | 30506.6 | 1670.85 | 49363.9 | 264.9 |
| KES301 | 38358.8 | 2404.1 | 64.522 | 35761.3 | 1644.29 | 49189.8 | 0.0 |
| KES302 | 42134.8 | 2412.1 | 64.189 | 28687.0 | 1615.72 | 48523.3 | 244.7 |
| KES303 | 45983.7 | 2439.3 | 64.573 | 33753.3 | 1644.46 | 48631.2 | 209.3 |
| KES304 | 40775.7 | 2405.0 | 64.093 | 37749.4 | 1650.65 | 48849.3 | 0.0 |
| KES305 | 43190.6 | 2475.7 | 64.901 | 33993.2 | 1620.28 | 48680.4 | 278.9 |
| KES306 | 42616.2 | 2325.0 | 65.318 | 42688.4 | 1644.83 | 49283.7 | 221.1 |
| KES307 | 44173.7 | 2311.6 | 65.926 | 33253.9 | 1612.37 | 48930.8 | 170.0 |
| KES308 | 43254.3 | 2007.0 | 68.086 | 33360.8 | 1660.85 | 49551.9 | 167.0 |
| KES309 | 43137.5 | 2067.0 | 67.015 | 34294.6 | 1585.55 | 48137.2 | 0.0 |
| KES310 | 44127.0 | 1984.2 | 68.130 | 30999.4 | 1575.09 | 48782.0 | 0.0 |
| KES311 | 47643.6 | 2075.1 | 71.228 | 35512.3 | 1723.79 | 50596.5 | 0.0 |
| KES312 | 39583.8 | 2086.2 | 68.493 | 34497.5 | 1671.22 | 49307.0 | 225.2 |
| KES313 | 40631.6 | 2045.4 | 69.486 | 36028.7 | 1677.72 | 49045.4 | 0.0 |
| KES314 | 40831.4 | 2083.5 | 70.875 | 35537.3 | 1709.51 | 50000.2 | 0.0 |
| KES315 | 43148.0 | 2070.7 | 66.190 | 39064.5 | 1676.93 | 49082.4 | 0.0 |
| KES316 | 41185.4 | 1803.7 | 66.086 | 36348.2 | 1555.50 | 47401.2 | 0.0 |
| KES317 | 42324.0 | 1977.1 | 66.172 | 40875.0 | 1581.19 | 48065.4 | 102.2 |
| KES318 | 51103.4 | 1016.5 | 31.286 | 38353.3 | 2168.90 | 53765.7 | 0.0 |
| KES319 | 42173.2 | 1063.4 | 33.797 | 35350.1 | 2339.91 | 52111.7 | 0.0 |
| KES320 | 33383.5 | 2574.8 | 65.856 | 39754.1 | 2317.28 | 58527.6 | 0.0 |
| KES321 | 38859.5 | 1866.6 | 50.981 | 33610.7 | 1811.58 | 50786.2 | 0.0 |
| KES322 | 38088.3 | 1225.3 | 36.491 | 31953.6 | 2291.76 | 52791.3 | 0.0 |
| KES323 | 47066.0 | 1507.7 | 43.476 | 33431.0 | 1802.49 | 47159.3 | 0.0 |
| KES324 | 42738.9 | 1673.6 | 43.342 | 32509.8 | 1829.50 | 47530.9 | 130.0 |
| KES325 | 40223.3 | 1645.7 | 43.818 | 33839.9 | 1845.62 | 48330.4 | 216.7 |
| KES326 | 44405.9 | 1652.4 | 43.452 | 37000.3 | 1848.12 | 48327.8 | 138.5 |
| KES327 | 46287.2 | 1605.6 | 44.501 | 31941.2 | 1840.40 | 47919.1 | 0.0 |
| KES328 | 44673.9 | 1644.6 | 45.499 | 35285.7 | 1889.70 | 49108.9 | 0.0 |
| KES329 | 47937.2 | 1707.8 | 45.729 | 39372.1 | 1899.51 | 49696.8 | 148.5 |
| KES330 | 42799.9 | 1603.8 | 43.808 | 32238.0 | 1847.61 | 48165.2 | 0.0 |
| KES331 | 47659.0 | 1656.8 | 45.118 | 34808.0 | 1883.18 | 49003.1 | 0.0 |
| KES335 | 45937.3 | 1638.1 | 44.316 | 34367.7 | 1882.11 | 48843.9 | 141.4 |
| KES336 | 43097.0 | 1581.1 | 44.514 | 42995.6 | 1864.85 | 48459.5 | 122.9 |
| KES337 | 44396.0 | 1708.9 | 43.692 | 35565.4 | 1827.80 | 47692.7 | 126.4 |
| KES338 | 49151.1 | 1638.6 | 43.966 | 37887.5 | 1854.59 | 48432.6 | 0.0 |
| KES339 | 45040.7 | 1667.7 | 43.667 | 40773.7 | 1835.26 | 47914.4 | 0.0 |
| KES340 | 52449.0 | 1614.0 | 44.013 | 41159.2 | 1874.18 | 49055.1 | 101.2 |
| KES341 | 51508.6 | 1088.9 | 32.960 | 34851.7 | 2328.57 | 52197.7 | 0.0 |
| KES342 | 42808.3 | 1291.6 | 42.927 | 40513.3 | 1976.01 | 48890.2 | 138.3 |
| KES343 | 42640.9 | 1042.1 | 35.004 | 35095.8 | 2338.18 | 52400.9 | 0.0 |
| KES344 | 45428.5 | 1194.1 | 34.151 | 38220.7 | 2345.31 | 52415.4 | 0.0 |
| KES345 | 58139.8 | 714.6 | 20.960 | 36089.6 | 2529.59 | 50857.2 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES346 | 61399.7 | 792.4 | 21.157 | 39780.2 | 2555.86 | 51636.7 | 0.0 |
| KES347 | 47033.9 | 1568.4 | 49.139 | 37366.8 | 1288.46 | 44467.5 | 0.0 |
| KES348 | 48238.8 | 1500.6 | 47.679 | 39083.1 | 1295.15 | 44406.7 | 0.0 |
| KES349 | 47436.5 | 1665.5 | 50.154 | 35913.7 | 1305.89 | 44794.5 | 0.0 |
| KES350 | 45443.1 | 2507.2 | 70.991 | 32408.2 | 1813.20 | 55962.0 | 0.0 |
| KES351 | 41354.3 | 2515.2 | 73.882 | 31588.5 | 1835.97 | 56775.8 | 0.0 |
| KES352 | 44793.1 | 2453.5 | 73.683 | 36146.4 | 1828.84 | 56738.2 | 0.0 |
| KES353 | 40370.5 | 2461.1 | 71.208 | 34051.2 | 1793.78 | 55688.6 | 0.0 |
| KES354 | 34947.0 | 2577.2 | 72.460 | 34086.9 | 1823.64 | 56306.2 | 0.0 |
| KES355 | 42222.4 | 2496.0 | 72.956 | 33706.0 | 1786.72 | 55423.3 | 0.0 |
| KES356 | 37358.3 | 2479.8 | 71.834 | 32605.7 | 1829.87 | 56494.5 | 0.0 |
| KES357 | 44796.2 | 2580.8 | 72.256 | 35911.6 | 1834.77 | 56756.2 | 0.0 |
| KES358 | 38519.5 | 2430.5 | 74.559 | 33089.0 | 1823.86 | 56303.4 | 0.0 |
| KES359 | 45244.3 | 2591.7 | 73.577 | 36033.2 | 1853.05 | 56785.0 | 0.0 |
| KES360 | 38295.8 | 2554.8 | 72.344 | 33903.1 | 1824.37 | 56420.8 | 0.0 |
| KES361 | 40524.9 | 2364.5 | 71.329 | 40711.9 | 1801.95 | 55662.2 | 0.0 |
| KES362 | 41818.0 | 2492.8 | 70.690 | 33141.0 | 1802.46 | 56129.0 | 0.0 |
| KES363 | 42260.0 | 2577.6 | 71.869 | 29277.6 | 1822.41 | 57166.3 | 0.0 |
| KES364 | 47243.2 | 2580.0 | 67.447 | 34647.7 | 1658.34 | 49670.7 | 0.0 |
| KES365 | 43220.6 | 2474.8 | 67.100 | 32623.7 | 1653.08 | 49586.5 | 0.0 |
| KES366 | 42511.6 | 2476.3 | 65.257 | 35559.2 | 1640.93 | 49029.8 | 0.0 |
| KES367 | 45810.6 | 2385.7 | 67.223 | 32478.7 | 1630.54 | 48796.5 | 0.0 |
| KES368 | 39592.5 | 2538.2 | 67.943 | 32712.2 | 1643.52 | 49834.4 | 0.0 |
| KES369 | 40539.2 | 2485.0 | 67.372 | 34745.4 | 1634.28 | 48792.4 | 0.0 |
| KES370 | 41252.0 | 1232.1 | 29.225 | 37467.1 | 2216.44 | 47124.5 | 0.0 |
| KES371 | 43160.2 | 1182.1 | 28.456 | 32247.1 | 2291.76 | 45987.0 | 0.0 |
| KES372 | 46197.3 | 1242.0 | 27.738 | 33350.8 | 2171.02 | 45837.0 | 0.0 |
| KES373 | 41943.3 | 1285.9 | 27.340 | 38533.7 | 2201.59 | 46561.1 | 0.0 |
| KES374 | 58238.8 | 1530.3 | 28.203 | 41290.9 | 449.09 | 38674.1 | 0.0 |
| KES375 | 57545.4 | 1473.9 | 26.036 | 32793.6 | 431.10 | 37184.1 | 75.9 |
| KES376 | 58156.7 | 1841.0 | 33.979 | 38846.0 | 468.89 | 40385.3 | 62.5 |
| KES377 | 58870.1 | 1889.7 | 33.137 | 34903.8 | 461.10 | 39818.7 | 77.5 |
| KES378 | 55789.7 | 1957.7 | 33.655 | 33733.0 | 461.20 | 39972.1 | 57.6 |
| KES379 | 56930.1 | 1643.4 | 34.571 | 36089.3 | 463.48 | 39842.5 | 0.0 |
| KES380 | 64425.7 | 1631.8 | 34.438 | 37985.4 | 473.11 | 40552.3 | 0.0 |
| KES381 | 58081.9 | 1680.2 | 34.795 | 36350.5 | 465.37 | 39750.2 | 64.2 |
| KES382 | 59609.1 | 1669.6 | 34.131 | 34357.4 | 471.11 | 40500.2 | 85.4 |
| KES383 | 58766.7 | 1613.9 | 33.863 | 35836.5 | 472.10 | 40277.1 | 0.0 |
| KES384 | 53354.6 | 1692.0 | 33.911 | 37893.0 | 467.80 | 40042.6 | 0.0 |
| KES385 | 62633.8 | 1630.1 | 33.370 | 35213.2 | 464.64 | 39895.6 | 66.0 |
| KES386 | 57755.4 | 1730.9 | 33.652 | 35868.4 | 468.93 | 40315.2 | 0.0 |
| KES387 | 61915.8 | 1610.2 | 34.219 | 36681.2 | 466.30 | 39845.9 | 0.0 |
| KES388 | 62539.2 | 1721.5 | 35.186 | 34931.5 | 476.22 | 40825.1 | 69.4 |
| KES389 | 65379.8 | 1643.0 | 34.967 | 38100.9 | 478.87 | 40793.4 | 0.0 |
| KES390 | 58879.6 | 1750.3 | 35.080 | 34776.3 | 473.95 | 40516.7 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | Cl | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES391 | 65423.5 | 1627.7 | 35.452 | 35050.3 | 486.69 | 41788.8 | 0.0 |
| KES392 | 57239.8 | 1665.9 | 36.837 | 36269.9 | 485.27 | 40502.0 | 0.0 |
| KES393 | 58379.7 | 1646.3 | 32.910 | 36604.7 | 462.98 | 40061.3 | 106.8 |
| KES394 | 62217.7 | 1534.9 | 37.326 | 37005.5 | 452.77 | 39856.1 | 102.1 |
| KES403 | 51067.7 | 1426.7 | 27.7967 | 35834.7 | 2165.31 | 45918.1 | 0.0 |
| KES404 | 54166.8 | 1479.9 | 28.9645 | 36985.6 | 2320.72 | 46387.5 | 0.0 |
| KES405 | 43041.0 | 1312.7 | 28.6762 | 38443.9 | 2329.30 | 46301.6 | 0.0 |
| KES406 | 37256.0 | 1358.1 | 28.3368 | 36712.0 | 2232.73 | 46987.1 | 0.0 |
| KES407 | 49927.1 | 1423.6 | 27.4910 | 39916.5 | 2166.31 | 45818.7 | 0.0 |
| KES408 | 68703.6 | 2167.8 | 38.3726 | 26739.8 | 4041.44 | 92923.5 | 0.0 |
| KES409 | 81724.2 | 2340.1 | 41.6976 | 40783.6 | 4162.81 | 95454.9 | 0.0 |
| KES410 | 70461.7 | 2290.6 | 40.3164 | 36698.3 | 4107.31 | 94287.0 | 0.0 |
| KES411 | 57329.7 | 1931.3 | 27.4926 | 41275.3 | 3043.57 | 66393.6 | 0.0 |
| KES412 | 53162.3 | 3850.7 | 38.5651 | 35999.9 | 1451.18 | 50290.8 | 145.1 |
| KES413 | 51800.1 | 3837.8 | 36.0300 | 34506.2 | 1470.70 | 51434.8 | 110.4 |
| KES414 | 64779.1 | 1531.6 | 20.7193 | 37895.2 | 2951.81 | 61979.0 | 0.0 |
| KES415 | 75616.3 | 1787.5 | 23.7351 | 36279.7 | 3025.65 | 63582.9 | 0.0 |
| KES416 | 65858.3 | 1807.0 | 21.3926 | 38098.0 | 2965.36 | 61546.2 | 0.0 |
| KES417 | 70476.0 | 1687.6 | 22.8927 | 34029.4 | 3005.91 | 63210.3 | 0.0 |
| KES418 | 67399.8 | 918.4 | 22.5488 | 34634.6 | 2420.13 | 58865.1 | 0.0 |
| KES419 | 86322.1 | 1284.8 | 16.8574 | 42822.4 | 2610.13 | 70207.5 | 0.0 |
| KES420 | 81544.0 | 1192.1 | 17.3727 | 41879.6 | 2630.83 | 71557.2 | 0.0 |
| KES421 | 88992.3 | 1234.0 | 19.1463 | 41266.0 | 2640.24 | 71854.4 | 0.0 |
| KES422 | 80771.0 | 1479.2 | 18.5464 | 41726.8 | 2639.24 | 55806.6 | 0.0 |
| KES423 | 77680.0 | 1187.1 | 16.6364 | 33947.6 | 2594.03 | 70716.8 | 0.0 |
| KES424 | 66253.1 | 1752.5 | 25.5142 | 40038.1 | 2421.64 | 69536.0 | 0.0 |
| KES425 | 72073.0 | 1397.8 | 25.6167 | 33452.1 | 2371.57 | 65039.7 | 0.0 |
| KES426 | 79508.4 | 1516.0 | 24.4505 | 33004.5 | 2295.34 | 65168.8 | 129.7 |
| KES427 | 84597.4 | 1420.8 | 26.4695 | 36199.0 | 2355.88 | 66351.7 | 0.0 |
| KES428 | 63032.2 | 1405.4 | 23.9832 | 33753.3 | 2304.78 | 64781.2 | 0.0 |
| KES429 | 82820.7 | 1521.8 | 17.4545 | 40656.9 | 2523.88 | 68850.6 | 0.0 |
| KES430 | 66489.6 | 746.7 | 12.4515 | 48712.3 | 1560.33 | 27183.0 | 501.1 |
| KES431 | 58377.1 | 977.5 | 16.4394 | 36827.9 | 900.55 | 37255.1 | 60.3 |
| KES432 | 63466.4 | 1154.4 | 16.8257 | 38665.5 | 884.68 | 37432.1 | 95.4 |
| KES433 | 50167.3 | 1305.0 | 19.3680 | 39215.3 | 1055.26 | 41155.0 | 0.0 |
| KES434 | 54416.2 | 1113.2 | 15.7694 | 40042.9 | 909.01 | 38309.7 | 97.0 |
| KES435 | 58856.8 | 1318.0 | 14.9229 | 40307.8 | 893.77 | 37043.5 | 114.4 |
| KES436 | 69485.1 | 0.0 | 20.6529 | 41259.6 | 2634.56 | 44539.8 | 0.0 |
| KES437 | 87031.7 | 1146.0 | 14.0616 | 43723.2 | 2476.65 | 64521.8 | 0.0 |
| KES438 | 87970.4 | 1268.8 | 16.9153 | 34506.2 | 2571.62 | 68635.5 | 0.0 |
| KES439 | 70135.4 | 1822.2 | 29.7717 | 39316.8 | 2629.01 | 65004.9 | 109.3 |
| KES440 | 53990.4 | 1878.8 | 31.0821 | 36378.9 | 2596.19 | 64136.4 | 0.0 |
| KES441 | 57580.9 | 1835.6 | 29.1392 | 34352.5 | 2708.35 | 64975.7 | 115.5 |
| KES442 | 72050.7 | 664.7 | 25.0103 | 41448.1 | 2769.28 | 59467.8 | 200.5 |
| KES443 | 69524.9 | 556.8 | 23.7235 | 43892.5 | 2804.46 | 60015.1 | 147.9 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES444 | 39249.9 | 2266.4 | 47.2820 | 32655.5 | 3004.76 | 53848.2 | 173.8 |
| KES445 | 38341.9 | 2126.3 | 47.2253 | 29541.7 | 2947.23 | 52640.8 | 140.2 |
| KES446 | 34682.3 | 3781.2 | 74.7878 | 35526.6 | 3098.70 | 60859.6 | 194.6 |
| KES447 | 48339.0 | 1347.3 | 26.9782 | 34419.1 | 3086.75 | 58806.1 | 0.0 |
| KES448 | 49354.1 | 1756.6 | 26.5260 | 28935.0 | 3080.58 | 59461.2 | 0.0 |
| KES449 | 53181.5 | 1761.8 | 23.1987 | 34395.7 | 2492.54 | 53078.7 | 0.0 |
| KES450 | 54748.2 | 1869.5 | 23.0573 | 37257.1 | 2471.53 | 54658.4 | 0.0 |
| KES451 | 48094.5 | 1662.6 | 21.6742 | 35454.6 | 2403.57 | 53210.7 | 98.1 |
| KES452 | 46731.4 | 1720.2 | 22.8476 | 35055.1 | 2423.24 | 53126.0 | 0.0 |
| KES453 | 48484.2 | 1676.5 | 22.6889 | 35713.4 | 2555.84 | 54476.4 | 0.0 |
| KES454 | 50764.3 | 1810.7 | 22.3791 | 33846.7 | 2402.34 | 53231.5 | 0.0 |
| KES455 | 48927.9 | 1715.2 | 21.6011 | 36375.2 | 2406.12 | 52714.0 | 0.0 |
| KES456 | 52271.6 | 1790.2 | 22.5125 | 37937.9 | 2416.39 | 53168.8 | 0.0 |
| KES457 | 47125.3 | 1504.1 | 23.4363 | 33972.3 | 2686.73 | 53978.0 | 231.7 |
| KES458 | 49289.3 | 1607.9 | 21.8285 | 35949.6 | 2664.03 | 54036.8 | 191.6 |
| KES459 | 89080.2 | 1039.7 | 16.9987 | 39028.8 | 2224.62 | 63914.3 | 0.0 |
| KES460 | 97550.7 | 922.6 | 14.4503 | 47275.5 | 2199.23 | 62454.9 | 0.0 |
| KES461 | 83527.9 | 1329.1 | 18.3784 | 43146.7 | 2624.91 | 70581.7 | 0.0 |
| KES462 | 86143.6 | 950.9 | 14.6733 | 39501.1 | 2165.62 | 62556.6 | 0.0 |
| KES463 | 81789.6 | 1286.1 | 17.6753 | 38368.6 | 2355.50 | 61395.5 | 0.0 |
| KES464 | 85697.6 | 1264.3 | 17.6501 | 48014.0 | 2374.79 | 61126.9 | 0.0 |
| KES465 | 83587.1 | 1197.8 | 16.7552 | 41495.9 | 2360.75 | 61237.9 | 0.0 |
| KES466 | 86404.8 | 1296.9 | 17.1919 | 46699.5 | 2380.14 | 62088.9 | 0.0 |
| KES467 | 41657.8 | 3222.4 | 69.9159 | 36257.4 | 1754.39 | 51608.1 | 219.5 |
| KES468 | 42100.6 | 3235.7 | 67.8909 | 33960.6 | 1715.60 | 50861.6 | 0.0 |
| KES469 | 35429.6 | 3330.9 | 68.3281 | 35673.7 | 1838.10 | 51235.3 | 0.0 |
| KES470 | 42491.2 | 3379.2 | 64.8706 | 35730.9 | 1696.76 | 50662.1 | 0.0 |
| KES471 | 46545.8 | 2161.9 | 41.2161 | 35968.7 | 2074.34 | 51345.4 | 0.0 |
| KES472 | 42184.7 | 2120.3 | 40.5722 | 34133.3 | 2048.93 | 50792.7 | 0.0 |
| KES473 | 42617.3 | 2190.9 | 41.5700 | 35163.5 | 2076.02 | 51216.2 | 0.0 |
| KES474 | 42805.1 | 1995.2 | 42.2342 | 30360.3 | 2057.91 | 51380.7 | 0.0 |
| KES475 | 46683.5 | 3268.6 | 69.0423 | 36513.9 | 1753.09 | 51929.2 | 296.4 |
| KES476 | 44249.3 | 3029.2 | 67.7062 | 39312.0 | 1714.81 | 50821.5 | 0.0 |
| KES477 | 45767.9 | 2160.4 | 41.3468 | 36593.9 | 2089.45 | 51679.6 | 0.0 |
| KES478 | 48551.8 | 2133.2 | 42.0119 | 33148.9 | 2074.10 | 51245.3 | 0.0 |
| KES479 | 46540.6 | 2717.8 | 55.6296 | 31784.1 | 1656.14 | 48500.5 | 0.0 |
| KES480 | 40597.3 | 1898.0 | 42.5506 | 35548.2 | 2160.82 | 54003.8 | 0.0 |
| KES481 | 44953.2 | 1889.0 | 39.3153 | 32907.4 | 1992.86 | 49917.3 | 0.0 |
| KES482 | 39350.0 | 1872.0 | 40.5553 | 28307.9 | 2006.47 | 49960.3 | 0.0 |
| KES483 | 42981.3 | 3018.7 | 64.2946 | 33941.7 | 1687.85 | 50902.0 | 0.0 |
| KES484 | 38684.6 | 3058.9 | 62.6561 | 35852.5 | 1654.21 | 49460.7 | 0.0 |
| KES485 | 42784.2 | 2656.1 | 53.9359 | 34150.0 | 1695.31 | 47876.4 | 0.0 |
| KES486 | 42205.5 | 2669.7 | 54.5768 | 39411.1 | 1641.00 | 48536.5 | 0.0 |
| KES487 | 50328.0 | 2735.3 | 51.8981 | 39190.4 | 1528.53 | 47136.5 | 0.0 |
| KES488 | 42851.3 | 2412.0 | 50.0614 | 40196.4 | 1588.51 | 45563.2 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Al | CI | Dy | K | Mn | Na | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES489 | 43558.3 | 2607.5 | 50.6044 | 33989.4 | 1692.20 | 46309.9 | 0.0 |
| KES490 | 36981.7 | 2443.6 | 52.5111 | 37191.9 | 1745.97 | 47951.0 | 0.0 |
| KES491 | 44043.4 | 2491.7 | 51.4370 | 40397.3 | 1722.91 | 47152.8 | 0.0 |
| KES492 | 45811.6 | 2748.0 | 51.9067 | 37976.9 | 1717.94 | 47170.9 | 0.0 |
| KES493 | 42304.0 | 2760.6 | 50.9092 | 31200.6 | 1703.92 | 46742.5 | 0.0 |
| KES494 | 42680.0 | 2674.8 | 51.2331 | 37409.0 | 1727.96 | 47214.7 | 0.0 |
| KES495 | 45541.5 | 2542.2 | 48.0831 | 36034.2 | 1687.82 | 46348.1 | 65.1 |
| KES496 | 54856.1 | 2125.8 | 31.4480 | 37467.8 | 1392.61 | 46028.7 | 67.5 |
| KES497 | 51564.8 | 2115.6 | 29.5586 | 38732.7 | 1409.05 | 45986.8 | 0.0 |
| KES498 | 55697.9 | 2185.5 | 29.2988 | 33482.2 | 1412.38 | 45348.7 | 0.0 |
| KES499 | 48622.2 | 2054.7 | 31.9971 | 38016.5 | 1421.29 | 45701.4 | 0.0 |
| KES500 | 53484.4 | 2078.2 | 29.2661 | 38286.1 | 1438.82 | 45837.6 | 0.0 |
| KES501 | 47841.2 | 2307.3 | 30.5031 | 34118.6 | 1518.38 | 44956.4 | 0.0 |
| KES502 | 48695.7 | 2203.2 | 30.3878 | 34167.9 | 1401.45 | 45073.4 | 0.0 |
| KES503 | 94175.4 | 1657.6 | 12.2381 | 44724.3 | 1979.48 | 63101.3 | 0.0 |
| KES504 | 88877.2 | 1632.5 | 11.4647 | 44970.6 | 1831.09 | 63759.9 | 0.0 |
| KES505 | 91274.4 | 1779.3 | 13.5928 | 39597.1 | 1796.51 | 64461.8 | 0.0 |
| KES506 | 91246.8 | 1721.3 | 12.3369 | 49333.6 | 1769.21 | 63140.9 | 0.0 |
| KES507 | 89683.8 | 1665.6 | 12.8491 | 42278.5 | 1796.26 | 64859.7 | 0.0 |
| KES508 | 91636.5 | 1758.5 | 12.3822 | 41410.1 | 1778.99 | 63978.8 | 0.0 |
| KES509 | 78515.6 | 815.3 | 18.7407 | 41091.6 | 2434.34 | 49741.2 | 129.5 |
| KES510 | 66452.8 | 912.2 | 19.2610 | 43641.6 | 2522.34 | 50625.4 | 0.0 |
| KES511 | 68420.0 | 961.9 | 19.4786 | 39490.0 | 2433.89 | 52138.3 | 102.5 |
| KES512 | 73822.2 | 975.3 | 18.8814 | 43865.2 | 2370.09 | 51151.5 | 70.7 |
| KES513 | 70884.7 | 618.3 | 16.3506 | 58683.8 | 2230.05 | 31941.8 | 115.7 |
| KES514 | 88069.9 | 1235.9 | 14.1379 | 39938.1 | 2218.09 | 58918.2 | 269.0 |
| KES515 | 81709.7 | 1148.0 | 14.2100 | 40256.0 | 2153.64 | 58148.2 | 294.6 |
| KES516 | 82648.8 | 1112.0 | 15.1491 | 35189.4 | 1138.39 | 44064.6 | 678.9 |
| KES517 | 79805.1 | 1033.3 | 15.3760 | 30474.9 | 1154.40 | 42898.2 | 644.9 |
| KES518 | 83203.1 | 831.3 | 15.7807 | 33668.3 | 1195.68 | 45685.2 | 654.1 |
| KES519 | 77858.9 | 1176.2 | 14.4376 | 34958.8 | 1125.67 | 43402.7 | 616.6 |
| KES520 | 80611.8 | 1062.0 | 14.8772 | 31321.5 | 1192.57 | 43950.7 | 655.2 |
| KES521 | 80931.4 | 1097.7 | 14.6717 | 33145.7 | 1122.96 | 43244.9 | 565.0 |
| KES522 | 81629.8 | 1141.3 | 14.8446 | 33020.2 | 1149.48 | 44540.9 | 651.3 |
| KES523 | 87588.9 | 1152.7 | 15.2361 | 34004.7 | 1145.46 | 44320.4 | 631.6 |
| KES524 | 77267.3 | 1113.8 | 15.1835 | 28484.8 | 1174.45 | 43729.9 | 676.1 |
| KES525 | 80517.5 | 985.3 | 15.4979 | 35062.3 | 1185.66 | 43859.8 | 615.9 |
| KES526 | 76109.0 | 1078.5 | 15.0509 | 34994.8 | 1109.18 | 44079.8 | 593.6 |
| KES527 | 80524.8 | 604.3 | 9.3380 | 41570.2 | 2079.82 | 52413.8 | 0.0 |
| KES528 | 87962.1 | 569.6 | 10.3347 | 47016.2 | 2000.50 | 54143.4 | 0.0 |
| KES529 | 78014.3 | 601.2 | 10.9268 | 46962.1 | 1966.58 | 53033.0 | 0.0 |
| KES530 | 77906.3 | 389.3 | 9.7746 | 47249.0 | 1829.55 | 43483.9 | 62.0 |
| KES531 | 80669.0 | 344.7 | 10.1121 | 44977.2 | 1874.49 | 43396.6 | 87.7 |
| KES532 | 82312.6 | 509.8 | 10.4363 | 43054.1 | 1855.50 | 46047.8 | 57.3 |
| KES533 | 82539.1 | 405.0 | 10.8257 | 44240.3 | 1908.92 | 45236.1 | 0.0 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | $\mathbf{A l}$ | $\mathbf{C l}$ | $\mathbf{D y}$ | $\mathbf{K}$ | $\mathbf{M n}$ | $\mathbf{N a}$ | Ba |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES534 | 62418.0 | 2020.4 | 35.3839 | 37510.7 | 441.59 | 39239.7 | 72.9 |
| KES535 | 56363.8 | 2030.9 | 36.2948 | 35445.2 | 441.87 | 39293.4 | 55.5 |
| KES536 | 47114.2 | 1267.8 | 28.5641 | 35512.3 | 2186.03 | 46105.2 | 68.2 |
| KES537 | 47868.4 | 1352.8 | 28.5459 | 36624.3 | 2190.22 | 46290.1 | 85.5 |
| KES538 | 44869.6 | 1446.7 | 28.2985 | 37070.1 | 2219.15 | 46703.3 | 67.2 |
| KES539 | 48381.3 | 1353.8 | 27.5955 | 35666.5 | 2195.49 | 46585.8 | 134.1 |
| KES540 | 74308.6 | 1856.1 | 27.7029 | 37821.6 | 3034.93 | 65896.3 | 42.6 |
| KES541 | 78424.9 | 1235.9 | 24.4409 | 34827.8 | 2233.02 | 62618.4 | 49.6 |
| KES542 | 69071.3 | 1381.9 | 25.0526 | 34794.3 | 2359.17 | 67333.1 | 0.0 |

Appendix 6 - Concentration data from NAA for obsidian from Kenya. All concentrations are in parts per million (ppm)

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES001 | 418.518 | 0.189 | 4.196 | 1.919 | 47950.2 | 39.154 | 215.510 |
| KES002 | 420.077 | 0.185 | 4.280 | 1.972 | 48491.3 | 39.158 | 216.151 |
| KES003 | 143.183 | 0.355 | 4.136 | 0.155 | 13291.3 | 16.442 | 71.697 |
| KES004 | 150.857 | 0.523 | 3.930 | 0.179 | 13321.8 | 16.846 | 76.394 |
| KES005 | 845.528 | 0.177 | 6.549 | 4.609 | 53752.6 | 77.659 | 438.370 |
| KES006 | 918.127 | 0.000 | 7.562 | 4.740 | 56682.8 | 80.665 | 469.564 |
| KES007 | 894.034 | 0.141 | 6.919 | 4.873 | 53553.7 | 82.869 | 466.748 |
| KES008 | 518.614 | 0.241 | 1.893 | 4.321 | 67104.0 | 41.300 | 259.103 |
| KES009 | 506.603 | 0.240 | 1.888 | 4.276 | 65503.9 | 40.504 | 256.228 |
| KES010 | 971.649 | 0.157 | 7.523 | 5.287 | 57024.7 | 89.741 | 504.070 |
| KES011 | 953.508 | 0.175 | 7.354 | 5.139 | 55917.4 | 87.599 | 495.179 |
| KES012 | 745.535 | 0.164 | 5.729 | 4.213 | 55854.7 | 66.995 | 389.466 |
| KES013 | 510.265 | 0.388 | 2.145 | 3.289 | 58312.4 | 43.082 | 258.984 |
| KES014 | 495.112 | 0.155 | 2.810 | 5.206 | 60359.6 | 40.000 | 249.782 |
| KES015 | 508.037 | 0.192 | 2.955 | 5.370 | 61779.0 | 42.249 | 259.304 |
| KES016 | 292.733 | 1.513 | 3.284 | 1.808 | 24391.2 | 28.331 | 158.774 |
| KES017 | 293.620 | 1.581 | 3.279 | 1.834 | 24763.6 | 28.445 | 159.212 |
| KES018 | 294.241 | 1.505 | 3.268 | 1.858 | 24398.1 | 28.344 | 158.332 |
| KES019 | 289.108 | 1.482 | 3.148 | 1.808 | 23895.4 | 28.017 | 157.362 |
| KES020 | 294.901 | 2.600 | 1.964 | 3.209 | 63784.2 | 22.475 | 169.103 |
| KES021 | 271.603 | 0.390 | 6.602 | 0.646 | 27305.1 | 41.890 | 130.605 |
| KES022 | 267.633 | 0.661 | 6.261 | 0.648 | 27347.1 | 39.690 | 129.204 |
| KES023 | 271.968 | 0.202 | 6.638 | 0.580 | 26516.2 | 42.077 | 129.313 |
| KES024 | 271.751 | 0.213 | 6.569 | 0.583 | 26251.1 | 41.478 | 131.380 |
| KES025 | 273.858 | 0.147 | 6.888 | 0.602 | 26794.9 | 43.682 | 129.393 |
| KES026 | 267.982 | 0.134 | 6.779 | 0.587 | 26241.2 | 42.845 | 128.044 |
| KES027 | 273.850 | 0.272 | 6.923 | 0.622 | 27241.3 | 43.522 | 130.810 |
| KES028 | 273.647 | 0.226 | 6.888 | 0.604 | 26960.5 | 43.453 | 129.770 |
| KES029 | 270.025 | 0.138 | 6.899 | 0.582 | 26356.4 | 43.081 | 128.435 |
| KES030 | 257.446 | 0.136 | 6.586 | 0.562 | 25001.8 | 41.128 | 122.867 |
| KES031 | 258.581 | 0.118 | 6.615 | 0.580 | 25240.7 | 41.184 | 122.792 |
| KES032 | 259.019 | 0.145 | 6.679 | 0.560 | 25149.4 | 41.185 | 122.897 |
| KES033 | 489.016 | 0.361 | 2.071 | 3.148 | 55792.8 | 40.898 | 246.935 |
| KES034 | 494.793 | 0.343 | 2.156 | 3.166 | 55904.8 | 41.384 | 249.981 |
| KES035 | 495.602 | 0.364 | 2.100 | 3.204 | 56341.9 | 41.583 | 250.978 |
| KES036 | 490.599 | 0.364 | 2.063 | 3.163 | 55926.0 | 41.123 | 247.989 |
| KES037 | 495.761 | 0.362 | 2.088 | 3.194 | 56256.8 | 41.499 | 249.799 |
| KES038 | 495.231 | 0.368 | 1.998 | 3.201 | 56123.2 | 41.335 | 249.217 |
| KES039 | 486.972 | 0.372 | 2.043 | 3.096 | 55352.7 | 40.796 | 247.591 |
| KES040 | 486.445 | 0.401 | 2.046 | 3.133 | 55648.7 | 40.886 | 248.055 |
| KES042 | 481.757 | 0.301 | 2.952 | 3.016 | 62421.3 | 39.814 | 253.501 |
| KES043 | 493.056 | 0.213 | 1.822 | 4.249 | 63335.3 | 38.946 | 247.731 |
| KES044 | 488.254 | 0.193 | 1.865 | 4.268 | 62861.5 | 38.868 | 246.691 |
| KES045 | 499.234 | 0.204 | 1.877 | 3.989 | 62490.2 | 40.337 | 252.071 |
| KES046 | 501.253 | 0.176 | 1.826 | 4.066 | 62614.5 | 40.425 | 253.892 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES047 | 242.315 | 0.126 | 6.222 | 0.542 | 23797.8 | 38.876 | 116.145 |
| KES048 | 258.102 | 0.116 | 6.574 | 0.554 | 24954.4 | 41.074 | 122.798 |
| KES049 | 257.266 | 0.121 | 6.583 | 0.560 | 24972.5 | 40.904 | 122.955 |
| KES050 | 259.607 | 0.102 | 6.361 | 0.572 | 25069.4 | 44.698 | 122.266 |
| KES051 | 261.277 | 0.116 | 6.456 | 0.566 | 25141.2 | 44.803 | 122.060 |
| KES052 | 330.139 | 0.130 | 8.042 | 0.717 | 28660.4 | 54.034 | 156.435 |
| KES053 | 329.022 | 0.138 | 8.193 | 0.729 | 28597.5 | 53.931 | 155.410 |
| KES054 | 449.587 | 0.210 | 2.088 | 3.435 | 67254.4 | 38.550 | 228.423 |
| KES055 | 452.442 | 0.223 | 2.162 | 3.391 | 67784.5 | 38.895 | 229.531 |
| KES056 | 453.462 | 0.216 | 2.124 | 3.412 | 67672.2 | 38.980 | 231.058 |
| KES057 | 418.967 | 0.182 | 1.621 | 2.836 | 64503.2 | 25.906 | 223.556 |
| KES058 | 777.255 | 0.142 | 6.086 | 4.151 | 49758.4 | 70.490 | 406.222 |
| KES059 | 873.751 | 0.041 | 7.313 | 4.483 | 54166.2 | 77.012 | 452.039 |
| KES060 | 595.128 | 0.161 | 4.927 | 3.185 | 45596.4 | 55.910 | 313.676 |
| KES061 | 598.426 | 0.157 | 4.958 | 3.220 | 45843.3 | 56.129 | 313.574 |
| KES062 | 716.337 | 0.216 | 5.487 | 3.992 | 53384.4 | 63.939 | 375.166 |
| KES063 | 707.813 | 0.447 | 5.378 | 3.968 | 52255.3 | 63.155 | 374.976 |
| KES064 | 707.145 | 0.159 | 5.432 | 3.963 | 52487.2 | 64.730 | 382.311 |
| KES065 | 760.070 | 0.195 | 5.838 | 4.134 | 51678.0 | 68.672 | 401.338 |
| KES066 | 282.568 | 1.055 | 1.617 | 3.055 | 61100.5 | 23.432 | 155.261 |
| KES067 | 706.204 | 0.193 | 5.502 | 3.782 | 49228.1 | 64.513 | 375.405 |
| KES068 | 712.149 | 0.171 | 5.494 | 3.799 | 49467.5 | 65.036 | 376.424 |
| KES069 | 707.662 | 0.208 | 5.443 | 3.800 | 49559.3 | 64.805 | 376.439 |
| KES070 | 878.221 | 0.421 | 7.289 | 4.503 | 54702.2 | 77.698 | 455.423 |
| KES071 | 875.489 | 0.111 | 7.322 | 4.517 | 54392.2 | 77.668 | 460.800 |
| KES072 | 872.520 | 0.121 | 7.203 | 4.473 | 54244.8 | 77.173 | 456.270 |
| KES073 | 880.096 | 0.085 | 7.235 | 4.504 | 54399.7 | 79.001 | 456.228 |
| KES074 | 879.421 | 0.110 | 7.126 | 4.498 | 54402.2 | 77.687 | 455.670 |
| KES075 | 238.790 | 0.101 | 6.390 | 0.530 | 24391.2 | 51.963 | 106.342 |
| KES076 | 147.181 | 0.401 | 4.111 | 0.154 | 14580.5 | 16.410 | 75.509 |
| KES077 | 137.458 | 0.336 | 4.177 | 0.163 | 13217.2 | 16.652 | 69.220 |
| KES078 | 145.733 | 0.634 | 4.027 | 0.175 | 17465.3 | 16.839 | 74.581 |
| KES079 | 804.424 | 0.161 | 6.218 | 4.374 | 51005.1 | 74.417 | 433.984 |
| KES080 | 216.714 | 0.312 | 0.723 | 2.260 | 68864.4 | 16.920 | 113.481 |
| KES081 | 496.269 | 0.376 | 2.016 | 3.175 | 56162.5 | 41.580 | 257.378 |
| KES082 | 501.019 | 0.416 | 2.064 | 3.203 | 56617.6 | 41.876 | 259.641 |
| KES083 | 498.441 | 0.400 | 2.053 | 3.220 | 56815.4 | 42.065 | 259.271 |
| KES084 | 503.767 | 0.402 | 2.118 | 3.196 | 56660.6 | 42.058 | 258.074 |
| KES085 | 498.330 | 0.399 | 2.105 | 3.199 | 56361.0 | 41.753 | 258.166 |
| KES086 | 495.210 | 0.411 | 2.190 | 3.177 | 56593.9 | 41.614 | 258.978 |
| KES087 | 894.679 | 0.198 | 6.857 | 4.812 | 52732.0 | 82.600 | 468.085 |
| KES088 | 868.253 | 0.161 | 6.704 | 4.771 | 51697.7 | 82.232 | 468.167 |
| KES089 | 326.620 | 0.152 | 1.184 | 3.588 | 69678.1 | 27.066 | 163.052 |
| KES090 | 334.117 | 0.136 | 1.250 | 3.645 | 70546.7 | 27.569 | 165.574 |
| KES091 | 744.823 | 0.243 | 5.696 | 4.134 | 54978.1 | 66.134 | 359.260 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES092 | 744.866 | 0.215 | 5.563 | 4.167 | 54935.9 | 66.374 | 366.929 |
| KES093 | 724.531 | 0.179 | 5.599 | 4.050 | 53891.4 | 64.852 | 366.022 |
| KES094 | 703.768 | 0.177 | 5.429 | 3.974 | 52605.8 | 63.450 | 365.940 |
| KES095 | 711.735 | 0.184 | 5.619 | 3.877 | 50130.3 | 65.270 | 369.356 |
| KES096 | 744.212 | 0.230 | 5.759 | 4.006 | 51942.1 | 67.887 | 382.265 |
| KES097 | 726.941 | 0.244 | 5.600 | 3.938 | 50699.1 | 66.212 | 373.513 |
| KES098 | 492.024 | 0.202 | 3.018 | 5.086 | 59112.3 | 40.555 | 248.619 |
| KES099 | 490.562 | 0.416 | 2.962 | 5.107 | 59245.6 | 40.484 | 247.493 |
| KES100 | 474.695 | 0.215 | 2.940 | 5.010 | 57837.6 | 39.498 | 244.276 |
| KES101 | 282.127 | 1.823 | 1.714 | 2.076 | 55460.7 | 23.246 | 153.099 |
| KES102 | 294.731 | 1.044 | 1.669 | 2.053 | 56310.9 | 23.461 | 151.747 |
| KES103 | 285.732 | 1.422 | 1.718 | 2.110 | 55017.0 | 22.949 | 154.906 |
| KES104 | 274.352 | 1.070 | 1.665 | 1.964 | 54707.7 | 22.616 | 148.935 |
| KES124 | 261.574 | 1.084 | 1.555 | 2.881 | 57931.7 | 21.693 | 139.947 |
| KES125 | 257.773 | 1.045 | 0.715 | 2.825 | 55533.3 | 21.200 | 136.247 |
| KES126 | 287.501 | 1.060 | 1.687 | 3.101 | 61908.6 | 24.106 | 152.692 |
| KES127 | 278.520 | 0.920 | 1.374 | 3.000 | 59777.3 | 23.435 | 148.856 |
| KES128 | 260.791 | 0.990 | 1.116 | 2.886 | 57312.3 | 22.016 | 140.213 |
| KES130 | 279.761 | 1.124 | 1.646 | 3.094 | 62114.5 | 23.284 | 151.402 |
| KES131 | 291.603 | 1.094 | 1.668 | 3.147 | 63423.5 | 24.227 | 154.407 |
| KES132 | 281.333 | 1.111 | 1.612 | 3.101 | 61472.1 | 23.380 | 148.828 |
| KES133 | 287.267 | 2.497 | 1.920 | 3.079 | 61385.5 | 21.927 | 162.890 |
| KES134 | 286.417 | 2.545 | 1.908 | 3.031 | 61099.2 | 21.723 | 162.389 |
| KES135 | 283.765 | 2.476 | 1.902 | 3.017 | 61092.6 | 21.522 | 161.330 |
| KES136 | 284.561 | 2.521 | 1.853 | 3.015 | 61241.0 | 21.764 | 162.997 |
| KES137 | 354.073 | 1.046 | 2.547 | 2.680 | 55447.3 | 28.667 | 202.890 |
| KES138 | 345.126 | 1.069 | 2.484 | 2.630 | 55169.3 | 28.407 | 201.321 |
| KES139 | 301.534 | 2.484 | 1.964 | 3.125 | 61126.9 | 23.539 | 172.592 |
| KES140 | 282.214 | 2.511 | 1.857 | 2.989 | 60797.4 | 21.440 | 158.111 |
| KES141 | 310.563 | 2.470 | 2.164 | 3.181 | 60663.0 | 24.459 | 176.204 |
| KES147 | 239.623 | 0.111 | 6.491 | 0.544 | 24450.7 | 52.150 | 105.067 |
| KES148 | 245.396 | 0.259 | 6.498 | 0.537 | 24883.1 | 53.237 | 105.202 |
| KES149 | 239.703 | 0.096 | 6.518 | 0.540 | 24478.9 | 52.144 | 106.339 |
| KES150 | 241.223 | 0.246 | 6.452 | 0.548 | 24651.1 | 52.245 | 106.601 |
| KES151 | 236.484 | 0.298 | 6.444 | 0.519 | 24206.3 | 51.624 | 106.194 |
| KES152 | 231.432 | 0.259 | 6.440 | 0.515 | 24109.4 | 51.247 | 103.363 |
| KES153 | 231.076 | 0.251 | 6.495 | 0.518 | 23963.3 | 50.870 | 103.493 |
| KES154 | 230.902 | 0.233 | 6.383 | 0.506 | 23841.6 | 50.816 | 102.123 |
| KES155 | 226.255 | 0.085 | 6.352 | 0.502 | 23633.7 | 50.217 | 102.947 |
| KES156 | 229.557 | 0.128 | 6.340 | 0.512 | 24101.7 | 51.048 | 102.566 |
| KES157 | 236.103 | 0.091 | 6.537 | 0.515 | 24433.6 | 51.901 | 104.844 |
| KES158 | 232.576 | 0.117 | 6.413 | 0.508 | 24151.4 | 51.357 | 103.255 |
| KES159 | 238.106 | 0.144 | 6.429 | 0.533 | 24190.7 | 51.595 | 107.398 |
| KES160 | 236.764 | 0.102 | 6.427 | 0.531 | 24230.7 | 51.704 | 105.954 |
| KES161 | 235.963 | 0.107 | 6.429 | 0.529 | 24231.0 | 51.428 | 106.703 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES162 | 240.029 | 0.235 | 6.439 | 0.541 | 24301.5 | 52.056 | 106.725 |
| KES163 | 234.358 | 0.233 | 6.401 | 0.525 | 24248.2 | 51.615 | 108.096 |
| KES164 | 232.665 | 0.263 | 6.349 | 0.516 | 23997.1 | 51.008 | 106.591 |
| KES165 | 488.312 | 0.294 | 3.208 | 2.487 | 59706.7 | 40.844 | 256.238 |
| KES166 | 502.484 | 0.300 | 3.354 | 2.619 | 62165.2 | 42.461 | 266.094 |
| KES167 | 507.894 | 0.274 | 3.260 | 2.615 | 62367.1 | 42.428 | 268.930 |
| KES168 | 436.839 | 0.066 | 2.374 | 3.807 | 61848.6 | 35.913 | 225.710 |
| KES169 | 329.666 | 0.090 | 3.285 | 1.808 | 41029.1 | 27.126 | 175.643 |
| KES170 | 374.882 | 0.307 | 2.781 | 1.544 | 37163.9 | 31.499 | 193.350 |
| KES171 | 371.173 | 0.307 | 2.768 | 1.531 | 37885.6 | 31.241 | 193.943 |
| KES172 | 375.029 | 0.065 | 2.872 | 1.501 | 36998.3 | 31.597 | 195.043 |
| KES173 | 377.597 | 0.051 | 2.829 | 1.505 | 37770.1 | 31.693 | 195.403 |
| KES174 | 373.212 | 0.076 | 2.821 | 1.541 | 37034.7 | 31.521 | 193.890 |
| KES175 | 381.443 | 0.049 | 2.912 | 1.563 | 37717.2 | 32.267 | 198.585 |
| KES176 | 129.592 | 0.354 | 4.023 | 0.156 | 12790.5 | 16.261 | 68.522 |
| KES177 | 133.351 | 0.470 | 4.091 | 0.162 | 12750.6 | 16.039 | 67.940 |
| KES178 | 136.454 | 0.578 | 4.237 | 0.160 | 13140.0 | 16.613 | 67.284 |
| KES179 | 139.463 | 0.368 | 4.245 | 0.163 | 13597.5 | 17.312 | 70.347 |
| KES180 | 140.945 | 0.455 | 4.077 | 0.176 | 13410.9 | 16.444 | 72.931 |
| KES181 | 136.691 | 0.639 | 4.061 | 0.182 | 13395.7 | 16.268 | 71.710 |
| KES182 | 702.639 | 0.334 | 5.673 | 3.906 | 49958.5 | 66.355 | 395.466 |
| KES183 | 718.563 | 0.164 | 5.649 | 3.933 | 50067.5 | 66.472 | 390.017 |
| KES184 | 688.618 | 0.165 | 5.544 | 3.882 | 49322.7 | 65.727 | 361.165 |
| KES185 | 676.920 | 0.384 | 5.429 | 3.870 | 48822.3 | 64.918 | 358.046 |
| KES186 | 675.684 | 0.367 | 5.549 | 3.835 | 48646.8 | 64.859 | 354.814 |
| KES187 | 674.194 | 0.159 | 5.574 | 3.825 | 48643.1 | 64.705 | 355.443 |
| KES188 | 665.798 | 0.159 | 5.472 | 3.820 | 48218.6 | 63.979 | 358.201 |
| KES189 | 651.812 | 0.169 | 5.429 | 3.757 | 47522.1 | 62.955 | 352.722 |
| KES190 | 669.828 | 0.371 | 5.525 | 3.839 | 48519.9 | 64.337 | 358.353 |
| KES191 | 664.533 | 0.416 | 5.542 | 3.765 | 48016.5 | 63.891 | 353.904 |
| KES192 | 675.939 | 0.110 | 5.558 | 3.904 | 48340.5 | 64.287 | 362.865 |
| KES193 | 682.794 | 0.342 | 5.642 | 3.886 | 48593.6 | 64.819 | 360.453 |
| KES194 | 683.537 | 0.116 | 5.578 | 3.960 | 48821.5 | 65.087 | 364.502 |
| KES195 | 667.477 | 0.298 | 5.541 | 3.812 | 48240.4 | 63.972 | 357.747 |
| KES196 | 684.563 | 0.369 | 5.560 | 3.872 | 49063.9 | 65.439 | 364.058 |
| KES197 | 675.593 | 0.317 | 5.504 | 3.887 | 48646.5 | 64.830 | 362.920 |
| KES198 | 677.769 | 0.329 | 5.578 | 3.826 | 48470.1 | 64.724 | 357.541 |
| KES199 | 668.114 | 0.149 | 5.509 | 3.796 | 48084.0 | 63.948 | 358.042 |
| KES200 | 675.377 | 0.167 | 5.529 | 3.834 | 48692.5 | 64.881 | 359.730 |
| KES201 | 685.342 | 0.178 | 5.661 | 3.923 | 49197.2 | 65.549 | 365.510 |
| KES202 | 677.928 | 0.183 | 5.540 | 3.839 | 48457.7 | 64.832 | 361.610 |
| KES203 | 682.263 | 0.301 | 5.535 | 3.867 | 49130.0 | 65.041 | 362.255 |
| KES204 | 675.513 | 0.333 | 5.605 | 3.842 | 48811.9 | 64.707 | 362.648 |
| KES205 | 676.376 | 0.206 | 5.672 | 3.906 | 49472.2 | 65.443 | 367.871 |
| KES206 | 673.658 | 0.316 | 5.454 | 3.805 | 48309.4 | 64.327 | 360.283 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES207 | 680.239 | 0.402 | 5.543 | 3.874 | 49085.2 | 65.171 | 367.858 |
| KES208 | 677.018 | 0.378 | 5.631 | 3.871 | 48835.4 | 64.883 | 364.725 |
| KES209 | 669.486 | 0.457 | 5.467 | 3.823 | 48578.6 | 63.208 | 354.401 |
| KES210 | 684.004 | 0.299 | 5.546 | 3.979 | 49008.9 | 65.148 | 369.807 |
| KES211 | 671.540 | 0.157 | 5.430 | 3.839 | 47877.2 | 63.650 | 361.324 |
| KES212 | 676.597 | 0.154 | 5.534 | 3.911 | 48394.6 | 64.209 | 394.681 |
| KES213 | 678.530 | 0.165 | 5.570 | 3.903 | 48378.1 | 64.430 | 394.838 |
| KES214 | 686.287 | 0.190 | 5.613 | 3.970 | 48911.5 | 65.158 | 400.019 |
| KES215 | 675.294 | 0.307 | 5.495 | 3.899 | 48157.8 | 64.066 | 394.837 |
| KES216 | 673.385 | 0.195 | 5.484 | 3.834 | 47946.5 | 63.879 | 394.060 |
| KES217 | 749.039 | 0.179 | 5.764 | 4.162 | 51805.4 | 68.223 | 392.723 |
| KES218 | 240.664 | 0.117 | 6.655 | 0.527 | 24790.1 | 52.814 | 106.902 |
| KES219 | 245.451 | 0.300 | 6.761 | 0.540 | 25280.2 | 53.521 | 108.412 |
| KES220 | 755.835 | 0.369 | 5.923 | 4.094 | 51460.0 | 68.204 | 382.751 |
| KES221 | 754.381 | 0.290 | 5.870 | 4.124 | 51851.9 | 68.401 | 388.254 |
| KES222 | 742.925 | 0.377 | 5.708 | 4.069 | 51213.8 | 67.675 | 385.848 |
| KES223 | 764.723 | 0.331 | 5.987 | 4.155 | 52225.5 | 68.912 | 391.878 |
| KES224 | 743.251 | 0.278 | 5.786 | 4.030 | 50798.7 | 67.264 | 382.674 |
| KES225 | 738.297 | 0.162 | 5.928 | 4.040 | 50862.1 | 67.028 | 381.820 |
| KES226 | 740.866 | 0.366 | 5.788 | 4.007 | 50606.4 | 66.834 | 381.878 |
| KES227 | 753.489 | 0.189 | 5.858 | 4.153 | 52214.8 | 68.638 | 390.940 |
| KES228 | 734.387 | 0.170 | 5.735 | 4.024 | 50449.0 | 66.621 | 384.075 |
| KES229 | 740.794 | 0.187 | 5.841 | 4.050 | 51477.1 | 68.076 | 388.751 |
| KES230 | 738.206 | 0.197 | 5.773 | 3.966 | 51000.6 | 67.600 | 382.624 |
| KES231 | 738.077 | 0.336 | 5.850 | 4.026 | 51463.3 | 67.759 | 385.834 |
| KES232 | 736.318 | 0.339 | 5.771 | 4.009 | 50494.1 | 66.814 | 380.393 |
| KES233 | 739.911 | 0.157 | 5.746 | 4.039 | 50543.6 | 67.014 | 386.282 |
| KES234 | 751.554 | 0.314 | 5.881 | 4.127 | 51274.6 | 68.085 | 387.568 |
| KES235 | 746.252 | 0.325 | 5.751 | 4.077 | 51005.7 | 67.545 | 388.756 |
| KES236 | 769.243 | 0.280 | 6.081 | 4.256 | 52733.7 | 69.656 | 400.789 |
| KES237 | 752.949 | 0.198 | 5.898 | 4.125 | 51453.5 | 68.135 | 389.994 |
| KES238 | 725.855 | 0.307 | 5.659 | 3.974 | 49858.2 | 66.079 | 381.103 |
| KES239 | 741.236 | 0.324 | 5.787 | 4.060 | 51200.5 | 67.355 | 389.168 |
| KES240 | 742.656 | 0.152 | 5.826 | 4.068 | 51078.7 | 67.473 | 388.982 |
| KES241 | 736.254 | 0.173 | 5.651 | 4.010 | 50526.4 | 66.657 | 383.552 |
| KES242 | 744.737 | 0.342 | 5.852 | 4.113 | 51437.6 | 67.903 | 391.571 |
| KES243 | 263.075 | 0.182 | 6.723 | 0.562 | 25312.6 | 41.827 | 125.916 |
| KES244 | 269.925 | 0.300 | 6.887 | 0.583 | 26063.4 | 42.850 | 130.889 |
| KES245 | 263.846 | 0.469 | 6.647 | 0.592 | 25979.9 | 41.947 | 127.761 |
| KES246 | 269.589 | 0.318 | 6.819 | 0.607 | 26404.8 | 42.681 | 128.263 |
| KES247 | 264.290 | 0.505 | 6.641 | 0.606 | 26695.3 | 41.564 | 128.546 |
| KES248 | 267.042 | 0.683 | 6.690 | 0.629 | 26762.4 | 41.675 | 128.230 |
| KES249 | 268.599 | 0.777 | 6.709 | 0.681 | 27877.7 | 41.843 | 130.045 |
| KES250 | 260.988 | 0.521 | 6.509 | 0.598 | 25520.0 | 40.436 | 123.948 |
| KES251 | 268.556 | 0.295 | 6.394 | 0.586 | 25589.9 | 39.863 | 127.128 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES252 | 260.402 | 0.338 | 6.568 | 0.567 | 25050.1 | 41.197 | 122.988 |
| KES253 | 266.977 | 0.318 | 6.754 | 0.570 | 25732.4 | 42.263 | 124.366 |
| KES254 | 265.644 | 0.366 | 6.621 | 0.579 | 25564.3 | 41.485 | 124.373 |
| KES255 | 286.418 | 1.557 | 3.181 | 1.781 | 23988.1 | 27.532 | 153.624 |
| KES256 | 288.986 | 1.469 | 3.066 | 1.786 | 23719.0 | 27.838 | 157.041 |
| KES257 | 293.380 | 1.751 | 3.229 | 1.788 | 24247.9 | 27.966 | 155.254 |
| KES258 | 281.858 | 1.539 | 3.141 | 1.751 | 22734.4 | 27.358 | 154.267 |
| KES259 | 289.978 | 1.484 | 3.194 | 1.751 | 24052.6 | 27.809 | 154.547 |
| KES260 | 293.508 | 1.561 | 3.205 | 1.791 | 24492.3 | 28.169 | 157.252 |
| KES261 | 277.492 | 1.650 | 3.189 | 1.713 | 23014.4 | 27.018 | 152.073 |
| KES262 | 288.481 | 1.481 | 3.239 | 1.784 | 23863.8 | 27.916 | 154.373 |
| KES263 | 294.293 | 1.670 | 3.205 | 1.787 | 24328.3 | 28.099 | 154.832 |
| KES264 | 290.266 | 1.454 | 3.185 | 1.772 | 23584.9 | 27.858 | 155.167 |
| KES265 | 293.066 | 1.544 | 3.198 | 1.810 | 24382.9 | 28.094 | 157.064 |
| KES266 | 288.476 | 1.510 | 3.175 | 1.767 | 23388.2 | 27.753 | 153.921 |
| KES267 | 290.353 | 1.467 | 3.196 | 1.787 | 23757.4 | 28.061 | 156.203 |
| KES268 | 292.424 | 1.486 | 3.239 | 1.771 | 23942.2 | 28.075 | 156.327 |
| KES269 | 286.005 | 1.408 | 3.169 | 1.753 | 23307.9 | 27.747 | 154.042 |
| KES270 | 289.376 | 1.497 | 3.214 | 1.794 | 24077.0 | 28.046 | 156.970 |
| KES271 | 280.832 | 1.527 | 3.093 | 1.761 | 23661.1 | 27.253 | 152.619 |
| KES272 | 287.110 | 1.549 | 3.201 | 1.775 | 24307.6 | 27.591 | 154.833 |
| KES273 | 285.414 | 1.492 | 3.123 | 1.782 | 23687.5 | 27.665 | 155.664 |
| KES274 | 288.332 | 1.539 | 3.190 | 1.771 | 24333.9 | 27.813 | 154.293 |
| KES275 | 290.256 | 1.667 | 3.234 | 1.782 | 23879.6 | 27.893 | 155.534 |
| KES276 | 283.339 | 1.483 | 3.199 | 1.763 | 22912.0 | 27.577 | 153.659 |
| KES277 | 287.644 | 1.555 | 3.154 | 1.790 | 23998.1 | 27.890 | 152.234 |
| KES278 | 261.656 | 0.408 | 6.686 | 0.614 | 26337.7 | 41.735 | 119.822 |
| KES279 | 259.445 | 0.373 | 6.722 | 0.591 | 25813.5 | 42.269 | 120.072 |
| KES281 | 260.095 | 1.114 | 6.600 | 0.734 | 27696.5 | 40.738 | 122.104 |
| KES282 | 256.538 | 0.473 | 6.373 | 0.609 | 26111.6 | 39.772 | 122.685 |
| KES283 | 266.238 | 0.580 | 6.278 | 0.642 | 26961.2 | 39.822 | 125.316 |
| KES284 | 258.363 | 0.704 | 6.144 | 0.660 | 27004.3 | 38.780 | 122.847 |
| KES285 | 404.852 | 0.179 | 1.622 | 2.822 | 64204.6 | 25.571 | 206.018 |
| KES286 | 412.497 | 0.205 | 1.543 | 2.845 | 64847.8 | 25.960 | 209.084 |
| KES287 | 361.499 | 0.197 | 1.421 | 2.007 | 64704.8 | 21.431 | 189.892 |
| KES288 | 369.520 | 0.235 | 1.584 | 2.011 | 65181.8 | 22.252 | 191.399 |
| KES289 | 371.944 | 0.208 | 1.517 | 2.027 | 65532.0 | 21.979 | 192.081 |
| KES290 | 365.733 | 0.390 | 1.531 | 2.042 | 64961.4 | 21.754 | 190.444 |
| KES291 | 379.488 | 0.232 | 1.528 | 2.090 | 67104.2 | 22.382 | 195.502 |
| KES292 | 361.139 | 0.170 | 1.486 | 2.007 | 65042.6 | 21.758 | 191.437 |
| KES293 | 377.867 | 0.160 | 1.491 | 2.084 | 66712.3 | 22.236 | 192.718 |
| KES294 | 373.704 | 0.187 | 1.605 | 2.050 | 66052.1 | 22.094 | 195.778 |
| KES295 | 375.513 | 0.211 | 1.560 | 2.070 | 66221.0 | 22.196 | 196.823 |
| KES296 | 375.905 | 0.398 | 1.572 | 2.053 | 66174.7 | 22.053 | 194.540 |
| KES297 | 246.064 | 0.113 | 6.669 | 0.543 | 25149.7 | 53.810 | 106.797 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES298 | 241.016 | 0.234 | 6.493 | 0.561 | 24723.6 | 52.677 | 106.104 |
| KES299 | 246.028 | 0.276 | 6.667 | 0.553 | 25192.7 | 53.650 | 106.150 |
| KES300 | 840.229 | 0.144 | 6.800 | 4.810 | 51919.4 | 81.414 | 405.028 |
| KES301 | 777.581 | 0.123 | 6.374 | 4.543 | 49353.1 | 76.327 | 384.529 |
| KES302 | 805.589 | 0.151 | 6.534 | 4.629 | 50384.8 | 78.382 | 398.208 |
| KES303 | 809.239 | 0.201 | 6.573 | 4.583 | 50237.0 | 78.438 | 396.230 |
| KES304 | 818.638 | 0.290 | 6.684 | 4.646 | 50332.7 | 78.687 | 397.558 |
| KES305 | 814.410 | 0.345 | 6.597 | 4.675 | 51126.5 | 79.128 | 401.544 |
| KES306 | 816.963 | 0.350 | 6.614 | 4.611 | 51115.5 | 78.752 | 397.312 |
| KES307 | 801.088 | 0.141 | 6.447 | 4.579 | 49869.6 | 77.486 | 395.355 |
| KES308 | 788.203 | 0.173 | 6.375 | 4.575 | 49579.8 | 76.372 | 391.841 |
| KES309 | 803.767 | 0.346 | 6.405 | 4.613 | 48538.7 | 78.096 | 398.202 |
| KES310 | 797.049 | 0.311 | 6.440 | 4.575 | 46690.1 | 78.441 | 391.617 |
| KES311 | 891.505 | 0.184 | 7.073 | 4.817 | 52477.2 | 82.618 | 457.949 |
| KES312 | 909.696 | 0.146 | 7.012 | 4.972 | 53671.5 | 84.067 | 460.665 |
| KES313 | 884.522 | 0.182 | 7.069 | 4.828 | 52000.8 | 82.060 | 454.018 |
| KES314 | 899.972 | 0.138 | 6.983 | 4.822 | 52840.5 | 83.253 | 456.249 |
| KES315 | 912.687 | 0.150 | 7.123 | 4.927 | 53598.8 | 84.305 | 464.390 |
| KES316 | 851.777 | 0.075 | 6.711 | 4.654 | 48445.4 | 79.975 | 435.181 |
| KES317 | 844.497 | 0.097 | 6.749 | 4.609 | 48432.5 | 79.514 | 439.193 |
| KES318 | 341.323 | 0.096 | 1.302 | 3.740 | 70582.2 | 27.875 | 174.140 |
| KES319 | 360.215 | 0.050 | 1.285 | 4.488 | 71379.6 | 28.872 | 184.811 |
| KES320 | 785.148 | 0.173 | 3.300 | 5.071 | 60251.2 | 65.100 | 392.674 |
| KES321 | 582.984 | 0.135 | 2.553 | 3.503 | 58922.3 | 50.657 | 288.757 |
| KES322 | 425.673 | 0.060 | 1.492 | 4.983 | 69628.8 | 33.855 | 210.195 |
| KES323 | 487.591 | 0.364 | 2.081 | 3.138 | 55612.9 | 41.125 | 247.952 |
| KES324 | 493.533 | 0.362 | 1.994 | 3.209 | 55646.2 | 41.494 | 245.593 |
| KES325 | 472.316 | 0.352 | 2.048 | 3.127 | 54737.8 | 40.375 | 245.029 |
| KES326 | 484.386 | 0.364 | 2.229 | 3.095 | 54935.4 | 40.753 | 245.184 |
| KES327 | 477.784 | 0.345 | 2.134 | 3.123 | 55017.3 | 40.662 | 241.985 |
| KES328 | 487.819 | 0.380 | 2.108 | 3.141 | 55378.4 | 41.020 | 247.956 |
| KES329 | 489.514 | 0.350 | 2.196 | 3.215 | 55870.5 | 41.346 | 247.087 |
| KES330 | 485.681 | 0.336 | 2.104 | 3.167 | 55402.1 | 40.960 | 245.848 |
| KES331 | 490.172 | 0.350 | 2.100 | 3.216 | 55817.0 | 41.284 | 246.860 |
| KES335 | 487.259 | 0.383 | 2.035 | 3.182 | 55551.9 | 41.327 | 248.798 |
| KES336 | 482.928 | 0.393 | 2.042 | 3.139 | 54949.5 | 40.732 | 244.417 |
| KES337 | 486.113 | 0.477 | 2.145 | 3.130 | 54984.6 | 40.908 | 245.163 |
| KES338 | 475.065 | 0.420 | 2.058 | 3.089 | 54335.3 | 40.286 | 245.663 |
| KES339 | 479.284 | 0.372 | 2.061 | 3.093 | 54966.5 | 40.592 | 247.596 |
| KES340 | 497.724 | 0.403 | 2.130 | 3.173 | 56022.5 | 41.686 | 250.101 |
| KES341 | 347.738 | 0.060 | 1.319 | 4.282 | 71862.3 | 27.724 | 178.384 |
| KES342 | 431.726 | 0.546 | 3.296 | 3.384 | 58303.4 | 36.439 | 220.945 |
| KES343 | 341.775 | 0.079 | 1.255 | 4.299 | 70746.1 | 27.354 | 176.276 |
| KES344 | 349.839 | 0.074 | 1.247 | 4.366 | 72716.7 | 27.852 | 179.651 |
| KES345 | 210.933 | 0.423 | 0.768 | 2.204 | 67102.1 | 16.479 | 107.474 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES346 | 207.012 | 0.305 | 0.686 | 2.216 | 66794.6 | 16.394 | 106.934 |
| KES347 | 589.286 | 0.166 | 4.844 | 3.177 | 45470.0 | 55.518 | 320.341 |
| KES348 | 598.913 | 0.193 | 4.865 | 3.254 | 45955.9 | 56.222 | 322.161 |
| KES349 | 605.057 | 0.175 | 4.869 | 3.290 | 46367.3 | 56.629 | 326.275 |
| KES350 | 872.591 | 0.095 | 7.180 | 4.547 | 54379.7 | 77.135 | 464.584 |
| KES351 | 871.388 | 0.104 | 7.108 | 4.534 | 54577.9 | 76.904 | 465.107 |
| KES352 | 872.336 | 0.060 | 7.111 | 4.513 | 54610.1 | 77.033 | 465.999 |
| KES353 | 856.781 | 0.030 | 7.104 | 4.507 | 54102.2 | 76.090 | 461.709 |
| KES354 | 871.155 | 0.261 | 7.140 | 4.556 | 54434.4 | 76.990 | 458.199 |
| KES355 | 858.583 | 0.000 | 7.079 | 4.491 | 54018.0 | 76.200 | 463.325 |
| KES356 | 860.499 | 0.331 | 7.093 | 4.469 | 53809.1 | 76.113 | 457.264 |
| KES357 | 867.187 | 0.321 | 7.210 | 4.535 | 54220.9 | 76.870 | 460.219 |
| KES358 | 854.876 | 0.000 | 7.117 | 4.484 | 54192.9 | 76.266 | 464.538 |
| KES359 | 851.691 | 0.138 | 7.009 | 4.404 | 53236.8 | 75.288 | 454.937 |
| KES360 | 858.322 | 0.000 | 7.031 | 4.470 | 53816.9 | 76.103 | 458.660 |
| KES361 | 861.206 | 0.070 | 7.080 | 4.463 | 53966.3 | 76.425 | 462.311 |
| KES362 | 854.660 | 0.128 | 7.073 | 4.456 | 53695.6 | 75.848 | 459.009 |
| KES363 | 867.332 | 0.109 | 7.182 | 4.511 | 54146.5 | 76.709 | 462.557 |
| KES364 | 883.545 | 0.169 | 6.735 | 4.855 | 52793.6 | 81.883 | 470.838 |
| KES365 | 849.226 | 0.144 | 6.637 | 4.659 | 51227.3 | 78.994 | 465.003 |
| KES366 | 859.672 | 0.360 | 6.708 | 4.769 | 52164.8 | 80.631 | 473.462 |
| KES367 | 854.978 | 0.179 | 6.665 | 4.714 | 51737.0 | 80.021 | 467.630 |
| KES368 | 861.233 | 0.212 | 6.655 | 4.749 | 51466.7 | 80.246 | 465.639 |
| KES369 | 889.347 | 0.343 | 6.830 | 4.856 | 53057.6 | 82.392 | 479.913 |
| KES370 | 489.814 | 0.346 | 2.958 | 5.121 | 59472.5 | 40.632 | 253.157 |
| KES371 | 488.367 | 0.224 | 2.920 | 5.046 | 59183.0 | 40.486 | 253.568 |
| KES372 | 487.786 | 0.215 | 2.911 | 5.125 | 59200.0 | 40.612 | 253.865 |
| KES373 | 479.139 | 0.333 | 2.920 | 5.039 | 58165.6 | 39.832 | 250.200 |
| KES374 | 248.996 | 0.155 | 5.430 | 0.484 | 23749.7 | 33.563 | 128.095 |
| KES375 | 252.436 | 0.168 | 5.414 | 0.494 | 23808.4 | 33.705 | 129.679 |
| KES376 | 267.859 | 0.317 | 6.725 | 0.598 | 26203.4 | 42.863 | 133.407 |
| KES377 | 270.092 | 0.173 | 6.833 | 0.594 | 26375.0 | 43.080 | 131.322 |
| KES378 | 265.136 | 0.145 | 6.714 | 0.576 | 25853.2 | 42.374 | 128.874 |
| KES379 | 258.728 | 0.145 | 6.572 | 0.573 | 25405.5 | 41.751 | 130.789 |
| KES380 | 245.995 | 0.185 | 6.484 | 0.566 | 24613.8 | 40.529 | 127.610 |
| KES381 | 249.863 | 0.151 | 6.612 | 0.563 | 24896.0 | 40.994 | 127.852 |
| KES382 | 251.710 | 0.149 | 6.575 | 0.564 | 24831.2 | 40.960 | 128.944 |
| KES383 | 251.120 | 0.174 | 6.608 | 0.571 | 24972.3 | 41.110 | 130.116 |
| KES384 | 255.996 | 0.155 | 6.523 | 0.568 | 24890.6 | 40.941 | 128.123 |
| KES385 | 250.210 | 0.311 | 6.524 | 0.578 | 24908.2 | 41.097 | 128.044 |
| KES386 | 248.737 | 0.156 | 6.490 | 0.555 | 24643.7 | 40.537 | 126.129 |
| KES387 | 251.688 | 0.328 | 6.389 | 0.560 | 24529.9 | 40.310 | 129.076 |
| KES388 | 251.510 | 0.157 | 6.420 | 0.557 | 24565.0 | 40.409 | 128.488 |
| KES389 | 249.310 | 0.335 | 6.470 | 0.563 | 24734.2 | 40.922 | 130.327 |
| KES390 | 247.681 | 0.132 | 6.392 | 0.551 | 24403.5 | 40.203 | 127.105 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES391 | 251.247 | 0.176 | 6.498 | 0.564 | 24741.5 | 40.809 | 128.962 |
| KES392 | 253.019 | 0.160 | 6.543 | 0.573 | 25142.3 | 41.609 | 130.245 |
| KES393 | 250.806 | 0.150 | 6.697 | 0.577 | 25041.5 | 41.073 | 130.019 |
| KES394 | 247.225 | 0.104 | 6.176 | 0.577 | 24464.7 | 43.705 | 127.072 |
| KES403 | 486.467 | 0.2570 | 2.9296 | 5.0011 | 58277.7 | 40.2599 | 244.251 |
| KES404 | 492.642 | 0.2617 | 2.9216 | 5.1354 | 59019.5 | 40.6947 | 246.922 |
| KES405 | 496.088 | 0.2489 | 2.8859 | 5.1116 | 59435.0 | 40.8493 | 249.392 |
| KES406 | 492.921 | 0.2619 | 2.9180 | 5.0411 | 58890.6 | 40.6383 | 248.745 |
| KES407 | 492.014 | 0.2739 | 2.8657 | 5.0631 | 59187.9 | 40.6957 | 248.605 |
| KES408 | 726.741 | 0.7268 | 4.4897 | 5.3765 | 65959.4 | 45.5118 | 404.621 |
| KES409 | 717.551 | 0.8110 | 4.3851 | 5.3382 | 65508.1 | 45.1032 | 400.764 |
| KES410 | 718.281 | 0.7200 | 4.4958 | 5.3602 | 65701.8 | 45.4123 | 404.650 |
| KES411 | 419.803 | 0.7192 | 2.2999 | 2.3044 | 75895.4 | 26.5238 | 223.103 |
| KES412 | 643.914 | 0.4881 | 4.3307 | 2.2235 | 45636.2 | 43.8973 | 356.511 |
| KES413 | 655.977 | 0.4898 | 4.4487 | 2.2407 | 46438.3 | 44.4313 | 358.820 |
| KES414 | 375.897 | 0.2285 | 1.5210 | 2.0063 | 65366.0 | 21.9288 | 203.186 |
| KES415 | 377.465 | 0.2452 | 1.5393 | 2.0058 | 65079.0 | 21.9472 | 199.972 |
| KES416 | 372.486 | 0.2707 | 1.4929 | 1.9767 | 64507.6 | 21.6447 | 197.812 |
| KES417 | 380.706 | 0.2823 | 1.5371 | 2.0517 | 66028.4 | 22.1768 | 205.057 |
| KES418 | 242.806 | 0.3284 | 1.4808 | 3.1884 | 69379.8 | 23.5812 | 124.467 |
| KES419 | 333.791 | 0.4650 | 1.6266 | 2.6826 | 55679.8 | 19.4531 | 182.704 |
| KES420 | 324.140 | 0.4542 | 1.6311 | 2.5796 | 54017.1 | 18.9756 | 179.215 |
| KES421 | 324.854 | 0.4877 | 1.5720 | 2.5990 | 54471.5 | 18.9129 | 178.740 |
| KES422 | 337.012 | 0.5011 | 1.6238 | 2.7368 | 56209.4 | 19.6176 | 185.349 |
| KES423 | 333.453 | 0.4861 | 1.6212 | 2.6984 | 55760.1 | 19.5538 | 182.946 |
| KES424 | 260.404 | 1.4437 | 1.9724 | 2.3730 | 67802.6 | 31.7827 | 132.152 |
| KES425 | 231.480 | 6.2314 | 1.5431 | 2.1659 | 65480.8 | 28.2157 | 118.214 |
| KES426 | 206.892 | 7.2157 | 1.5247 | 2.2244 | 64712.0 | 24.8744 | 104.611 |
| KES427 | 252.072 | 2.2371 | 1.7752 | 2.2798 | 65671.0 | 30.7514 | 127.142 |
| KES428 | 247.635 | 3.2687 | 1.8807 | 2.3197 | 67415.8 | 30.3210 | 124.981 |
| KES429 | 327.574 | 0.4274 | 1.6677 | 2.6223 | 54153.8 | 19.1471 | 179.479 |
| KES430 | 176.507 | 0.5468 | 1.1051 | 2.1795 | 38162.7 | 15.8853 | 93.446 |
| KES431 | 166.740 | 0.1256 | 0.9549 | 1.8907 | 28928.2 | 17.5674 | 82.883 |
| KES432 | 169.732 | 0.1517 | 1.0811 | 1.9435 | 29415.5 | 17.8835 | 85.102 |
| KES433 | 188.945 | 0.1299 | 1.1958 | 2.3035 | 35344.7 | 19.6021 | 94.921 |
| KES434 | 171.244 | 0.1422 | 0.9936 | 1.9434 | 29997.8 | 17.8930 | 85.009 |
| KES435 | 171.402 | 0.1843 | 1.0174 | 1.9358 | 30581.3 | 17.9761 | 84.233 |
| KES436 | 259.999 | 0.3587 | 1.3618 | 4.2634 | 68373.8 | 21.5980 | 136.053 |
| KES437 | 236.250 | 0.6242 | 0.9704 | 2.5644 | 52575.2 | 12.6514 | 128.775 |
| KES438 | 318.438 | 0.4647 | 1.5964 | 2.5708 | 53652.0 | 18.7670 | 175.487 |
| KES439 | 300.159 | 0.8916 | 2.0140 | 4.6251 | 74471.3 | 32.4764 | 154.879 |
| KES440 | 303.897 | 0.8612 | 2.0430 | 4.6500 | 74852.1 | 32.7879 | 157.553 |
| KES441 | 304.839 | 0.8724 | 2.1701 | 4.6635 | 74914.2 | 32.8523 | 157.504 |
| KES442 | 296.800 | 0.8797 | 1.5571 | 3.7616 | 64932.8 | 25.5202 | 154.997 |
| KES443 | 295.563 | 0.8905 | 1.4779 | 3.7199 | 64369.0 | 25.3865 | 153.422 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES444 | 638.754 | 0.4150 | 3.2783 | 7.6139 | 71965.0 | 60.7442 | 329.361 |
| KES445 | 625.134 | 0.4040 | 3.2382 | 7.3863 | 70795.5 | 59.2623 | 323.777 |
| KES446 | 1024.936 | 0.3887 | 5.5563 | 11.0579 | 65287.8 | 98.2014 | 525.392 |
| KES447 | 339.776 | 0.3107 | 1.5116 | 3.9984 | 75840.5 | 27.9877 | 177.412 |
| KES448 | 337.494 | 0.2698 | 1.4794 | 3.9646 | 76220.5 | 27.8689 | 178.822 |
| KES449 | 289.807 | 0.2924 | 2.4346 | 3.5756 | 67154.4 | 26.0329 | 154.839 |
| KES450 | 286.185 | 0.2820 | 2.3964 | 3.5539 | 66322.6 | 25.6371 | 153.604 |
| KES451 | 288.440 | 0.2621 | 2.3390 | 3.5972 | 66977.0 | 25.8714 | 153.699 |
| KES452 | 296.170 | 0.2597 | 2.4933 | 3.6602 | 68425.7 | 26.4383 | 156.823 |
| KES453 | 300.156 | 0.2723 | 2.4260 | 3.6847 | 68979.2 | 26.5120 | 158.635 |
| KES454 | 297.421 | 0.2716 | 2.4229 | 3.6662 | 68263.7 | 26.2454 | 155.765 |
| KES455 | 293.154 | 0.2918 | 2.4120 | 3.6221 | 67544.7 | 25.9535 | 154.212 |
| KES456 | 297.783 | 0.2717 | 2.4240 | 3.6880 | 68519.6 | 26.3000 | 156.421 |
| KES457 | 299.062 | 0.2654 | 2.3459 | 4.7079 | 75170.2 | 26.3715 | 155.576 |
| KES458 | 299.921 | 0.2519 | 2.2952 | 4.7734 | 75496.5 | 26.5808 | 155.147 |
| KES459 | 289.105 | 1.0176 | 2.7051 | 1.5747 | 43607.9 | 29.5758 | 156.653 |
| KES460 | 266.320 | 1.1343 | 2.4694 | 1.6716 | 43303.3 | 26.7121 | 143.182 |
| KES461 | 335.926 | 0.4114 | 1.6097 | 2.6849 | 55831.4 | 19.4527 | 182.233 |
| KES462 | 279.218 | 1.0740 | 2.5740 | 1.6243 | 43601.5 | 28.4820 | 152.353 |
| KES463 | 281.059 | 0.9900 | 2.0089 | 1.8088 | 46745.8 | 27.0382 | 150.097 |
| KES464 | 283.395 | 1.0813 | 1.9940 | 1.8387 | 47074.7 | 27.0494 | 150.900 |
| KES465 | 277.737 | 0.9772 | 1.9220 | 1.8107 | 46094.0 | 26.4905 | 148.219 |
| KES466 | 291.685 | 1.0618 | 2.0505 | 1.8381 | 47052.7 | 28.2645 | 157.350 |
| KES467 | 924.567 | 0.1247 | 7.1571 | 4.9938 | 53871.0 | 84.4805 | 481.986 |
| KES468 | 932.787 | 0.1685 | 7.2548 | 5.0855 | 54315.2 | 85.9033 | 492.297 |
| KES469 | 921.702 | 0.1717 | 7.0687 | 5.0106 | 53599.7 | 84.8702 | 486.175 |
| KES470 | 914.554 | 0.1599 | 7.1007 | 4.9412 | 53038.8 | 83.9410 | 480.136 |
| KES471 | 494.888 | 0.2193 | 1.9193 | 4.1314 | 64027.0 | 39.4652 | 252.986 |
| KES472 | 498.729 | 0.2560 | 1.9551 | 4.1939 | 64330.6 | 39.6907 | 249.238 |
| KES473 | 482.948 | 0.2429 | 1.8039 | 4.0350 | 62493.1 | 38.3971 | 247.270 |
| KES474 | 493.894 | 0.2248 | 1.7816 | 4.1439 | 63633.7 | 39.2695 | 253.266 |
| KES475 | 946.296 | 0.1257 | 7.2627 | 5.1176 | 54568.6 | 87.0855 | 502.394 |
| KES476 | 930.700 | 0.1636 | 7.2705 | 4.9982 | 53645.9 | 85.2153 | 487.311 |
| KES477 | 490.673 | 0.2437 | 1.8380 | 4.1331 | 63324.5 | 39.0499 | 247.745 |
| KES478 | 489.457 | 0.2269 | 1.8357 | 4.0730 | 63187.9 | 38.9259 | 249.669 |
| KES479 | 761.430 | 0.1403 | 6.0399 | 4.1014 | 48893.2 | 69.5215 | 406.674 |
| KES480 | 444.800 | 0.2229 | 2.1767 | 3.3469 | 66467.8 | 38.0823 | 226.477 |
| KES481 | 432.879 | 0.4856 | 3.1608 | 3.3816 | 58259.0 | 35.9463 | 220.013 |
| KES482 | 439.321 | 0.4631 | 3.2638 | 3.4285 | 58994.5 | 36.3817 | 225.215 |
| KES483 | 858.259 | 0.1964 | 6.6864 | 4.6758 | 50832.5 | 79.3374 | 458.662 |
| KES484 | 869.026 | 0.2861 | 6.7826 | 4.7459 | 51527.0 | 80.5053 | 468.331 |
| KES485 | 754.818 | 0.2683 | 5.8947 | 4.1411 | 51484.9 | 68.1427 | 401.768 |
| KES486 | 744.204 | 0.2529 | 5.8495 | 4.0931 | 50949.0 | 67.5158 | 400.577 |
| KES487 | 701.275 | 0.2770 | 5.4358 | 3.7615 | 48695.0 | 63.8773 | 373.230 |
| KES488 | 687.163 | 0.2870 | 5.3880 | 3.7130 | 48104.3 | 62.8627 | 369.530 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES489 | 713.869 | 0.2230 | 5.4407 | 4.0011 | 52517.8 | 63.6055 | 381.027 |
| KES490 | 732.300 | 0.2386 | 5.7009 | 4.0902 | 53985.6 | 65.2457 | 392.363 |
| KES491 | 722.012 | 0.2533 | 5.5238 | 4.0630 | 53184.8 | 64.4066 | 382.910 |
| KES492 | 720.713 | 0.2433 | 5.5532 | 4.0393 | 53182.6 | 64.2261 | 385.339 |
| KES493 | 717.902 | 0.2558 | 5.4753 | 4.0565 | 53012.5 | 63.9890 | 385.138 |
| KES494 | 718.577 | 0.2198 | 5.4917 | 4.0504 | 53479.7 | 64.1610 | 383.422 |
| KES495 | 707.907 | 0.2410 | 5.4774 | 3.8022 | 49226.4 | 64.4533 | 380.340 |
| KES496 | 409.515 | 0.2869 | 4.1606 | 1.8791 | 46476.6 | 38.0303 | 214.105 |
| KES497 | 431.731 | 0.2750 | 4.3024 | 1.9637 | 48497.3 | 39.6523 | 222.280 |
| KES498 | 422.068 | 0.2964 | 4.2701 | 1.9597 | 48002.1 | 39.4352 | 221.547 |
| KES499 | 404.380 | 0.2571 | 4.0891 | 1.8422 | 45625.9 | 37.5251 | 212.790 |
| KES500 | 405.510 | 0.2757 | 4.0133 | 1.8892 | 46817.7 | 37.4167 | 210.083 |
| KES501 | 421.100 | 0.2849 | 4.2412 | 1.9608 | 47954.3 | 39.1539 | 216.401 |
| KES502 | 406.173 | 0.2722 | 4.0543 | 1.8504 | 45835.1 | 37.7360 | 202.273 |
| KES503 | 334.261 | 0.8868 | 2.3730 | 1.2549 | 35592.4 | 28.3883 | 186.864 |
| KES504 | 331.715 | 0.8812 | 2.3841 | 1.2361 | 35308.5 | 28.1291 | 186.026 |
| KES505 | 336.230 | 0.9007 | 2.3898 | 1.2698 | 35795.0 | 28.6112 | 188.294 |
| KES506 | 338.685 | 0.9395 | 2.4128 | 1.2740 | 35690.7 | 28.7538 | 188.469 |
| KES507 | 337.276 | 0.9331 | 2.4312 | 1.2755 | 35922.4 | 28.7590 | 189.906 |
| KES508 | 339.127 | 0.9027 | 2.4672 | 1.2702 | 35904.4 | 28.8399 | 189.108 |
| KES509 | 289.227 | 1.1987 | 1.6381 | 3.1342 | 63153.4 | 23.9493 | 154.861 |
| KES510 | 299.156 | 1.1077 | 1.7352 | 3.2563 | 64171.6 | 24.9823 | 157.620 |
| KES511 | 288.954 | 1.1761 | 1.6808 | 3.1610 | 62909.5 | 24.1392 | 154.114 |
| KES512 | 284.276 | 1.2159 | 1.6592 | 3.1049 | 61917.5 | 23.5504 | 152.603 |
| KES513 | 278.451 | 1.0673 | 1.0979 | 2.9371 | 58894.9 | 22.5376 | 145.801 |
| KES514 | 278.422 | 0.2370 | 2.2276 | 3.4793 | 36240.2 | 22.7990 | 149.203 |
| KES515 | 282.312 | 0.2321 | 2.1872 | 3.4810 | 36799.8 | 22.9194 | 151.712 |
| KES516 | 163.245 | 7.2924 | 1.8091 | 2.1223 | 41060.3 | 18.9089 | 84.508 |
| KES517 | 166.192 | 7.0237 | 1.8843 | 2.0909 | 41132.3 | 19.5267 | 85.051 |
| KES518 | 166.305 | 7.0816 | 1.8464 | 2.0975 | 41141.5 | 19.2126 | 86.125 |
| KES519 | 158.407 | 7.7372 | 1.7756 | 2.1066 | 42404.7 | 18.1503 | 81.520 |
| KES520 | 161.341 | 7.4950 | 1.8375 | 2.0666 | 41473.8 | 18.7293 | 83.670 |
| KES521 | 157.648 | 7.8827 | 1.7883 | 2.0068 | 40569.6 | 18.3611 | 81.281 |
| KES522 | 161.396 | 6.5254 | 1.8343 | 2.0613 | 39392.7 | 18.7367 | 84.250 |
| KES523 | 164.136 | 6.6702 | 1.8259 | 2.1011 | 40118.8 | 19.1259 | 85.607 |
| KES524 | 161.721 | 6.7676 | 1.8593 | 2.0301 | 40397.1 | 18.8097 | 85.073 |
| KES525 | 163.924 | 7.5844 | 1.7826 | 2.0447 | 42297.8 | 19.1245 | 85.269 |
| KES526 | 145.702 | 11.4491 | 1.5670 | 2.1028 | 47043.6 | 16.6627 | 75.931 |
| KES527 | 213.709 | 0.5137 | 0.8968 | 2.6975 | 39905.3 | 14.5192 | 111.592 |
| KES528 | 215.551 | 0.5475 | 0.8343 | 2.7013 | 39710.2 | 14.5712 | 112.840 |
| KES529 | 210.900 | 0.4857 | 0.8240 | 2.6545 | 39135.3 | 14.3703 | 112.685 |
| KES530 | 206.518 | 0.5778 | 0.7436 | 2.6315 | 37934.3 | 13.9760 | 109.309 |
| KES531 | 202.354 | 0.6050 | 0.7075 | 2.5878 | 37165.0 | 13.7184 | 106.783 |
| KES532 | 203.711 | 0.6163 | 0.8062 | 2.5870 | 38506.9 | 13.9290 | 107.978 |
| KES533 | 205.484 | 0.6047 | 0.7410 | 2.6173 | 38046.9 | 13.9551 | 108.137 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ce | Co | Cs | Eu | Fe | Hf | La |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES534 | 269.101 | 0.1951 | 6.5472 | 0.6008 | 25816.0 | 46.1431 | 128.863 |
| KES535 | 269.233 | 0.1853 | 6.5926 | 0.5955 | 25818.7 | 46.1793 | 127.236 |
| KES536 | 495.813 | 0.2608 | 2.9657 | 5.1515 | 59644.4 | 41.1328 | 253.553 |
| KES537 | 488.041 | 0.2596 | 2.9056 | 5.0264 | 58845.0 | 40.4841 | 250.477 |
| KES538 | 485.529 | 0.3186 | 2.9006 | 5.0231 | 58578.6 | 40.3018 | 249.821 |
| KES539 | 490.250 | 0.2684 | 2.9248 | 5.0667 | 59077.9 | 40.7099 | 250.841 |
| KES540 | 417.321 | 0.7433 | 2.3296 | 2.3385 | 75777.5 | 26.4402 | 227.744 |
| KES541 | 229.448 | 7.0877 | 1.5654 | 2.1750 | 66381.8 | 27.3268 | 115.956 |
| KES542 | 249.816 | 5.1953 | 1.8824 | 2.2955 | 68982.2 | 30.5587 | 128.588 |

Appendix 6 - Concentration data from NAA for obsidian from Kenya. All concentrations are in parts per million ( ppm )

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES001 | 2.702 | 153.652 | 259.9 | 0.619 | 0.340 | 31.461 | 0.00 |
| KES002 | 2.597 | 158.509 | 260.2 | 0.625 | 0.390 | 31.404 | 0.00 |
| KES003 | 1.750 | 48.772 | 282.6 | 0.381 | 0.501 | 11.827 | 0.00 |
| KES004 | 1.774 | 51.017 | 269.6 | 0.416 | 0.620 | 12.090 | 0.00 |
| KES005 | 5.406 | 296.992 | 431.5 | 0.722 | 0.343 | 56.940 | 0.00 |
| KES006 | 6.289 | 326.999 | 461.7 | 0.831 | 0.171 | 64.926 | 0.00 |
| KES007 | 5.825 | 318.477 | 450.1 | 0.779 | 0.324 | 60.666 | 0.00 |
| KES008 | 4.008 | 203.554 | 219.9 | 0.406 | 0.359 | 41.971 | 0.00 |
| KES009 | 3.954 | 203.257 | 213.6 | 0.400 | 0.354 | 41.728 | 0.00 |
| KES010 | 6.258 | 335.873 | 480.8 | 0.757 | 0.316 | 65.978 | 0.00 |
| KES011 | 5.024 | 334.027 | 472.5 | 0.751 | 0.329 | 65.121 | 0.00 |
| KES012 | 4.858 | 274.702 | 390.4 | 0.664 | 0.356 | 51.822 | 0.00 |
| KES013 | 3.924 | 202.296 | 228.6 | 0.435 | 0.528 | 40.235 | 0.00 |
| KES014 | 2.258 | 190.371 | 216.6 | 0.559 | 1.106 | 35.099 | 0.00 |
| KES015 | 2.342 | 195.490 | 220.8 | 0.572 | 1.027 | 36.072 | 0.00 |
| KES016 | 1.772 | 94.630 | 210.4 | 0.682 | 2.809 | 16.168 | 47.90 |
| KES017 | 1.806 | 95.292 | 212.1 | 0.676 | 2.849 | 16.314 | 55.50 |
| KES018 | 1.779 | 93.279 | 213.1 | 0.661 | 2.891 | 16.318 | 60.20 |
| KES019 | 1.768 | 95.472 | 209.2 | 0.664 | 2.798 | 16.018 | 50.60 |
| KES020 | 2.012 | 97.183 | 173.2 | 0.618 | 3.656 | 16.678 | 0.00 |
| KES021 | 2.462 | 101.397 | 399.7 | 0.830 | 0.491 | 24.266 | 0.00 |
| KES022 | 2.400 | 97.058 | 383.7 | 0.715 | 0.776 | 23.529 | 0.00 |
| KES023 | 2.553 | 101.677 | 403.2 | 0.812 | 0.221 | 24.282 | 0.00 |
| KES024 | 2.474 | 101.582 | 400.2 | 0.785 | 0.246 | 24.022 | 0.00 |
| KES025 | 2.551 | 105.486 | 417.6 | 0.834 | 0.169 | 24.858 | 0.00 |
| KES026 | 2.572 | 103.206 | 410.0 | 0.786 | 0.142 | 24.448 | 0.00 |
| KES027 | 2.590 | 103.323 | 410.7 | 0.743 | 0.286 | 24.635 | 0.00 |
| KES028 | 2.581 | 103.282 | 410.7 | 0.000 | 0.261 | 24.688 | 0.00 |
| KES029 | 3.330 | 102.302 | 411.4 | 0.000 | 0.143 | 24.617 | 0.00 |
| KES030 | 2.597 | 91.707 | 395.3 | 0.701 | 0.131 | 23.212 | 0.00 |
| KES031 | 2.592 | 89.846 | 394.0 | 0.687 | 0.130 | 23.225 | 0.00 |
| KES032 | 2.573 | 87.087 | 394.7 | 0.708 | 0.134 | 23.220 | 0.00 |
| KES033 | 3.467 | 180.763 | 213.5 | 0.327 | 0.501 | 37.047 | 0.00 |
| KES034 | 3.446 | 185.444 | 214.6 | 0.368 | 0.509 | 37.302 | 0.00 |
| KES035 | 3.425 | 180.507 | 217.4 | 0.365 | 0.526 | 37.396 | 0.00 |
| KES036 | 3.428 | 181.198 | 219.3 | 0.377 | 0.507 | 37.081 | 0.00 |
| KES037 | 3.494 | 185.476 | 216.7 | 0.355 | 0.511 | 37.473 | 0.00 |
| KES038 | 3.508 | 182.982 | 215.4 | 0.358 | 0.515 | 37.504 | 0.00 |
| KES039 | 3.490 | 182.476 | 216.0 | 0.364 | 0.509 | 36.967 | 0.00 |
| KES040 | 3.380 | 185.993 | 214.2 | 0.373 | 0.512 | 36.874 | 0.00 |
| KES042 | 3.788 | 176.792 | 272.5 | 0.429 | 0.502 | 35.088 | 0.00 |
| KES043 | 3.527 | 194.901 | 209.7 | 0.333 | 0.321 | 39.314 | 0.00 |
| KES044 | 3.505 | 192.630 | 207.8 | 0.349 | 0.326 | 39.060 | 0.00 |
| KES045 | 3.633 | 193.175 | 217.2 | 0.332 | 0.285 | 39.615 | 0.00 |
| KES046 | 3.676 | 195.843 | 217.2 | 0.313 | 0.289 | 39.852 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES047 | 2.438 | 89.670 | 376.7 | 0.664 | 0.132 | 22.045 | 0.00 |
| KES048 | 2.663 | 96.161 | 395.7 | 0.723 | 0.130 | 23.606 | 0.00 |
| KES049 | 2.625 | 94.631 | 397.9 | 0.688 | 0.132 | 23.567 | 0.00 |
| KES050 | 2.832 | 98.715 | 393.1 | 0.661 | 0.122 | 25.128 | 0.00 |
| KES051 | 2.827 | 97.124 | 391.1 | 0.681 | 0.121 | 25.123 | 0.00 |
| KES052 | 3.374 | 120.712 | 469.0 | 0.832 | 0.145 | 30.164 | 0.00 |
| KES053 | 3.366 | 121.209 | 467.0 | 0.857 | 0.153 | 30.229 | 0.00 |
| KES054 | 3.778 | 182.861 | 215.3 | 0.330 | 0.164 | 36.517 | 0.00 |
| KES055 | 3.804 | 179.480 | 220.3 | 0.315 | 0.156 | 36.786 | 0.00 |
| KES056 | 3.809 | 175.392 | 222.2 | 0.321 | 0.165 | 37.076 | 0.00 |
| KES057 | 2.266 | 148.328 | 166.3 | 0.290 | 3.282 | 26.239 | 0.00 |
| KES058 | 4.578 | 280.083 | 404.7 | 0.571 | 0.295 | 53.928 | 0.00 |
| KES059 | 5.580 | 324.638 | 442.1 | 0.726 | 0.163 | 63.825 | 0.00 |
| KES060 | 3.873 | 219.239 | 336.0 | 0.424 | 0.355 | 41.078 | 0.00 |
| KES061 | 3.785 | 216.552 | 340.5 | 0.441 | 0.363 | 40.905 | 0.00 |
| KES062 | 4.138 | 255.970 | 366.7 | 0.504 | 0.338 | 47.781 | 0.00 |
| KES063 | 4.383 | 215.762 | 365.6 | 0.507 | 0.329 | 48.358 | 0.00 |
| KES064 | 4.404 | 205.067 | 365.3 | 0.511 | 0.324 | 48.765 | 0.00 |
| KES065 | 4.762 | 239.094 | 392.7 | 0.571 | 0.315 | 52.005 | 0.00 |
| KES066 | 2.014 | 103.378 | 146.9 | 0.300 | 4.707 | 19.525 | 0.00 |
| KES067 | 4.518 | 213.575 | 373.1 | 0.514 | 0.382 | 48.028 | 0.00 |
| KES068 | 4.522 | 219.087 | 371.9 | 0.545 | 0.350 | 48.499 | 0.00 |
| KES069 | 4.652 | 235.911 | 374.3 | 0.528 | 0.404 | 48.657 | 0.00 |
| KES070 | 5.991 | 274.205 | 446.9 | 0.735 | 0.164 | 62.714 | 0.00 |
| KES071 | 5.955 | 284.512 | 445.6 | 0.693 | 0.163 | 63.182 | 0.00 |
| KES072 | 5.873 | 282.029 | 442.1 | 0.679 | 0.159 | 62.472 | 0.00 |
| KES073 | 5.883 | 285.825 | 446.5 | 0.735 | 0.165 | 62.185 | 0.00 |
| KES074 | 5.865 | 286.036 | 444.3 | 0.696 | 0.160 | 62.698 | 0.00 |
| KES075 | 3.488 | 92.979 | 414.4 | 0.487 | 0.106 | 26.230 | 0.00 |
| KES076 | 1.430 | 50.802 | 275.7 | 0.349 | 0.496 | 11.708 | 0.00 |
| KES077 | 1.720 | 48.498 | 290.8 | 0.370 | 0.473 | 12.002 | 0.00 |
| KES078 | 1.450 | 47.066 | 277.5 | 0.370 | 0.604 | 11.753 | 0.00 |
| KES079 | 5.263 | 291.479 | 410.3 | 0.578 | 0.327 | 56.151 | 0.00 |
| KES080 | 1.782 | 103.629 | 98.3 | 0.135 | 2.320 | 20.010 | 0.00 |
| KES081 | 3.810 | 203.516 | 216.4 | 0.351 | 0.509 | 39.960 | 0.00 |
| KES082 | 3.872 | 216.011 | 218.9 | 0.350 | 0.510 | 40.648 | 0.00 |
| KES083 | 3.825 | 212.566 | 220.0 | 0.351 | 0.515 | 40.138 | 0.00 |
| KES084 | 3.861 | 210.376 | 215.8 | 0.378 | 0.514 | 40.331 | 0.00 |
| KES085 | 3.858 | 206.341 | 217.8 | 0.357 | 0.509 | 40.114 | 0.00 |
| KES086 | 3.884 | 218.951 | 213.9 | 0.359 | 0.519 | 40.659 | 0.00 |
| KES087 | 5.749 | 309.200 | 443.2 | 0.647 | 0.322 | 61.493 | 0.00 |
| KES088 | 5.673 | 314.314 | 434.4 | 0.613 | 0.316 | 60.723 | 0.00 |
| KES089 | 2.321 | 126.563 | 137.2 | 0.226 | 0.273 | 26.077 | 0.00 |
| KES090 | 2.390 | 134.783 | 140.1 | 0.232 | 0.268 | 26.060 | 0.00 |
| KES091 | 4.245 | 226.588 | 378.3 | 0.525 | 0.427 | 42.480 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES092 | 4.400 | 251.942 | 378.9 | 0.531 | 0.343 | 44.268 | 0.00 |
| KES093 | 4.355 | 247.486 | 374.2 | 0.519 | 0.339 | 44.426 | 0.00 |
| KES094 | 4.461 | 252.092 | 366.4 | 0.515 | 0.332 | 44.502 | 0.00 |
| KES095 | 4.327 | 240.376 | 378.2 | 0.483 | 0.362 | 48.446 | 0.00 |
| KES096 | 4.422 | 244.933 | 392.7 | 0.495 | 0.377 | 49.781 | 0.00 |
| KES097 | 4.293 | 244.135 | 383.8 | 0.497 | 0.361 | 49.633 | 0.00 |
| KES098 | 2.093 | 177.722 | 212.2 | 0.475 | 0.989 | 35.600 | 0.00 |
| KES099 | 2.111 | 185.699 | 209.6 | 0.439 | 0.988 | 35.671 | 0.00 |
| KES100 | 2.053 | 182.144 | 212.2 | 0.487 | 0.959 | 34.875 | 0.00 |
| KES101 | 1.670 | 96.664 | 149.8 | 0.322 | 4.332 | 18.064 | 0.00 |
| KES102 | 1.696 | 99.498 | 158.2 | 0.321 | 4.350 | 18.574 | 0.00 |
| KES103 | 1.665 | 98.985 | 151.0 | 0.330 | 4.292 | 18.276 | 0.00 |
| KES104 | 1.591 | 98.831 | 153.1 | 0.304 | 4.227 | 17.785 | 0.00 |
| KES124 | 1.628 | 92.481 | 129.9 | 0.331 | 4.807 | 17.680 | 108.31 |
| KES125 | 1.556 | 89.382 | 124.3 | 0.339 | 4.334 | 17.035 | 209.95 |
| KES126 | 1.730 | 99.857 | 152.6 | 0.301 | 4.667 | 19.243 | 0.00 |
| KES127 | 1.690 | 96.766 | 140.8 | 0.278 | 4.390 | 18.157 | 0.00 |
| KES128 | 1.593 | 97.178 | 129.5 | 0.315 | 4.420 | 17.573 | 40.94 |
| KES130 | 1.684 | 107.250 | 148.4 | 0.299 | 4.771 | 18.852 | 0.00 |
| KES131 | 1.758 | 106.866 | 153.7 | 0.345 | 4.704 | 19.230 | 0.00 |
| KES132 | 1.653 | 101.295 | 145.4 | 0.336 | 4.784 | 18.576 | 0.00 |
| KES133 | 1.563 | 90.582 | 165.3 | 0.595 | 3.504 | 16.356 | 0.00 |
| KES134 | 1.619 | 89.329 | 169.4 | 0.545 | 3.506 | 16.335 | 87.93 |
| KES135 | 1.560 | 87.145 | 168.1 | 0.579 | 3.513 | 16.044 | 52.03 |
| KES136 | 1.518 | 88.150 | 166.1 | 0.560 | 3.503 | 16.267 | 0.00 |
| KES137 | 2.045 | 112.484 | 200.6 | 0.715 | 1.790 | 19.942 | 0.00 |
| KES138 | 1.996 | 109.732 | 203.7 | 0.683 | 1.825 | 19.109 | 0.00 |
| KES139 | 1.711 | 96.895 | 168.8 | 0.596 | 3.421 | 17.112 | 50.53 |
| KES140 | 1.501 | 85.024 | 167.1 | 0.559 | 3.480 | 15.828 | 0.00 |
| KES141 | 1.719 | 96.375 | 175.2 | 0.665 | 3.346 | 17.414 | 0.00 |
| KES147 | 3.143 | 96.102 | 420.0 | 0.533 | 0.114 | 26.567 | 0.00 |
| KES148 | 3.174 | 95.037 | 427.5 | 0.577 | 0.114 | 26.689 | 0.00 |
| KES149 | 3.080 | 96.772 | 418.9 | 0.574 | 0.111 | 27.231 | 0.00 |
| KES150 | 3.146 | 99.744 | 418.3 | 0.539 | 0.107 | 27.276 | 0.00 |
| KES151 | 3.163 | 98.081 | 414.6 | 0.553 | 0.105 | 26.365 | 0.00 |
| KES152 | 3.169 | 96.481 | 411.7 | 0.556 | 0.110 | 26.194 | 0.00 |
| KES153 | 3.161 | 92.257 | 410.7 | 0.569 | 0.111 | 26.362 | 0.00 |
| KES154 | 3.147 | 91.763 | 412.6 | 0.571 | 0.109 | 25.946 | 0.00 |
| KES155 | 3.231 | 91.450 | 408.1 | 0.541 | 0.111 | 26.088 | 0.00 |
| KES156 | 3.124 | 93.884 | 411.6 | 0.564 | 0.108 | 25.601 | 0.00 |
| KES157 | 3.203 | 94.376 | 416.8 | 0.557 | 0.111 | 26.671 | 0.00 |
| KES158 | 3.129 | 93.066 | 411.0 | 0.591 | 0.113 | 26.253 | 0.00 |
| KES159 | 3.261 | 94.934 | 413.5 | 0.538 | 0.110 | 26.998 | 0.00 |
| KES160 | 3.203 | 100.033 | 410.1 | 0.551 | 0.106 | 26.968 | 0.00 |
| KES161 | 3.163 | 98.434 | 410.1 | 0.514 | 0.109 | 26.564 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES162 | 3.234 | 102.016 | 415.1 | 0.517 | 0.104 | 26.676 | 0.00 |
| KES163 | 3.255 | 100.782 | 412.9 | 0.546 | 0.111 | 27.262 | 0.00 |
| KES164 | 3.244 | 95.395 | 411.1 | 0.512 | 0.109 | 26.448 | 0.00 |
| KES165 | 3.887 | 174.367 | 282.0 | 0.463 | 0.430 | 35.513 | 0.00 |
| KES166 | 3.915 | 188.755 | 292.1 | 0.474 | 0.444 | 38.387 | 0.00 |
| KES167 | 3.953 | 187.695 | 293.3 | 0.438 | 0.456 | 38.352 | 0.00 |
| KES168 | 3.181 | 170.841 | 243.6 | 0.322 | 0.209 | 34.744 | 0.00 |
| KES169 | 2.094 | 134.304 | 222.6 | 0.285 | 0.226 | 24.658 | 0.00 |
| KES170 | 2.252 | 143.660 | 228.5 | 0.314 | 0.096 | 27.645 | 0.00 |
| KES171 | 2.256 | 143.168 | 231.4 | 0.303 | 0.112 | 27.364 | 0.00 |
| KES172 | 2.245 | 145.182 | 230.7 | 0.322 | 0.091 | 27.515 | 0.00 |
| KES173 | 2.241 | 146.721 | 233.2 | 0.321 | 0.094 | 27.672 | 0.00 |
| KES174 | 2.249 | 146.261 | 229.4 | 0.352 | 0.092 | 27.339 | 0.00 |
| KES175 | 2.299 | 151.524 | 236.3 | 0.312 | 0.092 | 27.896 | 0.00 |
| KES176 | 1.377 | 47.636 | 269.2 | 0.388 | 0.465 | 11.410 | 0.00 |
| KES177 | 1.375 | 46.068 | 277.6 | 0.379 | 0.451 | 11.669 | 0.00 |
| KES178 | 1.406 | 47.566 | 283.0 | 0.402 | 0.474 | 11.771 | 0.00 |
| KES179 | 1.474 | 48.373 | 283.2 | 0.403 | 0.467 | 12.291 | 0.00 |
| KES180 | 1.415 | 49.448 | 278.6 | 0.387 | 0.592 | 12.262 | 0.00 |
| KES181 | 1.392 | 50.176 | 274.9 | 0.353 | 0.583 | 11.941 | 0.00 |
| KES182 | 4.389 | 275.300 | 377.5 | 0.577 | 0.355 | 52.104 | 0.00 |
| KES183 | 4.349 | 277.449 | 377.6 | 0.565 | 0.340 | 50.716 | 0.00 |
| KES184 | 3.490 | 225.382 | 371.7 | 0.543 | 0.333 | 41.020 | 0.00 |
| KES185 | 3.904 | 225.850 | 366.1 | 0.537 | 0.337 | 40.163 | 0.00 |
| KES186 | 4.014 | 239.897 | 369.5 | 0.558 | 0.334 | 40.165 | 0.00 |
| KES187 | 4.051 | 245.308 | 371.7 | 0.540 | 0.329 | 40.588 | 0.00 |
| KES188 | 3.552 | 240.899 | 366.6 | 0.531 | 0.318 | 41.435 | 0.00 |
| KES189 | 3.896 | 238.717 | 359.9 | 0.535 | 0.319 | 40.044 | 0.00 |
| KES190 | 3.908 | 226.516 | 365.6 | 0.557 | 0.328 | 40.735 | 0.00 |
| KES191 | 4.067 | 242.398 | 364.4 | 0.534 | 0.337 | 40.566 | 0.00 |
| KES192 | 4.046 | 235.335 | 365.9 | 0.564 | 0.294 | 42.069 | 0.00 |
| KES193 | 4.105 | 249.748 | 367.2 | 0.545 | 0.289 | 42.012 | 0.00 |
| KES194 | 3.940 | 259.460 | 370.0 | 0.560 | 0.296 | 42.351 | 0.00 |
| KES195 | 3.931 | 246.391 | 365.1 | 0.514 | 0.325 | 41.518 | 0.00 |
| KES196 | 4.085 | 248.741 | 373.4 | 0.545 | 0.332 | 42.691 | 0.00 |
| KES197 | 4.118 | 246.080 | 368.6 | 0.548 | 0.329 | 42.167 | 0.00 |
| KES198 | 4.093 | 244.633 | 368.3 | 0.584 | 0.328 | 41.686 | 0.00 |
| KES199 | 4.031 | 243.655 | 364.8 | 0.531 | 0.319 | 42.156 | 0.00 |
| KES200 | 4.099 | 236.650 | 367.9 | 0.555 | 0.328 | 41.751 | 0.00 |
| KES201 | 4.081 | 248.617 | 371.2 | 0.553 | 0.334 | 42.713 | 0.00 |
| KES202 | 4.068 | 249.395 | 369.6 | 0.551 | 0.332 | 42.944 | 0.00 |
| KES203 | 4.057 | 251.603 | 367.4 | 0.544 | 0.328 | 42.516 | 0.00 |
| KES204 | 4.088 | 255.215 | 369.4 | 0.539 | 0.332 | 42.948 | 0.00 |
| KES205 | 4.035 | 246.497 | 374.0 | 0.553 | 0.357 | 42.466 | 0.00 |
| KES206 | 4.066 | 257.156 | 367.6 | 0.523 | 0.327 | 42.785 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES207 | 4.091 | 269.892 | 369.6 | 0.531 | 0.332 | 43.335 | 0.00 |
| KES208 | 4.067 | 256.671 | 372.9 | 0.521 | 0.325 | 42.875 | 0.00 |
| KES209 | 3.997 | 250.217 | 358.2 | 0.549 | 0.878 | 42.564 | 0.00 |
| KES210 | 4.150 | 259.703 | 373.6 | 0.557 | 0.298 | 43.596 | 0.00 |
| KES211 | 4.156 | 254.922 | 365.1 | 0.546 | 0.291 | 43.596 | 0.00 |
| KES212 | 4.910 | 298.365 | 367.8 | 0.560 | 0.288 | 50.732 | 0.00 |
| KES213 | 4.990 | 296.043 | 369.6 | 0.545 | 0.295 | 50.806 | 0.00 |
| KES214 | 4.787 | 312.511 | 368.0 | 0.542 | 0.301 | 51.710 | 0.00 |
| KES215 | 4.938 | 313.423 | 365.9 | 0.573 | 0.289 | 51.041 | 0.00 |
| KES216 | 4.263 | 315.079 | 365.5 | 0.559 | 0.290 | 51.502 | 0.00 |
| KES217 | 4.345 | 254.974 | 388.7 | 0.615 | 0.318 | 50.623 | 0.00 |
| KES218 | 3.067 | 97.409 | 430.7 | 0.589 | 0.115 | 28.158 | 0.00 |
| KES219 | 3.093 | 97.507 | 430.9 | 0.630 | 0.114 | 28.235 | 0.00 |
| KES220 | 4.381 | 255.769 | 394.2 | 0.598 | 0.318 | 51.612 | 0.00 |
| KES221 | 4.338 | 255.889 | 399.3 | 0.639 | 0.323 | 51.822 | 0.00 |
| KES222 | 4.291 | 235.748 | 389.1 | 0.595 | 0.310 | 51.042 | 0.00 |
| KES223 | 4.415 | 256.438 | 396.6 | 0.616 | 0.317 | 51.973 | 0.00 |
| KES224 | 4.369 | 253.169 | 391.7 | 0.600 | 0.309 | 51.849 | 0.00 |
| KES225 | 4.345 | 245.540 | 382.3 | 0.600 | 0.313 | 53.614 | 0.00 |
| KES226 | 4.458 | 256.281 | 384.0 | 0.600 | 0.305 | 52.563 | 0.00 |
| KES227 | 4.414 | 259.303 | 393.6 | 0.587 | 0.315 | 52.176 | 0.00 |
| KES228 | 4.456 | 258.200 | 384.1 | 0.586 | 0.316 | 51.716 | 0.00 |
| KES229 | 4.010 | 253.450 | 390.5 | 0.600 | 0.349 | 52.308 | 0.00 |
| KES230 | 4.394 | 249.930 | 387.0 | 0.615 | 0.357 | 52.646 | 0.00 |
| KES231 | 4.448 | 258.952 | 388.5 | 0.620 | 0.353 | 52.644 | 0.00 |
| KES232 | 4.416 | 256.863 | 382.7 | 0.614 | 0.309 | 52.312 | 0.00 |
| KES233 | 4.395 | 256.216 | 385.4 | 0.613 | 0.298 | 53.494 | 0.00 |
| KES234 | 4.441 | 262.982 | 384.0 | 0.595 | 0.312 | 52.783 | 0.00 |
| KES235 | 4.409 | 262.508 | 386.7 | 0.549 | 0.313 | 53.225 | 0.00 |
| KES236 | 4.468 | 265.119 | 396.8 | 0.585 | 0.314 | 54.793 | 0.00 |
| KES237 | 4.402 | 259.132 | 390.9 | 0.619 | 0.317 | 53.440 | 0.00 |
| KES238 | 4.374 | 253.763 | 383.6 | 0.623 | 0.295 | 52.494 | 0.00 |
| KES239 | 4.333 | 255.956 | 389.1 | 0.582 | 0.307 | 52.522 | 0.00 |
| KES240 | 4.436 | 264.217 | 391.1 | 0.574 | 0.316 | 53.217 | 0.00 |
| KES241 | 4.462 | 258.790 | 381.0 | 0.590 | 0.312 | 52.989 | 0.00 |
| KES242 | 4.389 | 266.993 | 386.0 | 0.578 | 0.312 | 53.821 | 0.00 |
| KES243 | 2.654 | 102.509 | 406.1 | 0.777 | 0.135 | 24.577 | 0.00 |
| KES244 | 3.141 | 101.880 | 411.0 | 0.751 | 0.139 | 25.264 | 0.00 |
| KES245 | 2.760 | 102.573 | 398.0 | 0.763 | 0.324 | 24.539 | 0.00 |
| KES246 | 2.751 | 105.973 | 408.3 | 0.772 | 0.271 | 24.825 | 0.00 |
| KES247 | 2.746 | 100.490 | 399.0 | 0.737 | 0.539 | 24.995 | 0.00 |
| KES248 | 2.768 | 100.079 | 399.1 | 0.723 | 0.573 | 24.530 | 0.00 |
| KES249 | 2.732 | 100.619 | 400.6 | 0.713 | 0.919 | 24.883 | 0.00 |
| KES250 | 2.504 | 91.379 | 390.3 | 0.554 | 0.349 | 22.364 | 0.00 |
| KES251 | 2.507 | 92.902 | 381.0 | 0.565 | 0.329 | 22.436 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES252 | 3.186 | 88.548 | 392.8 | 0.567 | 0.144 | 22.073 | 0.00 |
| KES253 | 2.709 | 97.916 | 405.5 | 0.584 | 0.147 | 22.937 | 0.00 |
| KES254 | 2.575 | 98.564 | 398.2 | 0.623 | 0.173 | 23.030 | 0.00 |
| KES255 | 1.792 | 91.297 | 207.7 | 0.479 | 3.029 | 15.641 | 65.00 |
| KES256 | 1.840 | 92.626 | 207.4 | 0.497 | 2.699 | 15.661 | 39.20 |
| KES257 | 1.833 | 89.352 | 211.1 | 0.481 | 3.028 | 15.952 | 35.50 |
| KES258 | 1.821 | 86.395 | 207.0 | 0.469 | 2.620 | 15.538 | 44.34 |
| KES259 | 1.841 | 92.112 | 210.0 | 0.490 | 2.773 | 15.921 | 53.79 |
| KES260 | 1.499 | 90.254 | 209.8 | 0.500 | 2.815 | 15.995 | 35.80 |
| KES261 | 1.799 | 88.812 | 201.9 | 0.456 | 2.713 | 15.442 | 26.54 |
| KES262 | 1.772 | 94.642 | 210.4 | 0.503 | 2.748 | 15.956 | 60.66 |
| KES263 | 1.862 | 89.796 | 212.5 | 0.500 | 2.785 | 16.024 | 37.68 |
| KES264 | 1.535 | 93.057 | 209.9 | 0.489 | 2.774 | 16.183 | 48.89 |
| KES265 | 1.498 | 98.733 | 212.0 | 0.478 | 2.782 | 16.330 | 50.70 |
| KES266 | 1.834 | 95.828 | 211.2 | 0.479 | 2.698 | 16.139 | 40.22 |
| KES267 | 1.382 | 95.083 | 210.2 | 0.512 | 2.691 | 16.337 | 44.44 |
| KES268 | 1.874 | 94.399 | 209.7 | 0.505 | 2.764 | 16.319 | 51.28 |
| KES269 | 1.868 | 94.089 | 207.0 | 0.482 | 2.723 | 16.273 | 50.58 |
| KES270 | 1.405 | 99.284 | 210.9 | 0.468 | 2.786 | 16.890 | 58.27 |
| KES271 | 1.814 | 94.058 | 204.8 | 0.487 | 2.860 | 16.408 | 29.49 |
| KES272 | 1.940 | 94.569 | 207.2 | 0.497 | 2.873 | 16.794 | 47.55 |
| KES273 | 1.881 | 95.360 | 209.9 | 0.494 | 2.793 | 16.906 | 60.77 |
| KES274 | 1.332 | 98.949 | 209.6 | 0.474 | 2.742 | 16.790 | 46.60 |
| KES275 | 1.868 | 97.607 | 208.0 | 0.484 | 2.744 | 16.876 | 52.91 |
| KES276 | 1.332 | 88.091 | 208.2 | 0.493 | 2.666 | 16.256 | 41.74 |
| KES277 | 1.444 | 88.641 | 211.0 | 0.518 | 2.785 | 14.803 | 74.10 |
| KES278 | 2.655 | 74.457 | 404.6 | 0.647 | 0.439 | 20.018 | 0.00 |
| KES279 | 2.696 | 87.132 | 407.6 | 0.656 | 0.147 | 20.351 | 0.00 |
| KES281 | 2.628 | 79.960 | 392.1 | 0.629 | 1.016 | 21.220 | 0.00 |
| KES282 | 2.600 | 88.403 | 381.9 | 0.624 | 0.525 | 21.238 | 0.00 |
| KES283 | 2.259 | 83.550 | 386.4 | 0.612 | 0.652 | 20.924 | 0.00 |
| KES284 | 2.495 | 87.869 | 376.6 | 0.610 | 0.830 | 19.793 | 0.00 |
| KES285 | 1.965 | 114.862 | 163.6 | 0.234 | 3.506 | 21.928 | 0.00 |
| KES286 | 1.988 | 120.124 | 167.7 | 0.237 | 3.311 | 21.557 | 0.00 |
| KES287 | 1.769 | 103.077 | 166.3 | 0.174 | 3.108 | 19.313 | 0.00 |
| KES288 | 1.821 | 109.304 | 168.6 | 0.184 | 3.112 | 20.013 | 0.00 |
| KES289 | 1.697 | 96.263 | 165.3 | 0.181 | 3.062 | 19.495 | 0.00 |
| KES290 | 1.728 | 101.205 | 167.3 | 0.210 | 3.022 | 19.525 | 0.00 |
| KES291 | 1.723 | 114.449 | 174.0 | 0.197 | 3.167 | 19.853 | 0.00 |
| KES292 | 1.610 | 107.005 | 166.7 | 0.187 | 3.042 | 19.783 | 0.00 |
| KES293 | 1.745 | 114.534 | 170.5 | 0.187 | 3.154 | 20.920 | 0.00 |
| KES294 | 1.722 | 121.986 | 170.8 | 0.179 | 3.103 | 20.303 | 0.00 |
| KES295 | 1.821 | 109.693 | 171.6 | 0.170 | 3.088 | 20.965 | 0.00 |
| KES296 | 1.717 | 114.530 | 170.3 | 0.190 | 3.124 | 20.920 | 0.00 |
| KES297 | 2.845 | 89.700 | 433.8 | 0.511 | 0.114 | 25.386 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES298 | 3.310 | 97.361 | 423.4 | 0.500 | 0.108 | 25.126 | 0.00 |
| KES299 | 3.237 | 96.358 | 430.8 | 0.510 | 0.110 | 24.904 | 0.00 |
| KES300 | 3.613 | 168.393 | 439.1 | 0.562 | 0.302 | 41.505 | 0.00 |
| KES301 | 3.466 | 173.420 | 415.8 | 0.564 | 0.302 | 38.133 | 0.00 |
| KES302 | 3.609 | 177.362 | 425.2 | 0.575 | 0.308 | 41.450 | 0.00 |
| KES303 | 3.681 | 166.257 | 422.0 | 0.597 | 0.309 | 41.356 | 0.00 |
| KES304 | 3.584 | 180.348 | 427.6 | 0.595 | 0.312 | 41.771 | 0.00 |
| KES305 | 3.666 | 179.751 | 428.3 | 0.613 | 0.310 | 42.072 | 0.00 |
| KES306 | 3.558 | 187.555 | 430.6 | 0.547 | 0.313 | 42.122 | 0.00 |
| KES307 | 3.724 | 175.975 | 420.1 | 0.585 | 0.309 | 41.583 | 0.00 |
| KES308 | 3.571 | 183.896 | 416.0 | 0.534 | 0.304 | 41.337 | 0.00 |
| KES309 | 3.705 | 201.447 | 423.2 | 0.597 | 0.292 | 42.889 | 0.00 |
| KES310 | 3.805 | 195.136 | 420.8 | 0.582 | 0.275 | 43.315 | 0.00 |
| KES311 | 4.969 | 284.968 | 443.5 | 0.652 | 0.309 | 57.985 | 0.00 |
| KES312 | 4.902 | 284.166 | 448.4 | 0.645 | 0.320 | 55.175 | 0.00 |
| KES313 | 4.883 | 280.083 | 444.0 | 0.658 | 0.311 | 55.458 | 0.00 |
| KES314 | 4.981 | 285.540 | 449.4 | 0.623 | 0.307 | 56.597 | 0.00 |
| KES315 | 5.031 | 292.226 | 452.8 | 0.623 | 0.314 | 59.536 | 0.00 |
| KES316 | 4.908 | 285.059 | 433.1 | 0.524 | 0.279 | 56.729 | 0.00 |
| KES317 | 4.855 | 285.778 | 430.7 | 0.600 | 0.294 | 56.642 | 0.00 |
| KES318 | 2.483 | 137.469 | 151.4 | 0.215 | 0.199 | 29.250 | 0.00 |
| KES319 | 2.509 | 143.484 | 148.9 | 0.221 | 0.291 | 31.340 | 0.00 |
| KES320 | 5.209 | 285.784 | 309.3 | 0.517 | 0.206 | 61.252 | 0.00 |
| KES321 | 3.843 | 206.627 | 244.6 | 0.402 | 0.183 | 42.509 | 0.00 |
| KES322 | 2.885 | 153.874 | 182.1 | 0.266 | 0.162 | 33.758 | 0.00 |
| KES323 | 3.425 | 180.905 | 215.4 | 0.290 | 0.512 | 37.884 | 0.00 |
| KES324 | 3.442 | 187.537 | 215.9 | 0.362 | 0.509 | 37.630 | 0.00 |
| KES325 | 3.325 | 178.260 | 215.6 | 0.341 | 0.504 | 36.911 | 0.00 |
| KES326 | 3.379 | 180.108 | 216.4 | 0.338 | 0.497 | 37.621 | 0.00 |
| KES327 | 3.331 | 177.210 | 214.3 | 0.335 | 0.499 | 36.366 | 0.00 |
| KES328 | 3.409 | 187.491 | 213.8 | 0.364 | 0.507 | 38.164 | 0.00 |
| KES329 | 3.384 | 185.788 | 219.6 | 0.372 | 0.519 | 37.650 | 0.00 |
| KES330 | 3.368 | 190.586 | 213.9 | 0.372 | 0.512 | 37.967 | 0.00 |
| KES331 | 3.461 | 189.078 | 214.0 | 0.367 | 0.503 | 38.204 | 0.00 |
| KES335 | 3.485 | 185.803 | 213.9 | 0.374 | 0.509 | 38.227 | 0.00 |
| KES336 | 3.464 | 185.512 | 212.0 | 0.332 | 0.506 | 37.410 | 0.00 |
| KES337 | 3.473 | 186.205 | 210.6 | 0.341 | 0.504 | 37.745 | 0.00 |
| KES338 | 3.442 | 187.416 | 211.9 | 0.314 | 0.500 | 38.175 | 0.00 |
| KES339 | 3.481 | 186.391 | 219.3 | 0.362 | 0.502 | 38.827 | 0.00 |
| KES340 | 3.522 | 197.791 | 220.1 | 0.393 | 0.509 | 39.459 | 0.00 |
| KES341 | 2.437 | 143.701 | 145.6 | 0.225 | 0.304 | 29.981 | 0.00 |
| KES342 | 3.296 | 171.939 | 245.6 | 0.358 | 0.613 | 36.523 | 0.00 |
| KES343 | 2.405 | 139.916 | 147.3 | 0.212 | 0.295 | 28.904 | 0.00 |
| KES344 | 2.483 | 149.786 | 151.5 | 0.215 | 0.303 | 30.246 | 0.00 |
| KES345 | 1.685 | 92.021 | 94.5 | 0.124 | 2.270 | 19.083 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES346 | 1.694 | 88.938 | 97.0 | 0.128 | 2.263 | 18.473 | 0.00 |
| KES347 | 3.855 | 219.002 | 333.1 | 0.474 | 0.359 | 43.769 | 0.00 |
| KES348 | 3.794 | 216.110 | 338.7 | 0.468 | 0.353 | 43.810 | 0.00 |
| KES349 | 3.795 | 223.683 | 340.7 | 0.490 | 0.360 | 44.698 | 0.00 |
| KES350 | 5.560 | 302.912 | 442.4 | 0.727 | 0.160 | 65.915 | 0.00 |
| KES351 | 5.447 | 310.168 | 440.5 | 0.738 | 0.170 | 64.734 | 0.00 |
| KES352 | 5.507 | 322.429 | 440.1 | 0.740 | 0.165 | 65.755 | 0.00 |
| KES353 | 5.355 | 316.022 | 438.1 | 0.748 | 0.164 | 64.867 | 0.00 |
| KES354 | 5.410 | 310.033 | 442.8 | 0.717 | 0.165 | 64.128 | 0.00 |
| KES355 | 5.345 | 312.773 | 439.2 | 0.785 | 0.166 | 64.943 | 0.00 |
| KES356 | 5.459 | 318.434 | 440.4 | 0.778 | 0.162 | 64.812 | 0.00 |
| KES357 | 5.517 | 319.521 | 441.9 | 0.745 | 0.171 | 65.108 | 0.00 |
| KES358 | 5.513 | 321.352 | 439.4 | 0.704 | 0.163 | 65.665 | 0.00 |
| KES359 | 5.296 | 322.736 | 428.5 | 0.756 | 0.163 | 64.760 | 0.00 |
| KES360 | 5.420 | 316.916 | 434.6 | 0.758 | 0.162 | 64.416 | 0.00 |
| KES361 | 5.426 | 327.164 | 435.9 | 0.711 | 0.172 | 64.934 | 0.00 |
| KES362 | 5.358 | 319.989 | 437.5 | 0.709 | 0.165 | 64.079 | 0.00 |
| KES363 | 5.401 | 319.447 | 441.1 | 0.752 | 0.164 | 64.930 | 0.00 |
| KES364 | 5.124 | 326.442 | 441.6 | 0.689 | 0.326 | 62.824 | 0.00 |
| KES365 | 5.085 | 309.548 | 425.7 | 0.668 | 0.316 | 61.718 | 0.00 |
| KES366 | 5.162 | 319.687 | 437.2 | 0.647 | 0.319 | 62.630 | 0.00 |
| KES367 | 5.115 | 316.883 | 432.9 | 0.677 | 0.322 | 61.102 | 0.00 |
| KES368 | 5.091 | 320.597 | 431.1 | 0.723 | 0.319 | 61.989 | 0.00 |
| KES369 | 5.286 | 322.970 | 442.7 | 0.682 | 0.320 | 63.606 | 0.00 |
| KES370 | 1.801 | 186.627 | 213.4 | 0.499 | 0.994 | 35.525 | 0.00 |
| KES371 | 1.820 | 191.513 | 212.5 | 0.515 | 0.991 | 35.639 | 0.00 |
| KES372 | 1.763 | 178.681 | 212.9 | 0.533 | 0.981 | 35.463 | 0.00 |
| KES373 | 1.761 | 185.912 | 210.6 | 0.512 | 0.970 | 35.101 | 0.00 |
| KES374 | 2.244 | 94.694 | 344.4 | 0.575 | 0.126 | 21.397 | 0.00 |
| KES375 | 2.268 | 94.672 | 345.2 | 0.600 | 0.124 | 21.715 | 0.00 |
| KES376 | 2.820 | 105.125 | 409.3 | 0.698 | 0.138 | 25.337 | 0.00 |
| KES377 | 2.747 | 105.465 | 411.7 | 0.703 | 0.138 | 24.851 | 0.00 |
| KES378 | 2.716 | 108.560 | 403.1 | 0.701 | 0.139 | 24.646 | 0.00 |
| KES379 | 3.322 | 105.717 | 399.5 | 0.703 | 0.129 | 25.717 | 0.00 |
| KES380 | 2.795 | 99.418 | 390.0 | 0.695 | 0.127 | 25.058 | 0.00 |
| KES381 | 3.297 | 97.534 | 394.2 | 0.698 | 0.129 | 24.893 | 0.00 |
| KES382 | 2.845 | 90.609 | 391.3 | 0.670 | 0.129 | 25.096 | 0.00 |
| KES383 | 3.189 | 98.524 | 395.8 | 0.688 | 0.125 | 25.665 | 0.00 |
| KES384 | 2.831 | 95.440 | 394.8 | 0.713 | 0.124 | 25.410 | 0.00 |
| KES385 | 3.268 | 96.210 | 395.3 | 0.673 | 0.129 | 25.351 | 0.00 |
| KES386 | 3.135 | 92.832 | 390.0 | 0.690 | 0.126 | 24.737 | 0.00 |
| KES387 | 3.174 | 100.635 | 388.4 | 0.671 | 0.127 | 25.617 | 0.00 |
| KES388 | 2.881 | 96.269 | 388.4 | 0.671 | 0.126 | 24.841 | 0.00 |
| KES389 | 2.875 | 97.477 | 391.5 | 0.698 | 0.130 | 25.666 | 0.00 |
| KES390 | 3.184 | 93.108 | 386.1 | 0.669 | 0.124 | 25.444 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES391 | 3.226 | 99.288 | 391.7 | 0.691 | 0.132 | 25.407 | 0.00 |
| KES392 | 2.940 | 100.274 | 396.6 | 0.695 | 0.134 | 25.984 | 0.00 |
| KES393 | 3.207 | 99.902 | 396.6 | 0.695 | 0.126 | 25.526 | 0.00 |
| KES394 | 3.403 | 109.797 | 382.3 | 0.644 | 0.118 | 26.390 | 0.00 |
| KES403 | 1.795 | 176.007 | 208.31 | 0.5262 | 0.9721 | 34.469 | 0.00 |
| KES404 | 1.803 | 188.940 | 213.59 | 0.5179 | 0.9762 | 34.746 | 0.00 |
| KES405 | 1.781 | 179.677 | 218.57 | 0.4901 | 0.9927 | 34.980 | 0.00 |
| KES406 | 1.786 | 187.735 | 213.92 | 0.5059 | 0.9798 | 35.230 | 0.00 |
| KES407 | 1.766 | 175.343 | 210.03 | 0.4719 | 0.9915 | 34.882 | 0.00 |
| KES408 | 3.464 | 223.852 | 254.31 | 0.5514 | 1.5248 | 39.987 | 0.00 |
| KES409 | 3.469 | 226.303 | 249.94 | 0.5781 | 1.4936 | 39.563 | 0.00 |
| KES410 | 3.440 | 225.216 | 253.69 | 0.5643 | 1.5081 | 39.581 | 0.00 |
| KES411 | 2.435 | 151.446 | 199.09 | 0.2912 | 0.7535 | 27.716 | 0.00 |
| KES412 | 3.051 | 207.550 | 344.03 | 0.6206 | 0.2285 | 35.520 | 0.00 |
| KES413 | 3.125 | 215.167 | 348.85 | 0.5836 | 0.2288 | 35.776 | 0.00 |
| KES414 | 2.020 | 133.650 | 172.24 | 0.2067 | 3.0475 | 23.985 | 0.00 |
| KES415 | 2.014 | 130.196 | 171.21 | 0.2031 | 3.0518 | 23.667 | 0.00 |
| KES416 | 2.027 | 134.836 | 165.87 | 0.1939 | 3.0248 | 23.399 | 0.00 |
| KES417 | 1.991 | 133.668 | 174.81 | 0.2476 | 3.1078 | 23.741 | 0.00 |
| KES418 | 1.915 | 94.594 | 133.55 | 0.2111 | 0.4367 | 20.578 | 0.00 |
| KES419 | 1.643 | 106.789 | 166.42 | 0.1951 | 4.4994 | 19.560 | 0.00 |
| KES420 | 1.623 | 104.998 | 160.79 | 0.1627 | 4.3549 | 19.096 | 0.00 |
| KES421 | 1.648 | 112.780 | 162.87 | 0.1578 | 4.3886 | 19.091 | 0.00 |
| KES422 | 1.670 | 119.333 | 166.02 | 0.2015 | 4.5428 | 19.692 | 0.00 |
| KES423 | 1.665 | 108.320 | 164.94 | 0.1868 | 4.5128 | 19.512 | 0.00 |
| KES424 | 2.357 | 101.892 | 163.32 | 0.2294 | 1.4048 | 22.062 | 0.00 |
| KES425 | 2.123 | 95.057 | 144.73 | 0.2094 | 2.3727 | 20.022 | 0.00 |
| KES426 | 1.948 | 81.093 | 128.24 | 0.2050 | 5.1305 | 17.933 | 61.52 |
| KES427 | 2.368 | 97.767 | 160.34 | 0.2235 | 1.4820 | 21.686 | 0.00 |
| KES428 | 2.285 | 96.211 | 155.94 | 0.2096 | 3.3454 | 21.309 | 0.00 |
| KES429 | 1.673 | 107.521 | 169.50 | 0.1796 | 4.3808 | 19.366 | 0.00 |
| KES430 | 0.870 | 68.535 | 113.55 | 0.2356 | 1.0984 | 13.868 | 0.00 |
| KES431 | 1.351 | 71.122 | 118.15 | 0.1261 | 0.2684 | 15.303 | 0.00 |
| KES432 | 1.376 | 69.742 | 121.73 | 0.1205 | 0.2758 | 15.486 | 0.00 |
| KES433 | 1.473 | 76.224 | 129.19 | 0.1748 | 0.1324 | 17.198 | 0.00 |
| KES434 | 1.345 | 78.281 | 121.06 | 0.1419 | 0.2877 | 15.484 | 0.00 |
| KES435 | 1.348 | 76.653 | 122.60 | 0.1409 | 0.3068 | 15.365 | 0.00 |
| KES436 | 1.750 | 105.851 | 188.91 | 0.2196 | 1.2077 | 21.438 | 0.00 |
| KES437 | 1.186 | 84.386 | 129.94 | 0.1073 | 7.3702 | 14.671 | 0.00 |
| KES438 | 1.620 | 110.723 | 161.42 | 0.1880 | 4.3320 | 18.761 | 0.00 |
| KES439 | 2.561 | 120.531 | 159.58 | 0.2855 | 0.4937 | 25.760 | 0.00 |
| KES440 | 2.573 | 119.342 | 161.11 | 0.2669 | 0.4571 | 26.181 | 0.00 |
| KES441 | 2.600 | 117.356 | 163.48 | 0.2483 | 0.4589 | 26.176 | 0.00 |
| KES442 | 2.052 | 112.189 | 152.56 | 0.1261 | 1.8848 | 22.796 | 0.00 |
| KES443 | 2.041 | 115.746 | 145.82 | 0.1352 | 1.8687 | 22.882 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES444 | 3.541 | 239.795 | 277.15 | 0.7374 | 0.8876 | 45.997 | 0.00 |
| KES445 | 3.532 | 226.667 | 275.08 | 0.6771 | 0.8559 | 45.213 | 0.00 |
| KES446 | 5.214 | 367.820 | 386.66 | 1.1902 | 0.3903 | 71.124 | 0.00 |
| KES447 | 2.369 | 132.945 | 159.82 | 0.3725 | 0.7911 | 25.489 | 0.00 |
| KES448 | 2.360 | 134.254 | 171.39 | 0.3602 | 0.7811 | 25.495 | 0.00 |
| KES449 | 1.842 | 107.069 | 185.45 | 0.4660 | 0.3842 | 21.859 | 0.00 |
| KES450 | 1.843 | 105.352 | 182.73 | 0.4625 | 0.3740 | 21.693 | 0.00 |
| KES451 | 1.842 | 118.070 | 193.27 | 0.4435 | 0.3793 | 21.706 | 0.00 |
| KES452 | 1.850 | 115.958 | 189.70 | 0.5128 | 0.3935 | 21.957 | 0.00 |
| KES453 | 1.988 | 105.003 | 190.17 | 0.5226 | 0.3869 | 22.186 | 0.00 |
| KES454 | 2.118 | 115.444 | 188.93 | 0.5454 | 0.3772 | 22.518 | 0.00 |
| KES455 | 1.966 | 115.490 | 184.91 | 0.5367 | 0.3819 | 22.439 | 0.00 |
| KES456 | 2.111 | 109.453 | 189.16 | 0.5065 | 0.3879 | 22.638 | 0.00 |
| KES457 | 1.986 | 102.117 | 171.35 | 0.5353 | 0.8721 | 22.451 | 0.00 |
| KES458 | 1.941 | 111.929 | 172.67 | 0.5117 | 0.8750 | 22.486 | 0.00 |
| KES459 | 2.165 | 96.398 | 229.93 | 0.2278 | 2.6126 | 16.621 | 0.00 |
| KES460 | 1.925 | 87.710 | 210.24 | 0.1917 | 2.9672 | 15.651 | 0.00 |
| KES461 | 1.749 | 112.460 | 163.99 | 0.1838 | 4.5145 | 19.544 | 0.00 |
| KES462 | 2.128 | 85.232 | 223.20 | 0.2078 | 2.7586 | 15.958 | 0.00 |
| KES463 | 2.128 | 83.504 | 190.54 | 0.2017 | 3.0377 | 17.997 | 0.00 |
| KES464 | 2.166 | 103.611 | 190.10 | 0.1783 | 3.2413 | 17.998 | 0.00 |
| KES465 | 2.169 | 91.173 | 187.82 | 0.1820 | 3.1138 | 17.292 | 0.00 |
| KES466 | 2.279 | 104.443 | 198.01 | 0.2015 | 3.1599 | 18.263 | 0.00 |
| KES467 | 5.651 | 314.381 | 451.37 | 0.7141 | 0.2999 | 63.208 | 0.00 |
| KES468 | 5.792 | 305.104 | 462.35 | 0.7549 | 0.3018 | 64.416 | 0.00 |
| KES469 | 5.653 | 298.345 | 454.85 | 0.7048 | 0.2995 | 63.255 | 0.00 |
| KES470 | 5.339 | 297.046 | 453.09 | 0.7434 | 0.2897 | 62.739 | 0.00 |
| KES471 | 3.666 | 219.673 | 209.13 | 0.3136 | 0.3479 | 41.281 | 0.00 |
| KES472 | 3.763 | 201.137 | 212.53 | 0.3275 | 0.3445 | 41.222 | 0.00 |
| KES473 | 3.734 | 193.947 | 208.15 | 0.3277 | 0.3342 | 40.338 | 0.00 |
| KES474 | 3.730 | 211.951 | 212.58 | 0.3232 | 0.3425 | 41.558 | 0.00 |
| KES475 | 5.853 | 309.894 | 469.57 | 0.7221 | 0.3000 | 64.881 | 0.00 |
| KES476 | 5.788 | 323.092 | 460.23 | 0.7160 | 0.2893 | 63.560 | 0.00 |
| KES477 | 3.754 | 202.467 | 208.90 | 0.3669 | 0.3390 | 40.612 | 0.00 |
| KES478 | 3.624 | 200.179 | 209.79 | 0.3170 | 0.3395 | 40.901 | 0.00 |
| KES479 | 4.959 | 268.175 | 396.56 | 0.6294 | 0.2864 | 53.480 | 0.00 |
| KES480 | 3.830 | 178.437 | 216.08 | 0.3082 | 0.1521 | 37.968 | 0.00 |
| KES481 | 3.471 | 179.312 | 244.26 | 0.4034 | 0.6032 | 36.250 | 0.00 |
| KES482 | 3.392 | 177.358 | 247.53 | 0.4385 | 0.6172 | 37.012 | 0.00 |
| KES483 | 5.574 | 300.629 | 430.12 | 0.6960 | 0.3145 | 59.973 | 0.00 |
| KES484 | 5.704 | 304.420 | 439.88 | 0.6883 | 0.3163 | 60.776 | 0.00 |
| KES485 | 4.770 | 283.569 | 389.90 | 0.6145 | 0.3063 | 53.052 | 0.00 |
| KES486 | 4.778 | 272.750 | 389.49 | 0.6021 | 0.3062 | 52.497 | 0.00 |
| KES487 | 4.648 | 259.121 | 371.80 | 0.5417 | 0.3490 | 49.474 | 0.00 |
| KES488 | 4.481 | 262.385 | 364.42 | 0.5505 | 0.3815 | 48.973 | 0.00 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES489 | 4.620 | 274.009 | 373.96 | 0.5780 | 0.3289 | 51.000 | 0.00 |
| KES490 | 4.635 | 281.406 | 379.70 | 0.5464 | 0.3336 | 52.905 | 0.00 |
| KES491 | 4.509 | 276.543 | 374.11 | 0.5766 | 0.3274 | 51.351 | 0.00 |
| KES492 | 4.713 | 277.243 | 373.93 | 0.6040 | 0.3267 | 51.511 | 0.00 |
| KES493 | 4.697 | 285.091 | 369.96 | 0.5499 | 0.3328 | 50.714 | 0.00 |
| KES494 | 4.600 | 266.563 | 374.86 | 0.5633 | 0.3262 | 50.957 | 0.00 |
| KES495 | 4.796 | 258.910 | 374.51 | 0.5717 | 0.3541 | 50.856 | 0.00 |
| KES496 | 2.625 | 162.965 | 252.22 | 0.5714 | 0.3393 | 31.782 | 0.00 |
| KES497 | 2.582 | 178.463 | 262.32 | 0.5978 | 0.3466 | 33.123 | 0.00 |
| KES498 | 2.621 | 173.003 | 259.92 | 0.5812 | 0.3298 | 33.102 | 0.00 |
| KES499 | 2.490 | 157.588 | 252.32 | 0.5532 | 0.3189 | 31.663 | 0.00 |
| KES500 | 2.501 | 164.918 | 249.78 | 0.5650 | 0.4938 | 32.158 | 0.00 |
| KES501 | 2.569 | 160.446 | 261.48 | 0.5808 | 0.3727 | 31.686 | 0.00 |
| KES502 | 2.345 | 145.517 | 248.55 | 0.5814 | 0.3432 | 29.218 | 0.00 |
| KES503 | 1.812 | 99.901 | 200.92 | 0.5046 | 2.3831 | 16.889 | 0.00 |
| KES504 | 1.815 | 93.540 | 201.46 | 0.5060 | 2.3792 | 16.876 | 0.00 |
| KES505 | 0.993 | 102.620 | 203.29 | 0.5044 | 2.4021 | 17.077 | 0.00 |
| KES506 | 1.877 | 103.297 | 203.18 | 0.5285 | 2.4069 | 17.104 | 0.00 |
| KES507 | 1.042 | 107.107 | 203.09 | 0.5244 | 2.4207 | 17.126 | 0.00 |
| KES508 | 0.996 | 102.824 | 202.74 | 0.5321 | 2.4392 | 17.191 | 0.00 |
| KES509 | 1.934 | 104.194 | 151.36 | 0.3130 | 5.4101 | 19.966 | 0.00 |
| KES510 | 2.004 | 107.083 | 155.85 | 0.3399 | 4.8199 | 20.245 | 0.00 |
| KES511 | 1.926 | 103.645 | 150.61 | 0.3106 | 4.8050 | 19.675 | 0.00 |
| KES512 | 1.906 | 103.987 | 149.12 | 0.2950 | 4.8226 | 19.499 | 0.00 |
| KES513 | 1.871 | 99.802 | 129.30 | 0.3021 | 4.5025 | 18.809 | 65.31 |
| KES514 | 1.551 | 99.470 | 148.46 | 0.4862 | 3.5768 | 17.803 | 0.00 |
| KES515 | 1.574 | 101.171 | 148.32 | 0.4573 | 3.6912 | 18.241 | 0.00 |
| KES516 | 1.522 | 64.439 | 133.83 | 0.3077 | 7.1696 | 13.898 | 77.07 |
| KES517 | 1.545 | 62.818 | 136.99 | 0.3405 | 7.1684 | 14.038 | 79.46 |
| KES518 | 1.434 | 63.291 | 135.65 | 0.3319 | 6.8138 | 14.157 | 58.29 |
| KES519 | 1.367 | 62.418 | 130.01 | 0.3088 | 7.8086 | 13.697 | 80.00 |
| KES520 | 1.520 | 65.226 | 131.65 | 0.3371 | 7.3366 | 13.910 | 79.28 |
| KES521 | 1.451 | 61.777 | 129.58 | 0.3231 | 10.7172 | 13.493 | 42.61 |
| KES522 | 1.495 | 63.935 | 133.15 | 0.3127 | 6.6201 | 13.783 | 83.09 |
| KES523 | 1.532 | 65.194 | 136.50 | 0.3423 | 6.7818 | 14.027 | 62.90 |
| KES524 | 1.511 | 65.073 | 133.57 | 0.3749 | 7.1648 | 13.993 | 67.31 |
| KES525 | 1.517 | 65.275 | 135.47 | 0.3255 | 7.2273 | 13.995 | 44.84 |
| KES526 | 1.385 | 60.774 | 112.99 | 0.2858 | 11.2641 | 12.883 | 142.41 |
| KES527 | 1.190 | 81.448 | 105.35 | 0.1762 | 6.3291 | 14.308 | 0.00 |
| KES528 | 1.161 | 79.887 | 104.58 | 0.1902 | 6.1841 | 14.509 | 0.00 |
| KES529 | 0.952 | 79.716 | 102.05 | 0.1885 | 6.1018 | 14.382 | 0.00 |
| KES530 | 0.897 | 76.548 | 97.46 | 0.2080 | 6.0715 | 13.824 | 208.07 |
| KES531 | 0.733 | 75.239 | 95.09 | 0.1907 | 5.8173 | 13.534 | 214.12 |
| KES532 | 1.134 | 74.821 | 97.11 | 0.1960 | 6.0223 | 13.857 | 103.12 |
| KES533 | 0.903 | 76.832 | 95.31 | 0.1965 | 5.9447 | 13.754 | 221.10 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Lu | Nd | Rb | Sb | Sc | Sm | Sr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES534 | 3.505 | 112.457 | 410.16 | 0.6969 | 0.1268 | 27.372 | 0.00 |
| KES535 | 3.437 | 112.443 | 403.02 | 0.6671 | 0.1279 | 26.902 | 0.00 |
| KES536 | 2.658 | 190.362 | 215.16 | 0.5344 | 0.9950 | 35.912 | 0.00 |
| KES537 | 2.594 | 190.703 | 213.23 | 0.5024 | 0.9760 | 35.939 | 0.00 |
| KES538 | 1.775 | 190.368 | 210.89 | 0.4914 | 0.9764 | 35.575 | 0.00 |
| KES539 | 2.715 | 191.088 | 213.16 | 0.5327 | 0.9820 | 36.030 | 0.00 |
| KES540 | 2.660 | 161.601 | 196.88 | 0.2738 | 0.7513 | 28.454 | 0.00 |
| KES541 | 2.132 | 93.352 | 140.36 | 0.1770 | 3.4080 | 19.747 | 76.67 |
| KES542 | 2.350 | 105.595 | 155.29 | 0.1943 | 1.9454 | 21.747 | 0.00 |

Appendix 6 - Concentration data from NAA for obsidian from Kenya. All concentrations are in parts per million ( ppm )

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES001 | 14.926 | 5.090 | 40.191 | 12.762 | 16.972 | 382.90 | 1419.30 |
| KES002 | 14.915 | 5.140 | 40.234 | 13.285 | 16.947 | 401.60 | 1484.90 |
| KES003 | 12.407 | 2.319 | 39.646 | 14.015 | 10.069 | 158.00 | 448.30 |
| KES004 | 12.316 | 2.243 | 39.426 | 13.433 | 10.078 | 146.40 | 468.20 |
| KES005 | 30.822 | 9.786 | 82.210 | 26.248 | 38.913 | 628.20 | 2957.40 |
| KES006 | 35.438 | 11.377 | 87.385 | 30.851 | 44.804 | 735.90 | 2991.90 |
| KES007 | 33.036 | 10.400 | 87.663 | 29.851 | 41.136 | 650.00 | 3174.70 |
| KES008 | 19.665 | 7.225 | 34.565 | 13.430 | 26.960 | 541.90 | 1510.70 |
| KES009 | 19.451 | 7.187 | 34.138 | 13.859 | 26.634 | 525.50 | 1556.70 |
| KES010 | 35.575 | 11.187 | 94.493 | 32.440 | 44.829 | 681.40 | 3455.30 |
| KES011 | 34.491 | 10.975 | 92.805 | 31.289 | 40.148 | 697.80 | 3339.90 |
| KES012 | 26.505 | 8.520 | 69.966 | 24.501 | 32.112 | 560.40 | 2554.20 |
| KES013 | 20.618 | 6.988 | 38.089 | 16.426 | 26.139 | 468.00 | 1568.20 |
| KES014 | 18.739 | 4.817 | 36.695 | 16.490 | 12.704 | 464.30 | 1605.20 |
| KES015 | 19.445 | 5.060 | 38.425 | 15.744 | 13.781 | 497.90 | 1665.20 |
| KES016 | 16.072 | 2.387 | 37.012 | 14.892 | 9.496 | 162.50 | 1075.50 |
| KES017 | 16.281 | 2.407 | 37.279 | 14.943 | 9.401 | 164.70 | 1055.40 |
| KES018 | 16.206 | 2.438 | 37.099 | 14.878 | 9.313 | 171.10 | 1097.30 |
| KES019 | 15.909 | 2.333 | 36.683 | 14.886 | 9.432 | 163.80 | 1043.60 |
| KES020 | 18.910 | 2.728 | 32.667 | 13.477 | 10.890 | 264.90 | 903.80 |
| KES021 | 22.316 | 4.807 | 66.342 | 24.285 | 19.346 | 320.30 | 1460.00 |
| KES022 | 21.295 | 4.638 | 63.056 | 22.966 | 18.420 | 305.60 | 1372.20 |
| KES023 | 22.328 | 4.792 | 66.741 | 24.311 | 19.504 | 324.40 | 1477.80 |
| KES024 | 22.104 | 4.778 | 65.917 | 23.616 | 19.735 | 321.10 | 1433.30 |
| KES025 | 23.334 | 4.979 | 69.396 | 25.310 | 20.089 | 332.30 | 1518.70 |
| KES026 | 22.925 | 4.875 | 68.095 | 24.632 | 19.836 | 328.10 | 1491.70 |
| KES027 | 23.040 | 4.934 | 68.607 | 25.626 | 20.212 | 331.40 | 1489.80 |
| KES028 | 23.130 | 4.936 | 68.812 | 25.695 | 20.384 | 324.90 | 1539.80 |
| KES029 | 23.191 | 4.915 | 68.257 | 25.646 | 19.910 | 317.20 | 1493.10 |
| KES030 | 21.969 | 4.655 | 65.692 | 23.649 | 18.656 | 271.60 | 1394.50 |
| KES031 | 21.933 | 4.681 | 65.809 | 23.161 | 18.878 | 274.20 | 1403.10 |
| KES032 | 22.008 | 4.648 | 65.941 | 23.262 | 18.619 | 277.90 | 1410.30 |
| KES033 | 19.634 | 6.686 | 36.593 | 14.680 | 24.798 | 401.60 | 1437.70 |
| KES034 | 19.793 | 6.689 | 36.956 | 15.402 | 24.968 | 400.40 | 1504.60 |
| KES035 | 19.782 | 6.701 | 37.158 | 14.744 | 24.934 | 386.30 | 1537.60 |
| KES036 | 19.717 | 6.678 | 36.557 | 15.683 | 24.666 | 393.30 | 1519.40 |
| KES037 | 19.681 | 6.734 | 36.904 | 14.839 | 24.773 | 394.80 | 1529.90 |
| KES038 | 19.772 | 6.694 | 36.929 | 15.182 | 25.242 | 390.90 | 1527.60 |
| KES039 | 19.560 | 6.636 | 36.509 | 15.817 | 24.706 | 394.90 | 1495.30 |
| KES040 | 19.639 | 6.692 | 36.396 | 16.233 | 24.560 | 400.80 | 1468.60 |
| KES042 | 22.467 | 6.380 | 40.483 | 18.357 | 25.956 | 418.30 | 1472.60 |
| KES043 | 18.839 | 7.063 | 33.145 | 15.191 | 25.168 | 442.50 | 1432.70 |
| KES044 | 18.856 | 6.969 | 32.889 | 14.576 | 24.971 | 445.30 | 1411.00 |
| KES045 | 19.440 | 7.071 | 34.337 | 15.471 | 25.633 | 449.30 | 1480.20 |
| KES046 | 19.517 | 7.075 | 34.421 | 15.483 | 25.978 | 451.50 | 1534.70 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES047 | 20.666 | 4.425 | 61.594 | 24.356 | 18.124 | 256.80 | 1346.80 |
| KES048 | 22.019 | 4.703 | 65.677 | 25.170 | 19.042 | 270.50 | 1418.40 |
| KES049 | 21.949 | 4.632 | 64.098 | 26.140 | 18.820 | 271.20 | 1410.50 |
| KES050 | 21.900 | 5.060 | 66.004 | 25.785 | 20.406 | 289.70 | 1537.10 |
| KES051 | 21.950 | 5.068 | 66.442 | 26.113 | 20.365 | 290.80 | 1532.50 |
| KES052 | 27.975 | 6.077 | 82.976 | 32.601 | 24.583 | 348.40 | 1866.40 |
| KES053 | 27.814 | 6.081 | 82.930 | 33.350 | 24.518 | 344.80 | 1834.60 |
| KES054 | 18.399 | 6.753 | 33.220 | 15.830 | 26.300 | 433.90 | 1444.00 |
| KES055 | 18.488 | 6.815 | 33.293 | 16.457 | 26.166 | 440.80 | 1421.10 |
| KES056 | 18.583 | 6.855 | 33.540 | 16.520 | 26.634 | 440.40 | 1398.00 |
| KES057 | 13.906 | 3.957 | 28.606 | 13.237 | 15.708 | 303.50 | 1011.50 |
| KES058 | 28.047 | 8.819 | 75.028 | 32.710 | 31.974 | 488.70 | 2654.40 |
| KES059 | 34.182 | 10.871 | 83.827 | 35.483 | 39.934 | 612.10 | 2795.50 |
| KES060 | 22.914 | 7.129 | 59.279 | 25.107 | 26.727 | 411.80 | 2041.70 |
| KES061 | 22.953 | 7.251 | 59.575 | 26.028 | 26.705 | 412.70 | 2069.90 |
| KES062 | 25.482 | 8.121 | 67.237 | 28.820 | 29.476 | 466.50 | 2399.10 |
| KES063 | 25.421 | 7.847 | 66.410 | 21.845 | 30.465 | 404.80 | 2362.40 |
| KES064 | 25.437 | 7.931 | 66.633 | 22.303 | 30.453 | 408.00 | 2341.00 |
| KES065 | 27.221 | 8.476 | 72.219 | 23.913 | 33.175 | 421.20 | 2525.30 |
| KES066 | 13.624 | 2.968 | 24.730 | 9.374 | 11.933 | 224.20 | 900.90 |
| KES067 | 25.747 | 7.886 | 67.987 | 23.329 | 31.132 | 402.60 | 2386.70 |
| KES068 | 25.800 | 8.026 | 68.525 | 22.221 | 31.042 | 400.60 | 2401.10 |
| KES069 | 25.649 | 7.958 | 68.195 | 24.067 | 31.198 | 401.70 | 2392.10 |
| KES070 | 34.421 | 10.767 | 83.856 | 29.414 | 41.054 | 528.30 | 2764.50 |
| KES071 | 34.439 | 10.691 | 83.788 | 29.586 | 41.323 | 525.20 | 2806.30 |
| KES072 | 34.304 | 10.625 | 83.704 | 29.082 | 40.697 | 528.60 | 2785.30 |
| KES073 | 34.590 | 10.729 | 84.164 | 28.891 | 40.597 | 534.50 | 2796.20 |
| KES074 | 34.638 | 10.721 | 84.165 | 30.210 | 40.873 | 530.80 | 2806.60 |
| KES075 | 21.304 | 5.546 | 65.910 | 21.705 | 23.117 | 279.90 | 1752.10 |
| KES076 | 12.344 | 2.255 | 38.831 | 13.646 | 10.885 | 106.60 | 442.50 |
| KES077 | 12.591 | 2.234 | 39.062 | 14.980 | 10.883 | 114.20 | 450.90 |
| KES078 | 12.398 | 2.273 | 39.075 | 13.804 | 9.680 | 114.10 | 448.60 |
| KES079 | 29.461 | 9.076 | 77.981 | 28.255 | 36.273 | 448.50 | 2736.50 |
| KES080 | 8.377 | 3.166 | 12.426 | 6.271 | 12.197 | 228.70 | 627.20 |
| KES081 | 19.904 | 6.657 | 36.713 | 17.251 | 25.405 | 348.30 | 1481.80 |
| KES082 | 20.022 | 6.676 | 37.181 | 15.921 | 25.961 | 351.10 | 1512.70 |
| KES083 | 20.146 | 6.796 | 37.093 | 14.976 | 25.565 | 352.90 | 1477.90 |
| KES084 | 20.096 | 6.712 | 37.218 | 15.899 | 25.665 | 354.80 | 1504.60 |
| KES085 | 20.067 | 6.701 | 36.931 | 16.605 | 25.657 | 352.30 | 1502.60 |
| KES086 | 20.065 | 6.698 | 36.751 | 16.001 | 25.680 | 347.50 | 1520.60 |
| KES087 | 32.734 | 10.063 | 86.999 | 32.723 | 35.643 | 488.60 | 3048.10 |
| KES088 | 32.359 | 9.890 | 85.100 | 31.887 | 35.437 | 478.60 | 2975.20 |
| KES089 | 12.562 | 4.687 | 20.138 | 10.303 | 16.517 | 310.60 | 987.60 |
| KES090 | 12.895 | 4.736 | 20.659 | 10.256 | 16.627 | 317.20 | 1011.00 |
| KES091 | 26.401 | 8.232 | 69.521 | 24.892 | 30.076 | 424.60 | 2491.70 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES092 | 26.446 | 8.274 | 69.727 | 25.945 | 30.087 | 423.70 | 2468.20 |
| KES093 | 26.009 | 8.123 | 67.967 | 27.373 | 30.122 | 411.70 | 2435.10 |
| KES094 | 25.445 | 7.976 | 66.462 | 26.658 | 30.092 | 402.90 | 2351.90 |
| KES095 | 25.750 | 7.962 | 68.861 | 22.423 | 30.132 | 476.51 | 2451.23 |
| KES096 | 26.554 | 8.185 | 71.813 | 23.540 | 31.111 | 490.57 | 2517.05 |
| KES097 | 25.922 | 8.079 | 70.148 | 22.010 | 29.884 | 478.04 | 2461.78 |
| KES098 | 18.842 | 4.688 | 37.360 | 16.683 | 12.283 | 428.24 | 1530.81 |
| KES099 | 18.769 | 4.678 | 37.358 | 15.809 | 13.347 | 430.37 | 1535.13 |
| KES100 | 18.442 | 4.567 | 36.399 | 15.516 | 12.142 | 415.25 | 1552.62 |
| KES101 | 13.448 | 2.907 | 24.825 | 10.369 | 11.446 | 215.81 | 898.28 |
| KES102 | 13.815 | 2.926 | 25.433 | 10.565 | 11.924 | 224.43 | 864.86 |
| KES103 | 13.409 | 2.956 | 24.716 | 10.910 | 11.250 | 215.61 | 859.52 |
| KES104 | 13.042 | 2.727 | 24.088 | 11.148 | 11.078 | 254.25 | 846.44 |
| KES124 | 12.589 | 2.750 | 23.084 | 10.875 | 10.854 | 215.04 | 804.66 |
| KES125 | 12.210 | 2.727 | 22.815 | 10.483 | 11.033 | 207.71 | 829.71 |
| KES126 | 13.893 | 3.003 | 25.441 | 11.404 | 12.060 | 235.36 | 925.44 |
| KES127 | 13.570 | 2.873 | 24.713 | 12.083 | 11.272 | 235.60 | 884.79 |
| KES128 | 12.686 | 2.729 | 23.100 | 11.227 | 11.100 | 222.70 | 820.36 |
| KES130 | 13.606 | 2.983 | 24.735 | 12.074 | 12.141 | 228.84 | 851.45 |
| KES131 | 14.059 | 3.028 | 25.523 | 11.978 | 12.365 | 236.43 | 909.36 |
| KES132 | 13.398 | 2.974 | 24.399 | 11.238 | 11.641 | 229.97 | 898.91 |
| KES133 | 18.683 | 2.635 | 32.077 | 15.425 | 10.777 | 223.86 | 815.38 |
| KES134 | 18.337 | 2.589 | 31.725 | 14.842 | 10.743 | 228.00 | 845.59 |
| KES135 | 18.254 | 2.597 | 31.339 | 14.783 | 10.422 | 219.00 | 825.56 |
| KES136 | 18.315 | 2.525 | 31.571 | 15.699 | 10.869 | 223.17 | 840.34 |
| KES137 | 22.702 | 3.195 | 38.784 | 19.658 | 13.930 | 268.31 | 1086.05 |
| KES138 | 22.762 | 3.146 | 38.134 | 18.208 | 13.533 | 264.69 | 1097.41 |
| KES139 | 19.271 | 2.711 | 34.681 | 16.305 | 11.476 | 229.35 | 925.65 |
| KES140 | 18.344 | 2.537 | 31.375 | 14.801 | 10.215 | 219.61 | 857.65 |
| KES141 | 20.110 | 2.851 | 36.552 | 17.767 | 11.825 | 233.56 | 960.06 |
| KES147 | 21.444 | 5.603 | 66.274 | 24.952 | 22.347 | 316.57 | 1715.17 |
| KES148 | 21.684 | 5.625 | 67.561 | 25.347 | 22.500 | 321.95 | 1784.51 |
| KES149 | 21.508 | 5.581 | 66.306 | 26.489 | 22.511 | 316.21 | 1743.50 |
| KES150 | 21.478 | 5.544 | 66.428 | 25.986 | 22.609 | 316.73 | 1750.68 |
| KES151 | 21.214 | 5.584 | 65.721 | 22.902 | 22.521 | 311.16 | 1747.92 |
| KES152 | 21.202 | 5.558 | 65.416 | 22.666 | 22.351 | 311.04 | 1700.45 |
| KES153 | 21.069 | 5.485 | 65.030 | 21.456 | 22.467 | 308.81 | 1710.06 |
| KES154 | 21.082 | 5.482 | 65.099 | 22.444 | 22.142 | 308.32 | 1718.73 |
| KES155 | 20.844 | 5.418 | 64.202 | 22.293 | 22.316 | 302.67 | 1682.00 |
| KES156 | 21.071 | 5.439 | 64.987 | 22.013 | 22.303 | 309.84 | 1691.42 |
| KES157 | 21.358 | 5.540 | 66.442 | 22.707 | 22.613 | 311.22 | 1739.10 |
| KES158 | 21.185 | 5.499 | 65.696 | 22.957 | 22.452 | 312.71 | 1721.73 |
| KES159 | 21.105 | 5.528 | 64.949 | 22.359 | 22.593 | 308.21 | 1743.67 |
| KES160 | 21.118 | 5.584 | 65.594 | 22.066 | 22.800 | 312.69 | 1734.27 |
| KES161 | 21.149 | 5.530 | 65.393 | 22.189 | 22.671 | 309.77 | 1718.06 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES162 | 21.311 | 5.643 | 66.298 | 23.054 | 22.628 | 310.99 | 1734.78 |
| KES163 | 21.102 | 5.544 | 65.534 | 22.824 | 22.782 | 312.64 | 1760.31 |
| KES164 | 20.952 | 5.546 | 64.714 | 23.100 | 22.595 | 307.76 | 1707.58 |
| KES165 | 22.489 | 6.496 | 41.592 | 17.535 | 25.973 | 407.82 | 1490.41 |
| KES166 | 23.140 | 6.681 | 43.134 | 16.621 | 26.923 | 429.03 | 1524.25 |
| KES167 | 23.010 | 6.648 | 43.093 | 17.811 | 27.036 | 426.24 | 1497.54 |
| KES168 | 16.655 | 5.830 | 33.267 | 13.272 | 22.027 | 441.63 | 1267.72 |
| KES169 | 11.793 | 3.918 | 32.326 | 11.676 | 14.535 | 274.42 | 1001.09 |
| KES170 | 13.687 | 4.230 | 35.612 | 11.985 | 15.283 | 252.15 | 1116.08 |
| KES171 | 13.704 | 4.275 | 35.343 | 12.671 | 15.139 | 257.94 | 1106.16 |
| KES172 | 13.729 | 4.287 | 35.617 | 11.968 | 15.128 | 251.73 | 1144.05 |
| KES173 | 13.712 | 4.212 | 35.838 | 10.922 | 15.645 | 257.93 | 1106.66 |
| KES174 | 13.771 | 4.261 | 35.545 | 12.473 | 15.591 | 254.52 | 1132.21 |
| KES175 | 14.097 | 4.337 | 36.508 | 13.225 | 15.798 | 262.09 | 1150.87 |
| KES176 | 12.091 | 2.181 | 37.151 | 14.023 | 9.806 | 108.12 | 438.15 |
| KES177 | 12.370 | 2.231 | 38.086 | 14.063 | 9.860 | 110.37 | 427.40 |
| KES178 | 12.594 | 2.266 | 38.870 | 14.677 | 9.859 | 113.54 | 443.15 |
| KES179 | 12.709 | 2.300 | 39.494 | 14.646 | 10.071 | 117.33 | 457.02 |
| KES180 | 12.480 | 2.248 | 38.622 | 15.097 | 10.073 | 112.76 | 437.94 |
| KES181 | 12.240 | 2.238 | 37.829 | 14.261 | 9.928 | 113.17 | 450.57 |
| KES182 | 26.517 | 8.199 | 69.648 | 29.811 | 31.012 | 467.03 | 2471.29 |
| KES183 | 26.540 | 8.234 | 69.779 | 27.901 | 30.949 | 461.43 | 2473.10 |
| KES184 | 25.809 | 8.151 | 67.446 | 19.883 | 29.187 | 418.42 | 2498.14 |
| KES185 | 25.691 | 8.067 | 66.600 | 18.292 | 28.896 | 411.40 | 2454.55 |
| KES186 | 25.680 | 8.134 | 66.931 | 20.246 | 29.068 | 412.46 | 2466.16 |
| KES187 | 25.670 | 8.042 | 66.694 | 20.422 | 28.783 | 412.95 | 2462.29 |
| KES188 | 25.507 | 7.930 | 66.186 | 18.401 | 28.342 | 410.07 | 2430.75 |
| KES189 | 25.281 | 7.861 | 65.213 | 18.360 | 27.944 | 405.02 | 2398.33 |
| KES190 | 25.569 | 8.010 | 66.220 | 17.426 | 29.487 | 414.18 | 2446.43 |
| KES191 | 25.523 | 7.972 | 66.242 | 20.533 | 28.602 | 410.93 | 2399.85 |
| KES192 | 25.720 | 8.167 | 66.966 | 19.995 | 28.783 | 415.12 | 2456.40 |
| KES193 | 25.842 | 8.098 | 67.266 | 20.970 | 29.250 | 412.97 | 2474.18 |
| KES194 | 26.149 | 8.179 | 67.629 | 21.759 | 29.216 | 417.98 | 2471.33 |
| KES195 | 25.571 | 8.028 | 66.583 | 19.675 | 28.943 | 406.59 | 2434.57 |
| KES196 | 26.045 | 8.153 | 67.314 | 20.589 | 29.287 | 417.37 | 2491.39 |
| KES197 | 25.583 | 8.055 | 66.126 | 21.096 | 28.888 | 412.57 | 2454.41 |
| KES198 | 25.699 | 8.030 | 66.710 | 19.684 | 28.680 | 412.95 | 2475.08 |
| KES199 | 25.517 | 7.989 | 65.856 | 20.819 | 29.604 | 407.90 | 2422.61 |
| KES200 | 25.889 | 8.078 | 66.757 | 19.706 | 29.265 | 413.58 | 2478.39 |
| KES201 | 26.031 | 8.106 | 67.568 | 20.194 | 27.209 | 416.66 | 2465.15 |
| KES202 | 25.697 | 8.034 | 66.722 | 20.588 | 29.729 | 414.42 | 2486.70 |
| KES203 | 25.985 | 8.070 | 67.222 | 20.752 | 26.594 | 413.34 | 2498.88 |
| KES204 | 25.724 | 8.050 | 66.544 | 22.494 | 29.695 | 414.79 | 2457.83 |
| KES205 | 26.073 | 8.166 | 66.857 | 21.837 | 29.788 | 416.12 | 2517.01 |
| KES206 | 25.629 | 7.978 | 66.209 | 21.507 | 29.988 | 408.31 | 2410.11 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES207 | 25.882 | 8.137 | 67.020 | 22.939 | 29.890 | 416.64 | 2471.22 |
| KES208 | 25.811 | 8.064 | 66.509 | 20.777 | 28.844 | 415.73 | 2451.25 |
| KES209 | 25.029 | 7.853 | 64.897 | 21.638 | 28.981 | 403.88 | 2447.83 |
| KES210 | 25.894 | 8.130 | 66.544 | 20.697 | 30.406 | 413.49 | 2470.63 |
| KES211 | 25.462 | 7.955 | 65.414 | 22.292 | 29.415 | 404.95 | 2412.56 |
| KES212 | 25.730 | 8.058 | 65.988 | 25.549 | 29.763 | 407.46 | 2452.35 |
| KES213 | 25.657 | 8.078 | 65.993 | 27.378 | 32.335 | 412.66 | 2468.88 |
| KES214 | 25.640 | 8.074 | 66.313 | 25.305 | 32.822 | 416.92 | 2478.55 |
| KES215 | 25.408 | 8.013 | 65.861 | 23.382 | 32.800 | 408.35 | 2457.09 |
| KES216 | 25.493 | 7.975 | 65.618 | 27.986 | 29.589 | 403.43 | 2446.01 |
| KES217 | 27.119 | 8.525 | 71.489 | 24.908 | 31.116 | 480.60 | 2520.26 |
| KES218 | 22.238 | 5.730 | 67.632 | 23.219 | 23.270 | 321.58 | 1759.16 |
| KES219 | 22.079 | 5.715 | 68.061 | 24.188 | 23.765 | 328.16 | 1798.31 |
| KES220 | 27.418 | 8.519 | 72.189 | 24.535 | 30.981 | 481.99 | 2528.94 |
| KES221 | 27.424 | 8.554 | 72.106 | 24.587 | 31.780 | 478.57 | 2571.68 |
| KES222 | 27.225 | 8.415 | 71.244 | 25.246 | 30.832 | 483.66 | 2546.45 |
| KES223 | 27.474 | 8.588 | 72.820 | 25.627 | 31.147 | 483.02 | 2561.90 |
| KES224 | 26.902 | 8.337 | 70.803 | 25.843 | 30.710 | 479.42 | 2532.24 |
| KES225 | 27.007 | 8.347 | 70.711 | 25.111 | 30.613 | 475.64 | 2494.74 |
| KES226 | 26.931 | 8.379 | 70.738 | 25.320 | 30.603 | 464.43 | 2501.78 |
| KES227 | 27.371 | 8.490 | 71.762 | 26.053 | 31.673 | 487.99 | 2530.47 |
| KES228 | 26.680 | 8.281 | 69.932 | 25.877 | 30.576 | 469.43 | 2486.66 |
| KES229 | 27.299 | 8.457 | 71.750 | 25.982 | 31.233 | 475.94 | 2497.65 |
| KES230 | 27.160 | 8.352 | 71.250 | 25.919 | 31.282 | 479.49 | 2518.08 |
| KES231 | 27.201 | 8.363 | 71.354 | 25.880 | 31.302 | 480.38 | 2541.52 |
| KES232 | 26.757 | 8.340 | 70.113 | 25.519 | 30.735 | 470.18 | 2517.46 |
| KES233 | 26.986 | 8.363 | 70.579 | 26.490 | 31.513 | 476.37 | 2513.24 |
| KES234 | 27.063 | 8.402 | 71.372 | 25.121 | 30.808 | 486.16 | 2523.42 |
| KES235 | 27.008 | 8.375 | 70.832 | 26.390 | 31.568 | 476.14 | 2496.70 |
| KES236 | 27.663 | 8.708 | 72.910 | 27.351 | 31.558 | 488.15 | 2593.29 |
| KES237 | 27.020 | 8.462 | 71.186 | 26.229 | 31.490 | 477.27 | 2523.33 |
| KES238 | 26.587 | 8.201 | 69.161 | 24.969 | 30.836 | 466.98 | 2460.60 |
| KES239 | 27.026 | 8.405 | 70.606 | 25.459 | 30.763 | 472.29 | 2487.98 |
| KES240 | 26.888 | 8.396 | 70.582 | 26.907 | 30.721 | 475.06 | 2513.00 |
| KES241 | 26.546 | 8.376 | 69.865 | 26.710 | 31.520 | 476.18 | 2493.90 |
| KES242 | 27.101 | 8.400 | 70.930 | 26.741 | 30.897 | 478.50 | 2539.37 |
| KES243 | 22.571 | 4.699 | 66.310 | 28.269 | 19.689 | 277.12 | 1409.54 |
| KES244 | 22.819 | 4.776 | 67.879 | 27.147 | 20.515 | 282.92 | 1444.56 |
| KES245 | 22.393 | 4.708 | 66.139 | 26.231 | 20.242 | 278.71 | 1413.50 |
| KES246 | 22.852 | 4.760 | 67.830 | 27.431 | 20.112 | 284.11 | 1435.52 |
| KES247 | 22.241 | 4.653 | 65.384 | 27.170 | 19.938 | 278.96 | 1403.26 |
| KES248 | 22.355 | 4.704 | 66.143 | 27.656 | 19.758 | 273.00 | 1412.07 |
| KES249 | 22.405 | 4.707 | 66.418 | 27.386 | 20.304 | 283.90 | 1402.61 |
| KES250 | 21.541 | 4.523 | 64.307 | 21.972 | 19.140 | 249.40 | 1382.31 |
| KES251 | 21.279 | 4.516 | 63.369 | 21.318 | 18.811 | 245.60 | 1373.48 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES252 | 21.923 | 4.568 | 66.521 | 22.504 | 19.172 | 252.69 | 1394.15 |
| KES253 | 22.467 | 4.719 | 67.002 | 22.971 | 19.422 | 253.15 | 1443.29 |
| KES254 | 22.078 | 4.625 | 65.967 | 22.773 | 19.435 | 249.86 | 1408.30 |
| KES255 | 15.771 | 2.325 | 36.170 | 13.981 | 9.503 | 133.24 | 1021.44 |
| KES256 | 15.593 | 2.281 | 36.257 | 13.707 | 9.470 | 134.26 | 1003.52 |
| KES257 | 15.980 | 2.338 | 36.831 | 14.322 | 9.608 | 133.86 | 1033.19 |
| KES258 | 15.602 | 2.256 | 35.845 | 14.143 | 9.389 | 128.44 | 1017.40 |
| KES259 | 15.830 | 2.322 | 36.572 | 14.451 | 9.485 | 130.09 | 1022.09 |
| KES260 | 16.018 | 2.334 | 36.977 | 14.220 | 9.676 | 133.47 | 1025.92 |
| KES261 | 15.533 | 2.280 | 35.433 | 14.451 | 9.184 | 128.75 | 1032.70 |
| KES262 | 15.843 | 2.303 | 36.505 | 14.444 | 9.521 | 130.37 | 1038.71 |
| KES263 | 15.925 | 2.303 | 36.824 | 14.732 | 9.632 | 134.68 | 1037.79 |
| KES264 | 15.793 | 2.318 | 36.530 | 14.932 | 9.670 | 132.14 | 1033.75 |
| KES265 | 15.971 | 2.303 | 36.654 | 14.392 | 9.468 | 134.41 | 1051.70 |
| KES266 | 15.824 | 2.315 | 36.447 | 14.942 | 9.274 | 134.76 | 1033.21 |
| KES267 | 16.010 | 2.317 | 36.677 | 14.522 | 9.423 | 129.95 | 1050.70 |
| KES268 | 15.880 | 2.253 | 36.895 | 14.768 | 9.537 | 130.31 | 1039.14 |
| KES269 | 15.684 | 2.277 | 36.214 | 15.171 | 9.393 | 129.33 | 1008.37 |
| KES270 | 15.899 | 2.272 | 36.543 | 15.733 | 9.606 | 137.34 | 1049.33 |
| KES271 | 15.546 | 2.233 | 35.427 | 14.803 | 9.224 | 134.14 | 1003.14 |
| KES272 | 15.713 | 2.298 | 36.054 | 14.039 | 9.369 | 131.34 | 1019.06 |
| KES273 | 15.737 | 2.276 | 35.963 | 15.594 | 9.567 | 135.50 | 1030.19 |
| KES274 | 15.953 | 2.272 | 36.294 | 15.266 | 9.361 | 135.01 | 1050.92 |
| KES275 | 15.884 | 2.302 | 36.562 | 15.425 | 9.529 | 132.17 | 1041.78 |
| KES276 | 15.817 | 2.292 | 36.016 | 15.204 | 9.319 | 131.95 | 1021.74 |
| KES277 | 15.853 | 2.325 | 36.689 | 12.638 | 9.466 | 140.58 | 1014.71 |
| KES278 | 22.199 | 4.742 | 66.079 | 20.069 | 18.861 | 256.46 | 1418.18 |
| KES279 | 22.503 | 4.777 | 66.752 | 19.948 | 19.216 | 253.06 | 1427.60 |
| KES281 | 21.604 | 4.626 | 64.301 | 20.296 | 18.833 | 250.88 | 1406.61 |
| KES282 | 21.138 | 4.554 | 62.953 | 20.478 | 18.153 | 249.90 | 1334.26 |
| KES283 | 21.182 | 4.521 | 62.992 | 19.409 | 18.547 | 245.42 | 1375.95 |
| KES284 | 20.599 | 4.432 | 61.154 | 19.384 | 17.970 | 243.68 | 1313.21 |
| KES285 | 13.664 | 3.885 | 27.831 | 7.262 | 14.257 | 295.11 | 962.28 |
| KES286 | 13.900 | 3.942 | 28.346 | 8.217 | 14.467 | 299.41 | 982.87 |
| KES287 | 12.785 | 3.385 | 27.342 | 7.766 | 12.358 | 267.22 | 841.09 |
| KES288 | 13.162 | 3.394 | 28.001 | 8.231 | 12.529 | 267.96 | 856.84 |
| KES289 | 13.062 | 3.471 | 28.010 | 7.522 | 12.468 | 269.45 | 815.28 |
| KES290 | 12.923 | 3.434 | 27.566 | 7.290 | 12.250 | 266.68 | 816.06 |
| KES291 | 13.309 | 3.548 | 28.575 | 7.814 | 12.737 | 273.71 | 868.19 |
| KES292 | 12.979 | 3.389 | 27.626 | 7.089 | 12.353 | 268.27 | 818.75 |
| KES293 | 13.155 | 3.506 | 28.348 | 7.820 | 12.610 | 272.29 | 827.79 |
| KES294 | 13.069 | 3.437 | 28.205 | 9.009 | 12.853 | 268.28 | 841.30 |
| KES295 | 13.162 | 3.470 | 28.312 | 8.471 | 12.867 | 271.02 | 865.72 |
| KES296 | 13.077 | 3.554 | 28.334 | 8.241 | 12.590 | 266.79 | 826.19 |
| KES297 | 21.935 | 5.813 | 68.108 | 22.256 | 23.503 | 296.20 | 1844.35 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES298 | 21.540 | 5.738 | 66.551 | 21.492 | 23.198 | 290.06 | 1804.10 |
| KES299 | 21.947 | 5.842 | 68.403 | 21.916 | 23.188 | 293.37 | 1838.15 |
| KES300 | 32.286 | 10.058 | 83.385 | 18.635 | 31.520 | 503.84 | 3073.72 |
| KES301 | 30.622 | 9.526 | 78.543 | 19.074 | 29.801 | 468.07 | 2868.59 |
| KES302 | 31.227 | 9.736 | 80.826 | 21.006 | 30.699 | 482.34 | 2954.73 |
| KES303 | 31.163 | 9.756 | 80.543 | 20.579 | 30.932 | 479.58 | 2954.66 |
| KES304 | 31.407 | 9.744 | 81.276 | 18.886 | 30.757 | 479.03 | 2932.87 |
| KES305 | 31.251 | 9.806 | 81.318 | 21.081 | 31.065 | 487.63 | 2986.35 |
| KES306 | 31.199 | 9.752 | 81.220 | 20.010 | 31.312 | 486.49 | 2962.88 |
| KES307 | 30.976 | 9.606 | 79.865 | 21.748 | 30.817 | 474.69 | 2907.20 |
| KES308 | 30.518 | 9.552 | 78.623 | 22.203 | 30.916 | 471.53 | 2843.30 |
| KES309 | 31.038 | 9.670 | 80.415 | 21.478 | 30.826 | 475.26 | 2949.40 |
| KES310 | 31.292 | 9.765 | 80.624 | 20.696 | 31.717 | 472.64 | 2937.31 |
| KES311 | 33.038 | 10.056 | 87.753 | 16.178 | 38.372 | 554.66 | 3073.56 |
| KES312 | 33.197 | 10.149 | 89.196 | 18.564 | 38.451 | 566.79 | 3140.98 |
| KES313 | 32.957 | 9.938 | 87.414 | 18.119 | 38.464 | 554.32 | 3025.75 |
| KES314 | 33.342 | 10.039 | 88.730 | 17.515 | 38.395 | 556.85 | 3105.69 |
| KES315 | 33.431 | 10.058 | 89.358 | 19.002 | 37.905 | 568.01 | 3165.78 |
| KES316 | 32.144 | 9.602 | 85.301 | 17.099 | 37.740 | 528.43 | 2978.97 |
| KES317 | 31.764 | 9.575 | 84.287 | 16.737 | 36.987 | 525.38 | 2953.36 |
| KES318 | 13.272 | 4.836 | 21.620 | 4.668 | 18.032 | 382.62 | 1071.82 |
| KES319 | 13.780 | 5.091 | 23.442 | 3.833 | 18.495 | 378.37 | 1120.73 |
| KES320 | 31.565 | 10.436 | 58.373 | 9.902 | 38.893 | 581.43 | 2374.54 |
| KES321 | 22.773 | 7.575 | 44.098 | 8.385 | 28.707 | 454.60 | 1844.28 |
| KES322 | 16.307 | 5.959 | 28.089 | 5.626 | 20.453 | 429.72 | 1293.63 |
| KES323 | 19.663 | 6.448 | 36.454 | 7.581 | 24.665 | 388.31 | 1484.77 |
| KES324 | 19.834 | 6.468 | 36.841 | 7.073 | 25.316 | 402.28 | 1479.78 |
| KES325 | 19.679 | 6.349 | 35.771 | 6.331 | 24.076 | 387.10 | 1460.78 |
| KES326 | 19.602 | 6.511 | 36.205 | 6.580 | 24.719 | 389.35 | 1465.98 |
| KES327 | 19.810 | 6.366 | 36.138 | 6.520 | 24.158 | 388.69 | 1500.61 |
| KES328 | 19.838 | 6.431 | 36.580 | 7.396 | 25.188 | 385.51 | 1497.07 |
| KES329 | 19.944 | 6.448 | 36.819 | 6.489 | 25.245 | 392.02 | 1529.68 |
| KES330 | 19.637 | 6.454 | 36.438 | 7.806 | 25.727 | 392.45 | 1515.33 |
| KES331 | 19.756 | 6.411 | 36.570 | 7.716 | 25.598 | 393.54 | 1488.17 |
| KES335 | 19.746 | 6.387 | 36.374 | 6.240 | 24.831 | 390.60 | 1508.16 |
| KES336 | 19.620 | 6.392 | 36.331 | 6.945 | 24.791 | 384.84 | 1489.90 |
| KES337 | 19.513 | 6.342 | 36.111 | 7.008 | 25.144 | 390.61 | 1494.35 |
| KES338 | 19.410 | 6.310 | 35.616 | 8.142 | 24.693 | 382.33 | 1419.51 |
| KES339 | 19.635 | 6.355 | 36.118 | 8.093 | 25.193 | 390.98 | 1489.64 |
| KES340 | 19.798 | 6.503 | 36.919 | 6.795 | 25.746 | 396.55 | 1505.58 |
| KES341 | 13.268 | 4.875 | 22.393 | 11.705 | 18.320 | 362.57 | 1059.31 |
| KES342 | 17.479 | 6.136 | 37.125 | 7.924 | 23.131 | 408.26 | 1301.71 |
| KES343 | 13.295 | 4.875 | 22.116 | 4.655 | 17.439 | 363.39 | 1045.12 |
| KES344 | 13.428 | 4.800 | 22.579 | 4.349 | 17.757 | 367.42 | 1063.14 |
| KES345 | 8.241 | 2.986 | 12.218 | 2.491 | 11.910 | 255.78 | 601.75 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES346 | 8.224 | 3.062 | 11.856 | 6.576 | 11.796 | 249.17 | 621.77 |
| KES347 | 22.936 | 6.872 | 59.085 | 14.296 | 28.066 | 394.87 | 2056.29 |
| KES348 | 23.093 | 6.919 | 59.751 | 14.207 | 28.306 | 394.24 | 2094.73 |
| KES349 | 23.221 | 6.967 | 60.100 | 13.431 | 28.836 | 407.73 | 2073.34 |
| KES350 | 34.178 | 10.379 | 83.811 | 19.628 | 40.965 | 588.31 | 2855.14 |
| KES351 | 34.080 | 10.359 | 83.550 | 19.962 | 41.228 | 586.59 | 2848.28 |
| KES352 | 34.147 | 10.387 | 83.723 | 20.916 | 41.005 | 590.82 | 2836.32 |
| KES353 | 33.892 | 10.340 | 82.656 | 19.208 | 40.904 | 565.84 | 2815.99 |
| KES354 | 34.210 | 10.373 | 83.587 | 19.927 | 40.608 | 590.06 | 2861.44 |
| KES355 | 33.920 | 10.317 | 82.780 | 20.748 | 40.197 | 582.25 | 2827.82 |
| KES356 | 33.859 | 10.289 | 82.922 | 18.732 | 40.172 | 573.26 | 2811.84 |
| KES357 | 34.225 | 10.350 | 83.549 | 20.781 | 40.658 | 580.21 | 2865.36 |
| KES358 | 33.895 | 10.319 | 82.384 | 19.128 | 40.891 | 577.33 | 2819.89 |
| KES359 | 33.367 | 10.097 | 81.402 | 19.530 | 40.622 | 570.17 | 2790.08 |
| KES360 | 33.830 | 10.274 | 82.558 | 19.903 | 40.506 | 578.57 | 2812.84 |
| KES361 | 33.904 | 10.319 | 82.701 | 18.645 | 40.739 | 569.68 | 2842.64 |
| KES362 | 33.917 | 10.273 | 82.434 | 19.817 | 40.645 | 576.76 | 2834.13 |
| KES363 | 33.883 | 10.333 | 83.122 | 19.194 | 40.679 | 582.41 | 2850.28 |
| KES364 | 32.572 | 9.907 | 86.627 | 19.556 | 38.800 | 536.17 | 3100.54 |
| KES365 | 31.721 | 9.526 | 83.779 | 19.862 | 37.805 | 513.49 | 2993.74 |
| KES366 | 32.157 | 9.661 | 85.045 | 20.705 | 38.285 | 533.29 | 3045.42 |
| KES367 | 32.122 | 9.685 | 84.618 | 19.210 | 37.799 | 525.32 | 3048.63 |
| KES368 | 32.064 | 9.600 | 84.846 | 18.812 | 38.095 | 527.47 | 3056.94 |
| KES369 | 32.615 | 9.912 | 87.091 | 19.618 | 38.945 | 538.77 | 3138.90 |
| KES370 | 18.868 | 4.650 | 37.378 | 10.469 | 13.057 | 408.55 | 1612.92 |
| KES371 | 18.882 | 4.638 | 37.182 | 9.584 | 13.076 | 409.47 | 1586.36 |
| KES372 | 18.838 | 4.618 | 37.192 | 9.647 | 13.074 | 410.45 | 1583.05 |
| KES373 | 18.661 | 4.530 | 36.808 | 9.590 | 12.987 | 401.57 | 1575.62 |
| KES374 | 18.411 | 3.797 | 54.827 | 13.961 | 16.656 | 231.75 | 1160.76 |
| KES375 | 18.452 | 3.842 | 55.161 | 13.636 | 16.697 | 228.53 | 1159.47 |
| KES376 | 22.667 | 4.657 | 69.425 | 16.326 | 20.968 | 277.54 | 1487.40 |
| KES377 | 22.893 | 4.692 | 68.168 | 16.523 | 20.499 | 274.69 | 1500.84 |
| KES378 | 22.549 | 4.638 | 67.252 | 16.032 | 20.179 | 275.41 | 1456.54 |
| KES379 | 22.090 | 4.565 | 65.872 | 23.510 | 20.931 | 270.92 | 1423.30 |
| KES380 | 21.686 | 4.495 | 64.121 | 22.843 | 20.610 | 270.37 | 1392.83 |
| KES381 | 21.902 | 4.509 | 64.941 | 23.047 | 20.668 | 271.57 | 1407.17 |
| KES382 | 21.733 | 4.483 | 64.780 | 22.398 | 20.853 | 269.25 | 1386.72 |
| KES383 | 21.901 | 4.505 | 64.969 | 22.491 | 20.919 | 265.03 | 1404.26 |
| KES384 | 21.981 | 4.510 | 65.119 | 23.491 | 20.553 | 270.47 | 1397.74 |
| KES385 | 21.954 | 4.493 | 64.810 | 23.243 | 20.891 | 268.61 | 1398.79 |
| KES386 | 21.759 | 4.457 | 64.144 | 23.465 | 20.343 | 268.17 | 1378.65 |
| KES387 | 21.479 | 4.425 | 63.885 | 23.543 | 21.200 | 261.10 | 1368.18 |
| KES388 | 21.570 | 4.481 | 63.902 | 23.391 | 20.667 | 271.63 | 1373.26 |
| KES389 | 21.791 | 4.483 | 64.749 | 23.312 | 20.636 | 270.28 | 1384.20 |
| KES390 | 21.505 | 4.408 | 63.647 | 23.264 | 20.452 | 266.94 | 1375.40 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES391 | 21.783 | 4.464 | 64.502 | 24.330 | 20.659 | 268.99 | 1389.23 |
| KES392 | 21.999 | 4.592 | 65.267 | 24.559 | 21.028 | 274.54 | 1424.55 |
| KES393 | 22.070 | 4.545 | 65.071 | 24.026 | 20.858 | 272.70 | 1399.59 |
| KES394 | 21.438 | 4.755 | 63.725 | 24.557 | 22.404 | 278.22 | 1497.80 |
| KES403 | 18.6984 | 4.8091 | 36.9758 | 10.375 | 12.662 | 408.80 | 1525.15 |
| KES404 | 18.9480 | 4.8863 | 37.3595 | 9.895 | 12.960 | 409.44 | 1520.82 |
| KES405 | 18.8737 | 4.9538 | 37.5761 | 9.280 | 13.050 | 414.17 | 1569.33 |
| KES406 | 19.0287 | 4.8483 | 37.3830 | 9.279 | 12.966 | 413.88 | 1512.42 |
| KES407 | 18.9580 | 4.8241 | 37.3435 | 9.186 | 13.079 | 420.63 | 1557.80 |
| KES408 | 30.5571 | 6.5308 | 73.5349 | 15.819 | 25.532 | 473.35 | 1793.56 |
| KES409 | 30.4043 | 6.4689 | 72.7759 | 14.470 | 24.896 | 462.97 | 1804.22 |
| KES410 | 30.4700 | 6.4394 | 72.7818 | 14.784 | 25.106 | 473.27 | 1865.11 |
| KES411 | 17.1074 | 4.5014 | 31.7168 | 5.830 | 16.849 | 329.80 | 1026.57 |
| KES412 | 23.4231 | 5.5419 | 65.1304 | 13.680 | 21.924 | 380.95 | 1570.32 |
| KES413 | 23.6847 | 5.7069 | 66.0873 | 14.729 | 22.170 | 381.64 | 1618.76 |
| KES414 | 13.0658 | 3.5475 | 28.2627 | 5.951 | 13.457 | 277.62 | 831.72 |
| KES415 | 13.0111 | 3.4938 | 28.1250 | 4.901 | 13.287 | 277.47 | 845.38 |
| KES416 | 13.0394 | 3.4563 | 28.0091 | 5.601 | 13.468 | 271.48 | 814.16 |
| KES417 | 13.3226 | 3.5484 | 28.5445 | 5.448 | 13.455 | 282.36 | 831.63 |
| KES418 | 11.0051 | 3.5366 | 18.3181 | 3.788 | 12.855 | 230.83 | 884.41 |
| KES419 | 12.6294 | 2.8882 | 25.6263 | 5.177 | 11.410 | 230.88 | 733.34 |
| KES420 | 12.3558 | 2.8395 | 24.9100 | 5.001 | 10.777 | 226.32 | 736.40 |
| KES421 | 12.4832 | 2.8739 | 25.0549 | 4.951 | 11.431 | 229.79 | 699.31 |
| KES422 | 12.7203 | 2.9511 | 25.9499 | 5.368 | 11.415 | 237.51 | 742.90 |
| KES423 | 12.6512 | 2.9134 | 25.6299 | 5.124 | 11.372 | 237.05 | 760.95 |
| KES424 | 12.8330 | 4.0405 | 20.2812 | 5.668 | 16.438 | 251.81 | 1187.99 |
| KES425 | 11.4666 | 3.6696 | 18.0070 | 4.513 | 14.529 | 232.15 | 1052.89 |
| KES426 | 10.0833 | 3.2777 | 15.8362 | 3.828 | 12.910 | 207.28 | 922.36 |
| KES427 | 12.4422 | 3.9487 | 19.6770 | 4.713 | 15.722 | 244.87 | 1119.19 |
| KES428 | 12.1851 | 3.9112 | 19.3025 | 4.853 | 15.659 | 245.13 | 1128.06 |
| KES429 | 12.5124 | 2.8587 | 25.2505 | 4.951 | 11.163 | 234.61 | 715.84 |
| KES430 | 7.7707 | 1.9881 | 14.4559 | 3.478 | 5.888 | 143.87 | 601.03 |
| KES431 | 5.5919 | 2.5009 | 11.8372 | 2.390 | 9.230 | 177.93 | 655.08 |
| KES432 | 5.6446 | 2.5569 | 12.0490 | 2.699 | 9.287 | 179.05 | 625.76 |
| KES433 | 6.9221 | 2.9082 | 13.8670 | 3.414 | 10.468 | 214.65 | 726.75 |
| KES434 | 5.7180 | 2.5524 | 12.0568 | 2.555 | 9.389 | 179.99 | 664.77 |
| KES435 | 5.6564 | 2.5498 | 12.1754 | 2.604 | 9.181 | 183.72 | 651.85 |
| KES436 | 10.8734 | 3.4761 | 17.3976 | 2.307 | 12.020 | 230.14 | 847.68 |
| KES437 | 8.2991 | 2.1110 | 15.6869 | 3.244 | 7.856 | 190.40 | 473.21 |
| KES438 | 12.2593 | 2.8025 | 24.6844 | 5.297 | 10.991 | 226.64 | 722.96 |
| KES439 | 13.5886 | 4.6369 | 24.2311 | 4.586 | 17.157 | 291.24 | 1198.14 |
| KES440 | 13.7891 | 4.7886 | 24.5952 | 5.526 | 18.188 | 291.23 | 1224.65 |
| KES441 | 13.9377 | 4.8021 | 24.4982 | 5.524 | 17.541 | 287.16 | 1242.32 |
| KES442 | 14.4518 | 3.8135 | 22.2916 | 6.388 | 13.840 | 253.86 | 966.03 |
| KES443 | 14.1689 | 3.7149 | 22.2344 | 5.224 | 13.849 | 247.33 | 974.86 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES444 | 31.9521 | 7.4218 | 51.4894 | 11.620 | 24.794 | 509.00 | 2280.09 |
| KES445 | 31.3272 | 7.3637 | 50.2538 | 11.349 | 24.443 | 491.43 | 2287.27 |
| KES446 | 52.0942 | 11.7920 | 85.0438 | 18.957 | 37.924 | 667.11 | 3825.44 |
| KES447 | 15.6869 | 4.2114 | 25.3297 | 5.003 | 15.978 | 332.07 | 1096.17 |
| KES448 | 15.6193 | 4.2039 | 25.0959 | 5.877 | 16.123 | 333.92 | 1078.09 |
| KES449 | 14.9088 | 3.5524 | 29.0341 | 5.740 | 12.614 | 270.52 | 953.64 |
| KES450 | 14.7921 | 3.5618 | 28.5882 | 6.524 | 12.554 | 267.11 | 1009.66 |
| KES451 | 14.8386 | 3.6122 | 28.8970 | 6.785 | 12.709 | 269.77 | 1025.96 |
| KES452 | 15.2213 | 3.6240 | 29.5846 | 6.652 | 12.980 | 277.35 | 1011.58 |
| KES453 | 15.2858 | 3.6429 | 30.0362 | 8.753 | 12.788 | 276.17 | 1060.26 |
| KES454 | 15.2715 | 3.6437 | 29.8304 | 9.907 | 12.610 | 271.78 | 1056.56 |
| KES455 | 14.9732 | 3.5987 | 29.3933 | 11.328 | 12.522 | 264.65 | 1051.81 |
| KES456 | 15.2884 | 3.6011 | 29.9440 | 9.901 | 12.864 | 268.41 | 1068.75 |
| KES457 | 14.9262 | 3.6880 | 28.7628 | 11.027 | 12.799 | 284.63 | 1067.93 |
| KES458 | 15.0130 | 3.7183 | 28.8994 | 10.979 | 12.471 | 288.51 | 1074.79 |
| KES459 | 20.4488 | 2.5622 | 32.0187 | 14.742 | 11.456 | 223.17 | 1223.73 |
| KES460 | 18.3899 | 2.4081 | 28.3265 | 13.275 | 10.224 | 211.26 | 1075.56 |
| KES461 | 12.7344 | 2.8528 | 25.7944 | 8.515 | 10.795 | 260.73 | 811.65 |
| KES462 | 19.5437 | 2.4576 | 30.5184 | 13.498 | 10.615 | 216.58 | 1157.28 |
| KES463 | 17.2947 | 2.7668 | 26.5172 | 12.120 | 11.777 | 223.83 | 1087.92 |
| KES464 | 17.3617 | 2.7838 | 26.6021 | 11.385 | 11.733 | 223.00 | 1079.44 |
| KES465 | 16.9720 | 2.6915 | 26.0407 | 12.273 | 11.738 | 219.82 | 1057.36 |
| KES466 | 18.0331 | 2.8725 | 27.8690 | 12.480 | 12.123 | 228.63 | 1105.06 |
| KES467 | 33.4856 | 10.5001 | 89.2249 | 30.097 | 38.646 | 557.67 | 3259.10 |
| KES468 | 33.9262 | 10.6576 | 90.6396 | 31.513 | 39.560 | 562.23 | 3272.39 |
| KES469 | 33.5991 | 10.5507 | 89.7581 | 28.830 | 39.251 | 563.95 | 3249.68 |
| KES470 | 33.5406 | 10.4743 | 89.0426 | 27.957 | 38.794 | 558.85 | 3243.18 |
| KES471 | 19.1968 | 6.9301 | 33.2873 | 14.379 | 25.262 | 437.83 | 1511.53 |
| KES472 | 19.4300 | 7.0734 | 33.8885 | 14.674 | 25.044 | 439.30 | 1523.04 |
| KES473 | 18.8389 | 6.8508 | 32.6907 | 13.692 | 24.767 | 429.27 | 1471.49 |
| KES474 | 19.2682 | 7.0055 | 33.4866 | 11.968 | 24.874 | 431.74 | 1474.78 |
| KES475 | 34.5761 | 10.8432 | 91.8300 | 33.606 | 40.533 | 586.13 | 3345.93 |
| KES476 | 33.9473 | 10.7312 | 90.2319 | 32.318 | 39.218 | 569.62 | 3294.62 |
| KES477 | 19.0227 | 6.9006 | 33.1809 | 13.518 | 24.870 | 431.02 | 1475.18 |
| KES478 | 19.0305 | 6.9214 | 32.9596 | 14.090 | 24.429 | 432.91 | 1475.08 |
| KES479 | 27.8277 | 8.6740 | 73.6901 | 26.611 | 31.993 | 460.74 | 2697.21 |
| KES480 | 18.2562 | 6.7306 | 32.7622 | 13.269 | 25.699 | 414.71 | 1409.48 |
| KES481 | 17.5128 | 6.2818 | 37.0268 | 14.838 | 22.579 | 405.26 | 1353.44 |
| KES482 | 17.6275 | 6.2780 | 37.2931 | 15.264 | 23.253 | 405.02 | 1351.30 |
| KES483 | 31.5741 | 9.9270 | 83.9185 | 29.259 | 36.878 | 525.00 | 3070.40 |
| KES484 | 32.0131 | 10.0249 | 84.9341 | 29.328 | 37.336 | 538.90 | 3058.71 |
| KES485 | 27.1371 | 8.5562 | 72.1399 | 26.565 | 33.179 | 479.63 | 2636.71 |
| KES486 | 26.9037 | 8.4799 | 71.3702 | 27.824 | 31.016 | 469.69 | 2599.88 |
| KES487 | 25.4494 | 7.9996 | 67.8920 | 25.761 | 30.316 | 440.53 | 2461.66 |
| KES488 | 25.1949 | 7.8689 | 66.6994 | 25.775 | 31.562 | 432.95 | 2431.95 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | U | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES489 | 25.3886 | 8.0584 | 67.0920 | 25.146 | 30.120 | 449.83 | 2468.07 |
| KES490 | 26.1316 | 8.3212 | 69.0131 | 26.619 | 31.030 | 469.60 | 2555.18 |
| KES491 | 25.8937 | 8.2130 | 68.0719 | 26.447 | 30.229 | 451.68 | 2501.19 |
| KES492 | 25.5945 | 8.2047 | 67.6049 | 25.184 | 31.032 | 458.30 | 2482.05 |
| KES493 | 25.7404 | 8.1446 | 67.8549 | 26.071 | 31.576 | 451.04 | 2498.88 |
| KES494 | 25.8512 | 8.1788 | 67.6767 | 25.935 | 30.194 | 456.29 | 2472.39 |
| KES495 | 25.6351 | 8.0126 | 68.3588 | 26.774 | 30.651 | 445.63 | 2486.46 |
| KES496 | 14.6327 | 4.9248 | 39.0874 | 15.507 | 16.260 | 320.49 | 1451.74 |
| KES497 | 15.1105 | 5.1436 | 40.8512 | 16.558 | 17.119 | 332.03 | 1465.95 |
| KES498 | 15.0265 | 5.1718 | 40.2508 | 15.464 | 17.103 | 336.66 | 1460.82 |
| KES499 | 14.5524 | 4.8570 | 38.8110 | 15.207 | 17.196 | 325.84 | 1410.93 |
| KES500 | 14.3165 | 4.8720 | 38.4643 | 13.725 | 17.257 | 318.44 | 1406.70 |
| KES501 | 14.9112 | 5.0999 | 40.0343 | 15.788 | 16.603 | 335.03 | 1466.53 |
| KES502 | 14.4998 | 4.8862 | 38.7148 | 13.582 | 15.703 | 323.69 | 1388.65 |
| KES503 | 16.9261 | 2.1264 | 37.8922 | 14.720 | 7.392 | 178.28 | 1118.33 |
| KES504 | 16.8651 | 2.1202 | 37.6090 | 13.808 | 7.329 | 176.38 | 1093.95 |
| KES505 | 16.9863 | 2.1825 | 38.2973 | 14.154 | 7.623 | 178.29 | 1122.62 |
| KES506 | 17.0636 | 2.1561 | 38.4002 | 14.790 | 7.490 | 181.08 | 1122.77 |
| KES507 | 17.0939 | 2.1675 | 38.3852 | 14.676 | 7.560 | 176.79 | 1138.64 |
| KES508 | 17.0532 | 2.1523 | 38.4218 | 14.692 | 7.554 | 179.14 | 1127.77 |
| KES509 | 13.7880 | 3.0769 | 25.4609 | 10.252 | 11.904 | 242.26 | 946.24 |
| KES510 | 14.2432 | 3.2129 | 26.4590 | 11.037 | 12.022 | 252.18 | 953.88 |
| KES511 | 13.7991 | 3.0634 | 25.4913 | 10.401 | 11.800 | 244.62 | 928.03 |
| KES512 | 13.5129 | 2.9906 | 24.9500 | 10.230 | 11.758 | 236.30 | 901.45 |
| KES513 | 13.0226 | 2.9157 | 23.8222 | 8.986 | 11.327 | 227.33 | 858.56 |
| KES514 | 12.5167 | 2.2871 | 26.5513 | 11.782 | 6.892 | 207.60 | 940.05 |
| KES515 | 12.5917 | 2.3378 | 26.6433 | 10.942 | 7.001 | 211.24 | 913.69 |
| KES516 | 7.9307 | 2.4290 | 16.2118 | 7.183 | 9.909 | 151.46 | 731.00 |
| KES517 | 8.0900 | 2.4522 | 16.5633 | 7.154 | 9.725 | 156.10 | 769.59 |
| KES518 | 8.1271 | 2.4376 | 16.6350 | 7.308 | 10.101 | 152.88 | 742.74 |
| KES519 | 7.6571 | 2.3641 | 15.5750 | 6.551 | 9.599 | 150.42 | 706.88 |
| KES520 | 7.8731 | 2.3895 | 16.0689 | 6.991 | 9.540 | 150.96 | 743.26 |
| KES521 | 7.6946 | 2.3446 | 15.6697 | 6.803 | 9.581 | 150.57 | 742.42 |
| KES522 | 7.9215 | 2.4113 | 16.1880 | 6.664 | 9.757 | 150.04 | 743.56 |
| KES523 | 8.1047 | 2.4159 | 16.5592 | 7.795 | 9.995 | 149.48 | 773.45 |
| KES524 | 7.9134 | 2.3657 | 16.1576 | 7.925 | 9.638 | 152.53 | 735.93 |
| KES525 | 7.9728 | 2.4185 | 16.2512 | 6.915 | 9.847 | 156.91 | 756.99 |
| KES526 | 7.0457 | 2.2014 | 13.8876 | 6.055 | 8.941 | 152.63 | 662.00 |
| KES527 | 8.3865 | 1.8372 | 14.5878 | 7.427 | 5.558 | 183.97 | 583.75 |
| KES528 | 8.3574 | 1.8951 | 14.6689 | 6.891 | 5.485 | 182.30 | 591.95 |
| KES529 | 8.2102 | 1.8421 | 14.4299 | 6.618 | 5.665 | 175.63 | 573.73 |
| KES530 | 7.9375 | 1.7758 | 14.0328 | 6.803 | 5.239 | 174.87 | 561.14 |
| KES531 | 7.8446 | 1.7780 | 13.7940 | 6.235 | 5.283 | 171.51 | 542.89 |
| KES532 | 7.9364 | 1.7679 | 13.8743 | 6.689 | 5.322 | 174.22 | 568.52 |
| KES533 | 7.9706 | 1.7908 | 13.9677 | 7.015 | 5.387 | 175.00 | 579.31 |

Appendix 6 - continued. All concentrations are in parts per million (ppm).

| Sample ID | Ta | Tb | Th | $\mathbf{U}$ | Yb | Zn | Zr |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| KES534 | 22.5647 | 5.1553 | 68.2553 | 28.082 | 21.777 | 309.33 | 1571.82 |
| KES535 | 22.3827 | 5.0858 | 68.0606 | 27.401 | 21.634 | 309.96 | 1563.87 |
| KES536 | 18.9083 | 4.8288 | 37.6299 | 19.020 | 13.168 | 432.97 | 1575.45 |
| KES537 | 18.7959 | 4.8041 | 37.2496 | 19.408 | 12.829 | 427.45 | 1558.47 |
| KES538 | 18.7759 | 4.7718 | 36.9414 | 19.009 | 12.770 | 429.08 | 1556.03 |
| KES539 | 18.8399 | 4.7714 | 37.3189 | 19.604 | 12.936 | 435.20 | 1565.90 |
| KES540 | 16.9828 | 4.3436 | 31.5709 | 14.657 | 17.000 | 345.74 | 989.14 |
| KES541 | 11.0307 | 3.4767 | 17.5919 | 9.879 | 13.911 | 229.32 | 1029.82 |
| KES542 | 12.3366 | 3.9216 | 19.5630 | 10.657 | 15.647 | 262.00 | 1149.18 |

Appendix 7 - Results of concentration calculations for the standards analyzed for both the standard INAA procedure and the new ENAA procedure for titanium. NIST SRM 278 (obsidian) was used as a primary standard.
A. Standard INAA procedure (Ti-50(n, $\gamma$ ) Ti-51)

| Sample ID | $\begin{aligned} & \text { Mass } \\ & \text { (mg) } \end{aligned}$ | $\begin{gathered} \text { Counts } \\ (320 \mathrm{keV}) \end{gathered}$ | Calculated Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| AGV-1a | 200.61 | 7842 | 6680 | 6790 | 290 |
| AGV-1b | 200.15 | 8370 | 7120 |  |  |
| AGV-1c | 200.75 | 7724 | 6560 |  |  |
| JA-1a | 200.03 | 6015 | 6090 | 6210 | 270 |
| JA-1b | 200.37 | 6471 | 6510 |  |  |
| JA-1c | 200.70 | 5983 | 6010 |  |  |
| JB-2a | 200.47 | 7487 | 718 | 6390 | 803. |
| JB-2b | 200.37 | 6703 | 643 |  |  |
| JB-2c | 200.64 | 5818 | 5570 |  |  |
| JR-1a | 200.80 | -- | -- | 610. | 539. |
| JR-1b | 200.66 | 1135 | 1030 |  |  |
| JR-1c | 200.40 | 889 | 804. |  |  |
| SRM 1633381 | 100.29 | 7156 | 9210 | 10400 | 1330 |
| SRM 1633382 | 99.99 | 7078 | 9140 |  |  |
| SRM 1633383 | 99.43 | 7026 | 9110 |  |  |
| SRM 1633393 | 99.77 | 9137 | 11900 |  |  |
| SRM 1633394 | 99.73 | 8484 | 11100 |  |  |
| SRM 1633395 | 99.98 | 8926 | 11600 |  |  |

B. Epithermal INAA procedure $(\mathrm{Ti}-47(\mathrm{n}, \gamma) \mathrm{Sc}-47)$

| Sample ID | $\begin{aligned} & \text { Mass } \\ & \text { (mg) } \end{aligned}$ | $\begin{gathered} \text { Counts } \\ (159 \mathrm{keV}) \\ \hline \end{gathered}$ | Calculated Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| FA7257 | 99.56 | 28067 | 7960 | 7970 | 23 |
| FA7258 | 99.21 | 22443 | 7990 |  |  |
| AGV-1a | 200.26 | 45165 | 6350 | 6460 | 94 |
| AGV-1b | 201.62 | 42209 | 6520 |  |  |
| AGV-1c | 201.58 | 38098 | 6510 |  |  |
| JA-1a | 200.51 | 37117 | 5160 | 5240 | 71 |
| JA-1b | 200.91 | 34486 | 5290 |  |  |
| JA-1c | 200.58 | 30917 | 5260 |  |  |
| JB-2a | 199.79 | 52295 | 7380 | 7480 | 118 |
| JB-2b | 199.90 | 48686 | 7610 |  |  |
| JB-2c | 200.45 | 43233 | 7460 |  |  |
| JR-1a | 200.22 | 4905 | 705. | 678. | 57 |
| JR-1b | 200.80 | 3873 | 613. |  |  |
| JR-1c | 200.55 | 4088 | 717. |  |  |

Appendix 8 - Results of concentration calculations for the standards analyzed for both the standard INAA procedure and the new ENAA procedure for barium. NIST SRM 278 (obsidian) was used as a primary standard.
A. Standard INAA procedure $(\mathrm{Ba}-138(\mathrm{n}, \gamma) \mathrm{Ba}-139)$

| Sample ID | $\begin{aligned} & \text { Mass } \\ & (\mathrm{mg}) \end{aligned}$ | $\begin{gathered} \text { Counts } \\ (165 \mathrm{keV}) \\ \hline \end{gathered}$ | Calculated Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| AGV-1a | 200.61 | 14166 | 1060 | 1020 | 38 |
| AGV-1b | 200.15 | 13544 | 1010 |  |  |
| AGV-1c | 200.75 | 13186 | 981 |  |  |
| JA-1a | 200.03 | 2804 | 249 | 226 | 31 |
| JA-1b | 200.37 | 2165 | 191 |  |  |
| JA-1c | 200.70 | 2703 | 238 |  |  |
| JB-2a | 200.47 | 1870 | 157 | 129 | 31 |
| JB-2b | 200.37 | 1603 | 135 |  |  |
| JB-2c | 200.64 | 1136 | 95.2 |  |  |
| JR-1a | 200.80 | -- | -- | -- | -- |
| JR-1b | 200.66 | -- | -- |  |  |
| JR-1c | 200.40 | -- | -- |  |  |
| FA 381 | 100.29 | 11015 | 1240 | 1360 | 136 |
| FA 382 | 99.99 | 11416 | 1290 |  |  |
| FA 383 | 99.43 | 10446 | 1190 |  |  |
| FA 393 | 99.77 | 13221 | 1511 |  |  |
| FA 394 | 99.73 | 13018 | 1490 |  |  |
| FA 395 | 99.98 | 12520 | 1430 |  |  |

B. Epithermal INAA procedure ( $\mathrm{Ba}-130(\mathrm{n}, \gamma) \mathrm{Ba}-131)$

| Sample ID | $\begin{aligned} & \hline \text { Mass } \\ & \text { (mg) } \end{aligned}$ | $\begin{gathered} \text { Counts } \\ (496 \mathrm{keV}) \end{gathered}$ | Calculated Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| FA7257 | 99.56 | 18441 | 1290 | 1280 | 19 |
| FA7258 | 99.21 | 16968 | 1270 |  |  |
| AGV-1a | 200.26 | 34511 | 1190 | 1190 | 8 |
| AGV-1b | 201.62 | 33647 | 1190 |  |  |
| AGV-1c | 201.58 | 33175 | 1200 |  |  |
| JA-1a | 200.51 | 8561 | 291 | 294 | 2 |
| JA-1b | 200.91 | 8473 | 296 |  |  |
| JA-1c | 200.58 | 8154 | 294 |  |  |
| JB-2a | 199.79 | 6151 | 212 | 216 | 6 |
| JB-2b | 199.90 | 6289 | 223 |  |  |
| JB-2c | 200.45 | 5890 | 214 |  |  |
| JR-1a | 200.22 | 1552 | 54.2 | 53.8 | 0.4 |
| JR-1b | 200.80 | 1502 | 53.7 |  |  |
| JR-1c | 200.55 | 1453 | 53.5 |  |  |

Appendix 9 - Results of concentration calculations for the standards analyzed for the new ENAA procedure for arsenic after both a 7-day and 5-day decay. NIST SRM 278 (obsidian) was used as a primary standard.
A. ENAA procedure, 7-day decay

| Sample ID | Mass <br> (mg) | Counts <br> (496 keV) | Calculated <br> Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| FA7257 | 99.56 | 61412 | 152 | 150 | 2 |
| FA7258 | 99.21 | 29754 | 149 |  |  |
| AGV-1a | 200.26 | 873 | 1.09 | 1.22 | 0.28 |
| AGV-1b | 201.62 | 609 | 1.03 |  |  |
| AGV-1c | 201.58 | 664 | 1.54 |  | 0.17 |
| JA-1a | 200.51 | 2501 | 3.11 | 3.09 |  |
| JA-1b | 200.91 | 1713 | 2.91 |  | 0.37 |
| JA-1c | 200.58 | 1397 | 3.26 |  |  |
| JB-2a | 199.79 | 2648 | 3.38 |  | 0.4 |
| JB-2b | 199.90 | 1985 | 3.47 |  |  |
| JB-2c | 200.45 | 1701 | 4.06 |  |  |
| JR-1a | 200.22 | 11949 | 15.7 | 15.9 |  |
| JR-1b | 200.80 | 8891 | 15.9 |  |  |
| JR-1c | 200.55 | 6677 | 16.3 |  |  |

B. ENAA procedure, 5-day decay

| Sample ID | Mass <br> (mg) | Counts <br> (496 keV) | Calculated <br> Concentration | Average | Standard <br> Deviation |
| :---: | :---: | :---: | :---: | :---: | :---: |
| FA 1633 A | 200.75 | 294126 | 141 |  | 142 |
| FA 1633 B | 201.50 | 218575 | 141 | 2 |  |
| FA 1633 C | 201.5 | 161765 | 143 |  |  |
| JR-1d | 199.36 | 34167 | 16.3 | 16.4 | 0.2 |
| JR-1e | 199.72 | 25461 | 16.3 |  |  |
| JR-1f | 200.30 | 19044 | 16.7 |  |  |

## VitA

Magen Elizabeth Coleman was born to Charles and Patricia Coleman on October 25, 1984, in Bergenfield, New Jersey. Interested in ancient history, science, and art from an early age, she attended the Academy of the Holy Angels in Demarest, NJ for high school, where she was able to pursue those interests, especially in taking AP courses in Latin and Chemistry. After swearing off chemistry, she attended the University of Mary Washington, in Fredericksbug, VA, with the goal of majoring in Classics/Latin, and was re-inspired in science after taking analytical chemistry. Under the guidance of her advisors, Dr. Liane Houghtalin of the Classics Department and Dr. Raymond Scott of the Chemistry Department, she graduated in May 2006 with a double major in Chemistry and Classics. Not quite willing to give up either of her two majors, she discovered the field of archaeometry through Dr. Michael Glascock of the University of Missouri Research Reactor. In August 2006, she began the journey toward a Ph.D. in chemistry with Dr. J. David Robertson as her advisor. After receiving her doctorate in May 2010, she will be headed to Los Alamos National Laboratory, in Los Alamos, NM for a postdoctoral position under Dr. Lav Tandon.

